

# Voids, superclusters and the ISW effect

Shaun Hotchkiss

(with Seshadri Nadathur)

(also Subir Sarkar, Samuel Flender and Jubilee)



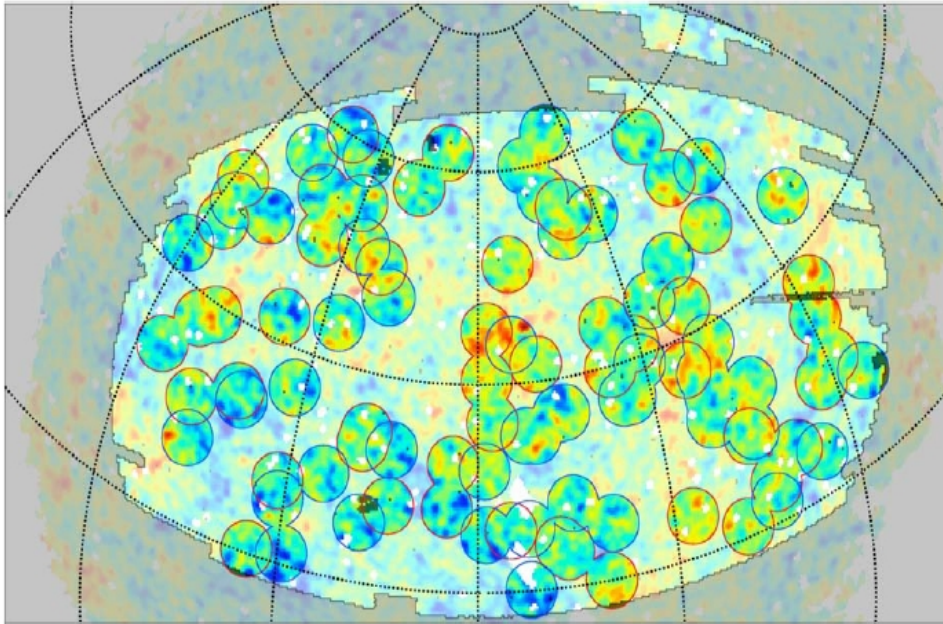
University of Sussex

What I want you to take from this talk

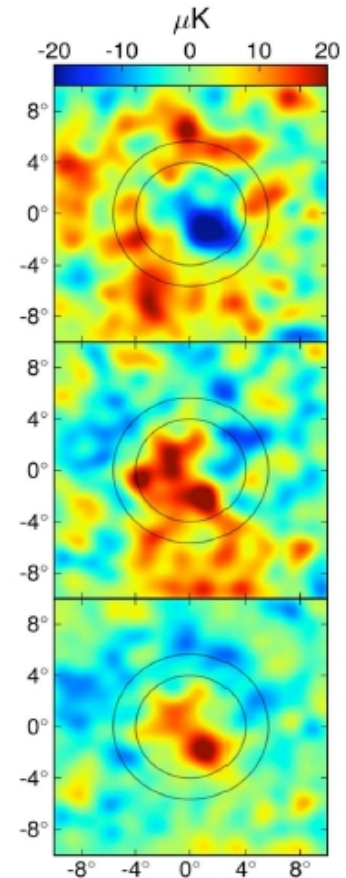
What I want you to take from this talk

Knowledge that the ISW effect from voids in  
LCDM is extremely small (unobservable)

$A > 4\sigma$  anomaly, it would seem...



Granett, Neyrinck, Szapudi 2008

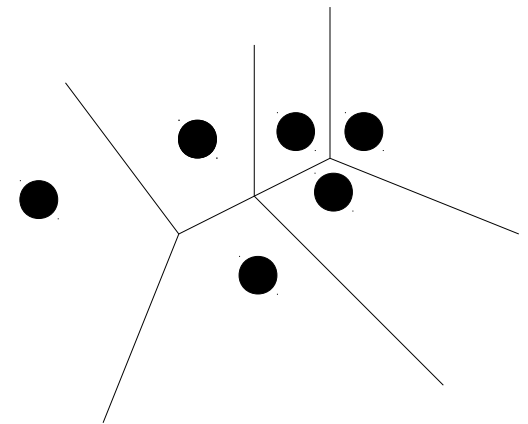
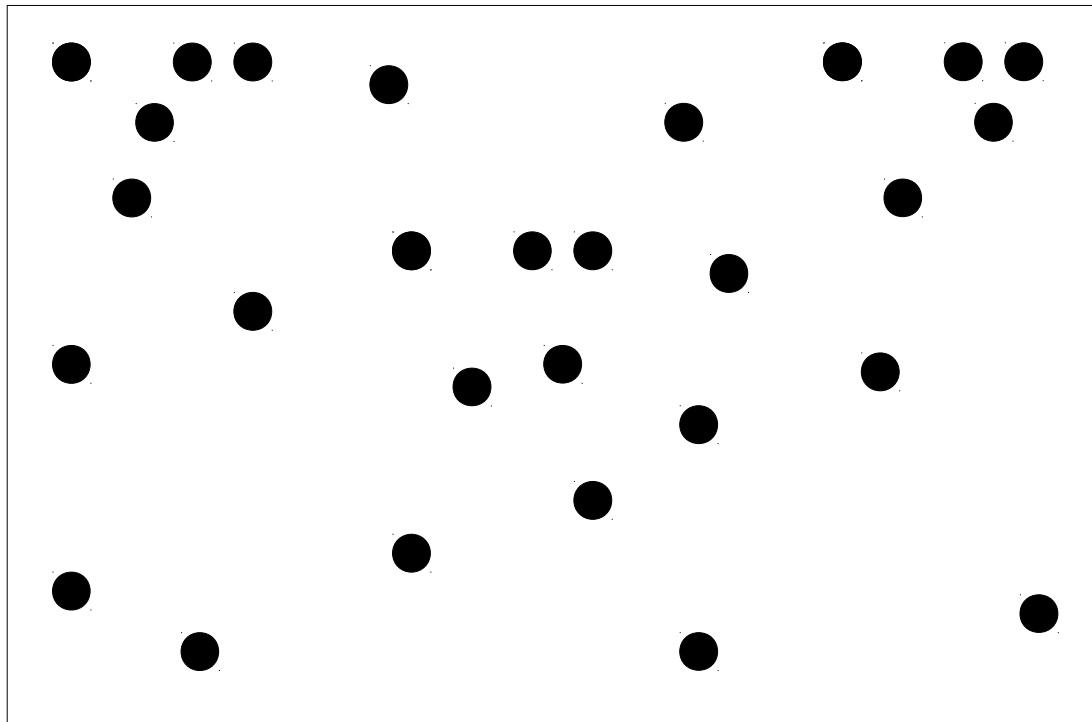


Why are voids  
interesting?

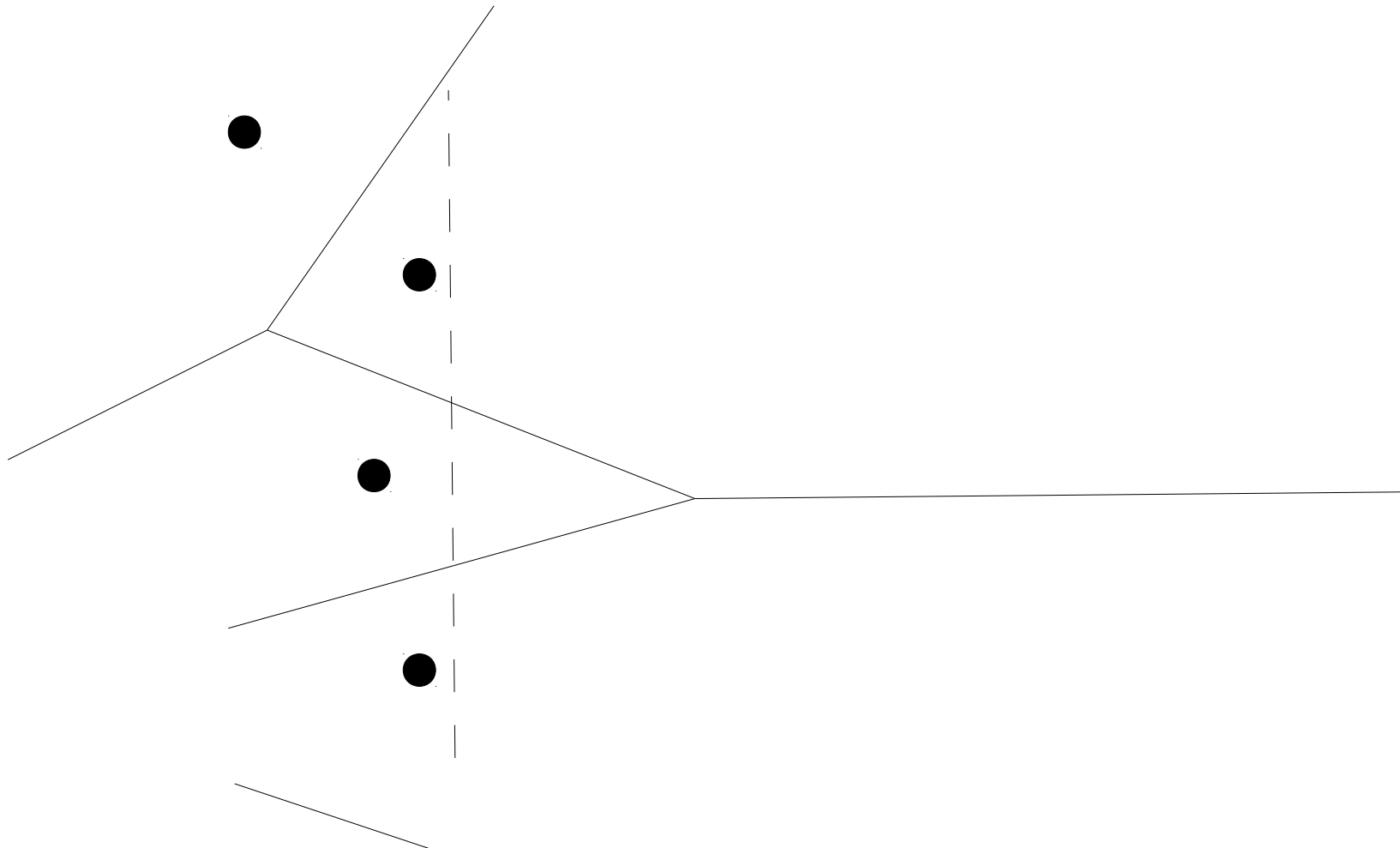
# How to find these things?

ZOBOV (Neyrinck 2008)

# Voronoi tessellation

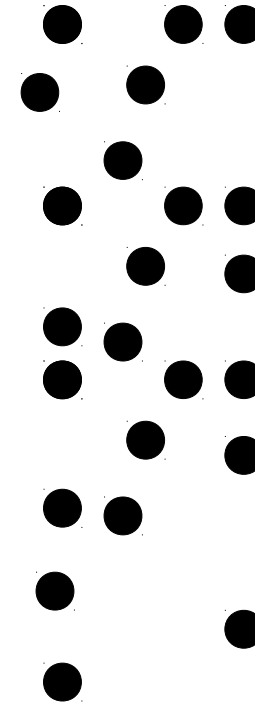
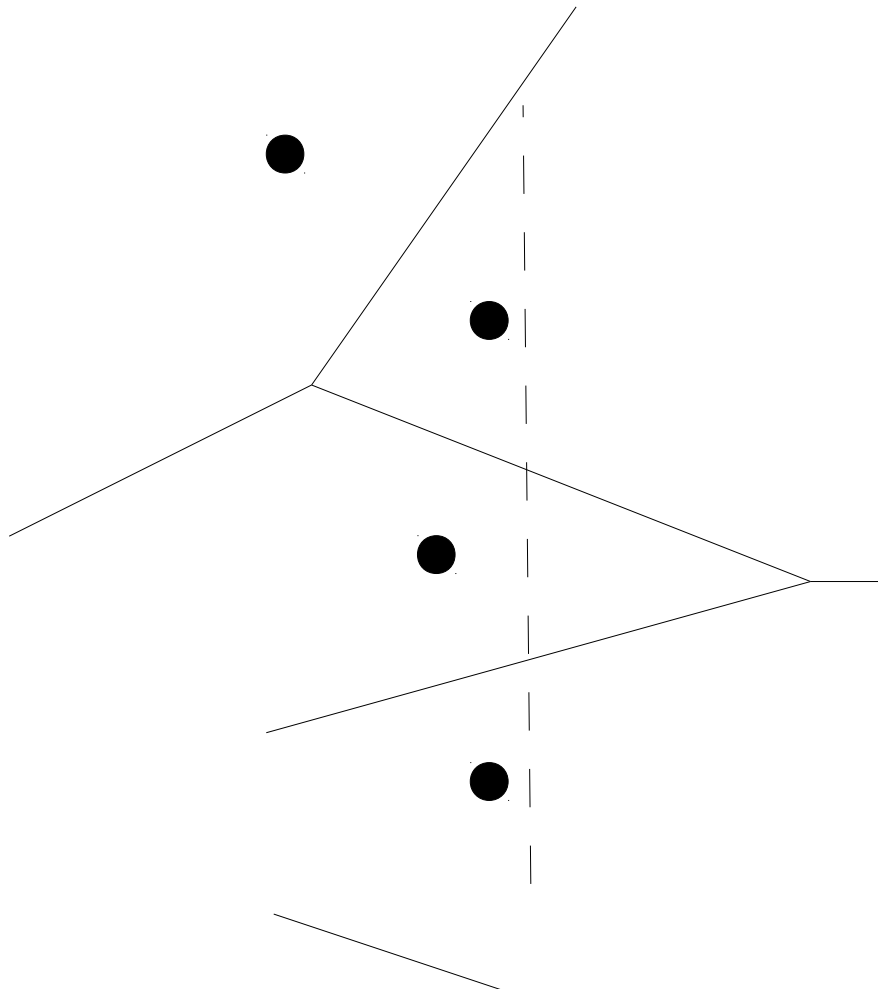


# Voronoi tessellation at survey edge

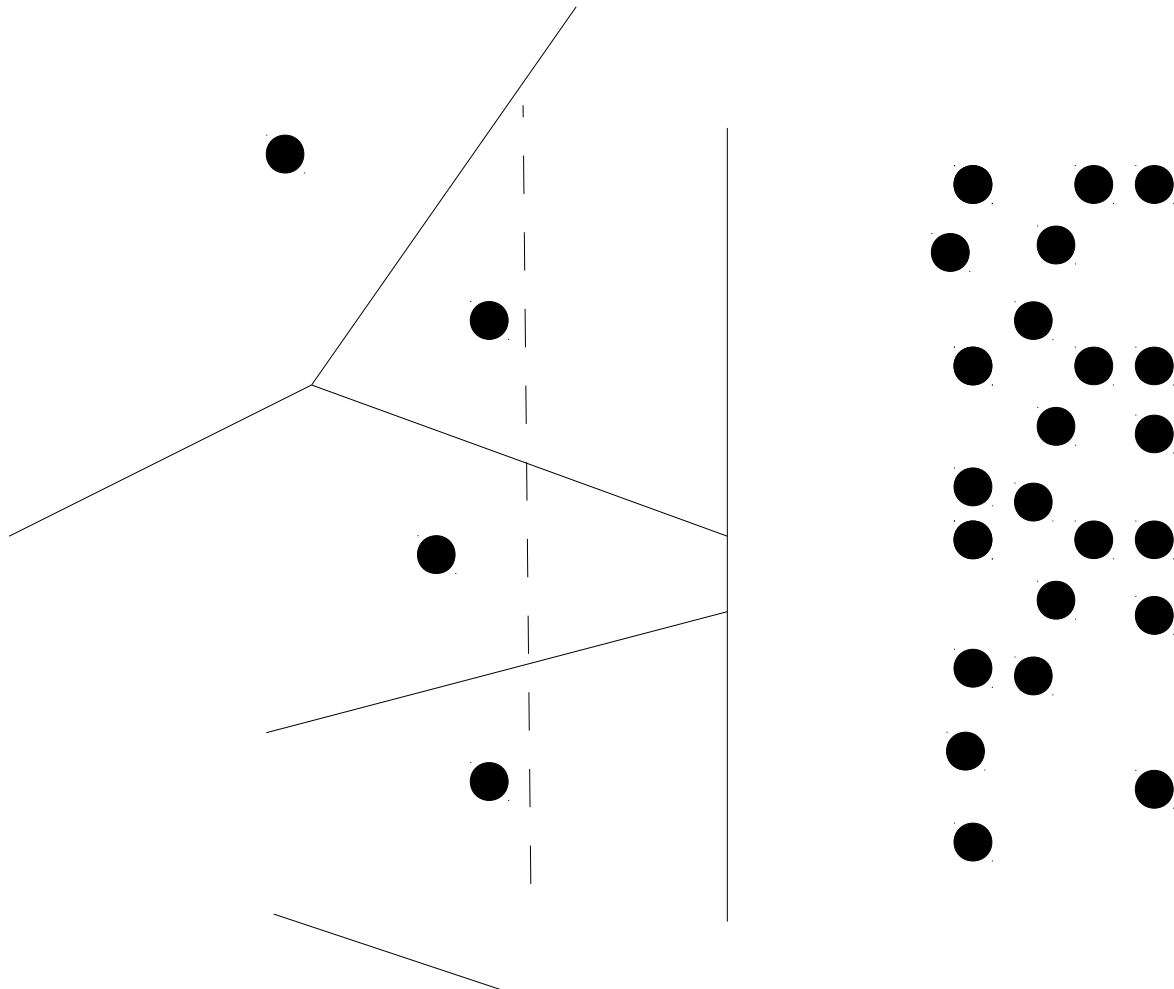




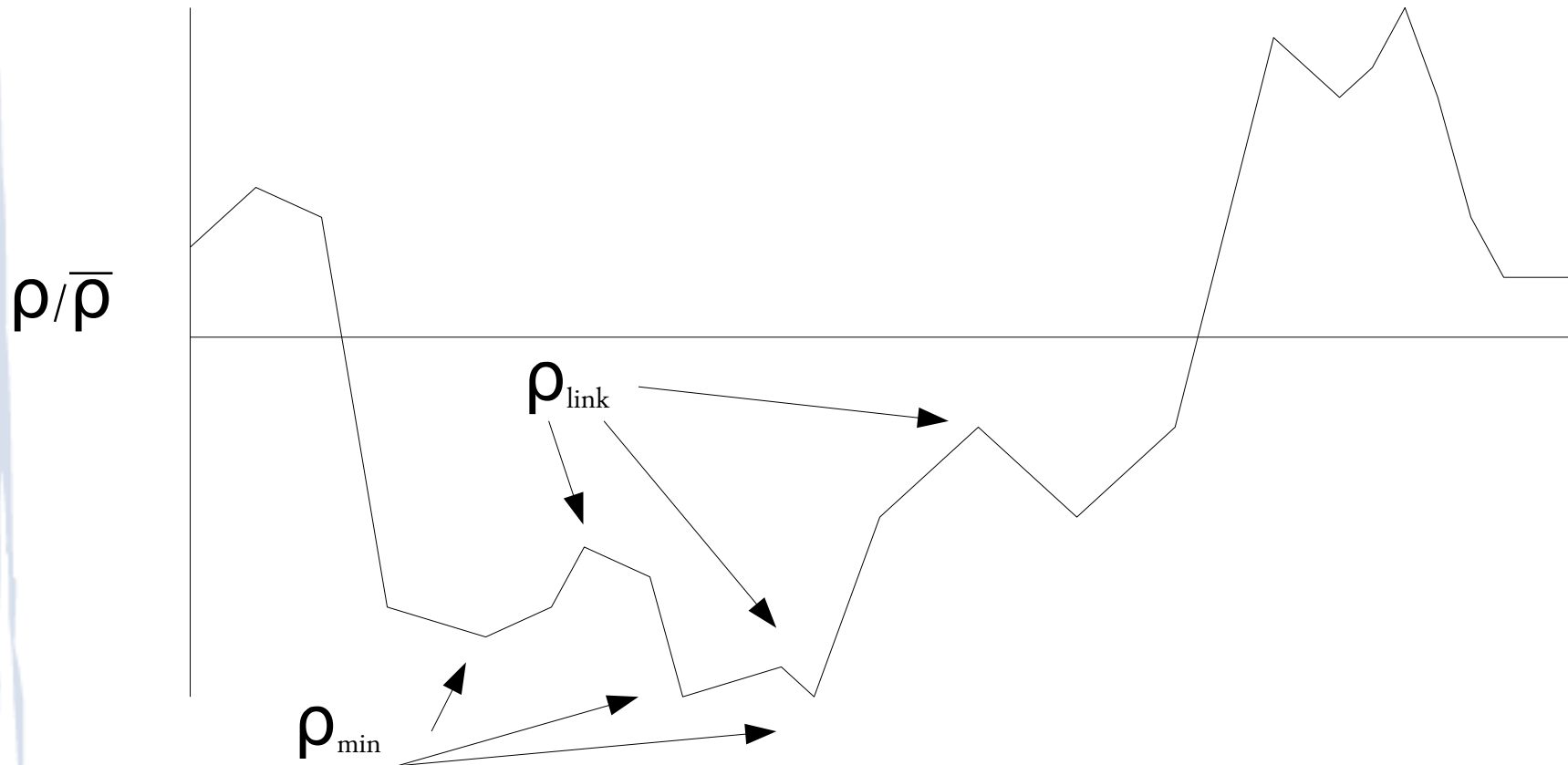
# Voronoi tessellation at survey edge



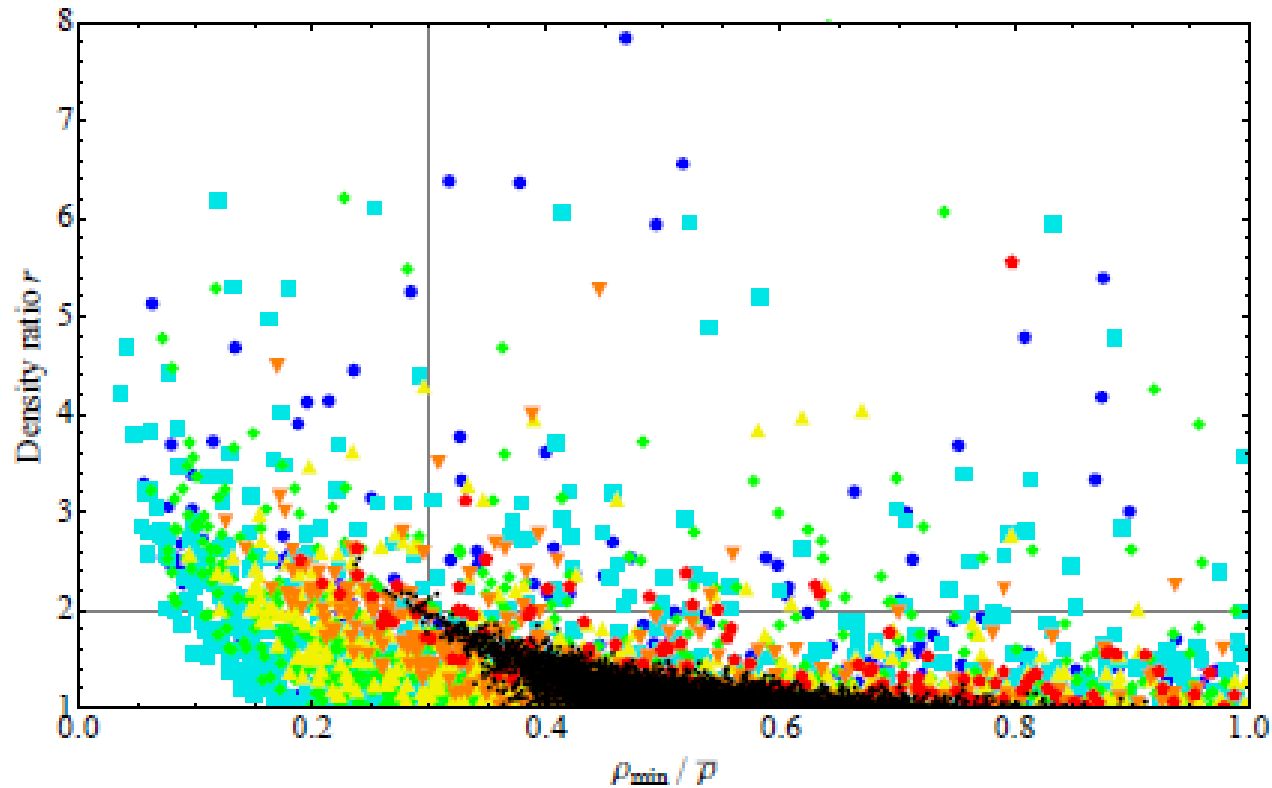
# Voronoi tessellation at survey edge



# ZOBOV and the density field



# Distinguishing from Poisson



Type 1:  $\rho_{\min} < 0.3$

and also Type 2:

$\rho_{\min}$  and  $\rho_{\text{link}} < 0.2$

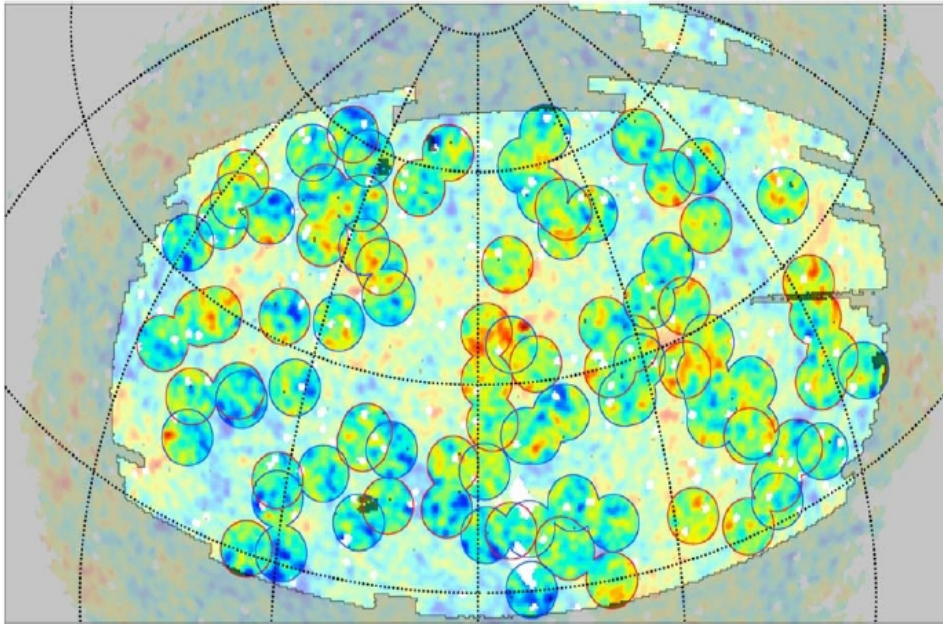
Where is the void and how big is it?

$$\mathbf{X}_v = \frac{1}{\sum_i V_i} \sum_i \mathbf{x}_i V_i$$

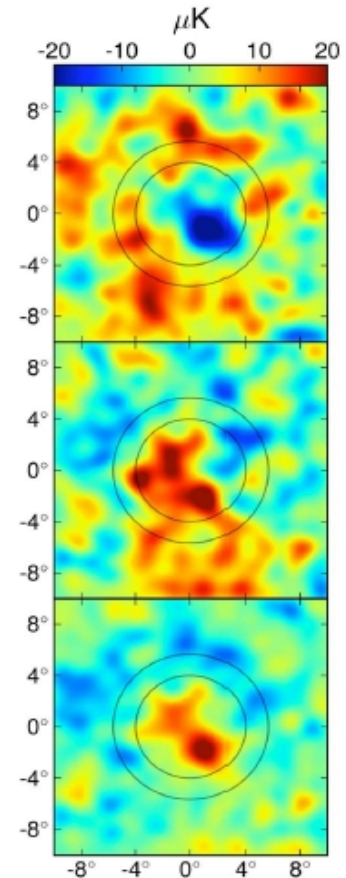
$$R_{\text{eff}} = \left( \frac{3}{4\pi} V \right)^{1/3}$$

# The “ISW” effect of superstructures

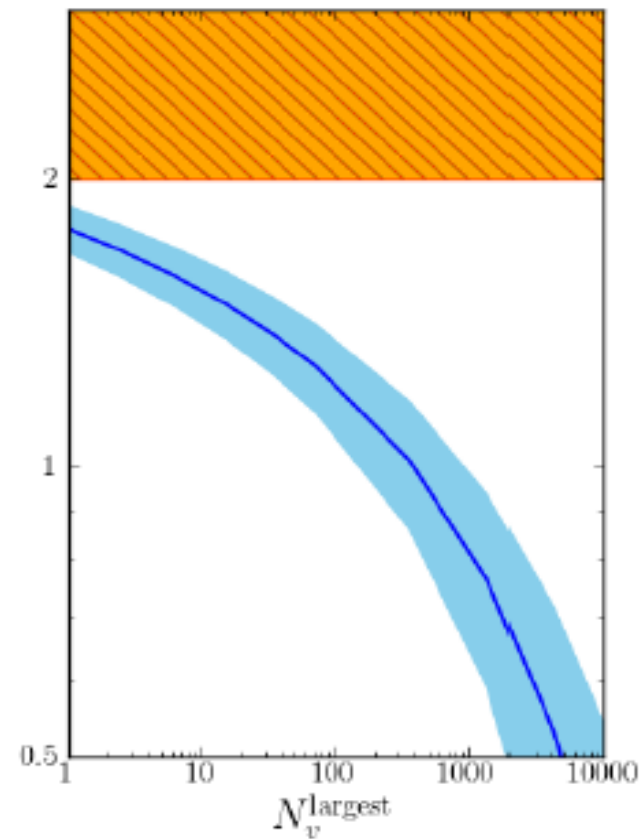
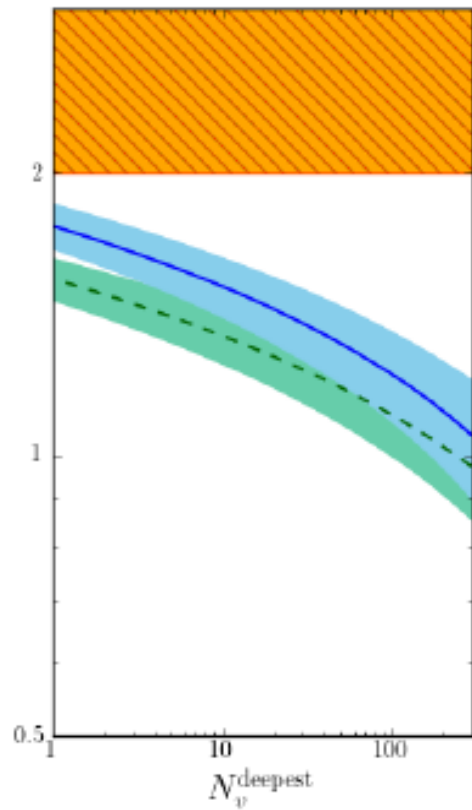
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Granett, Neyrinck, Szapudi 2008

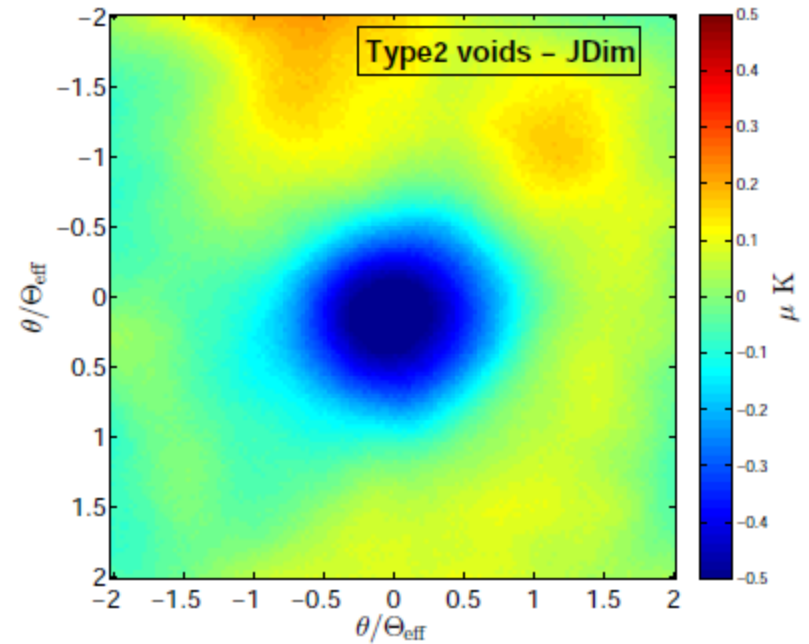
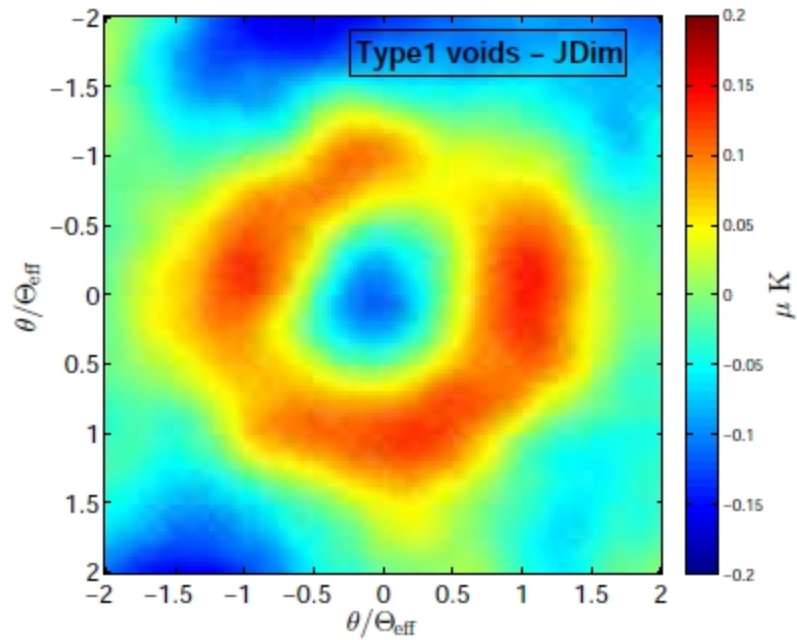


# Theory (us)

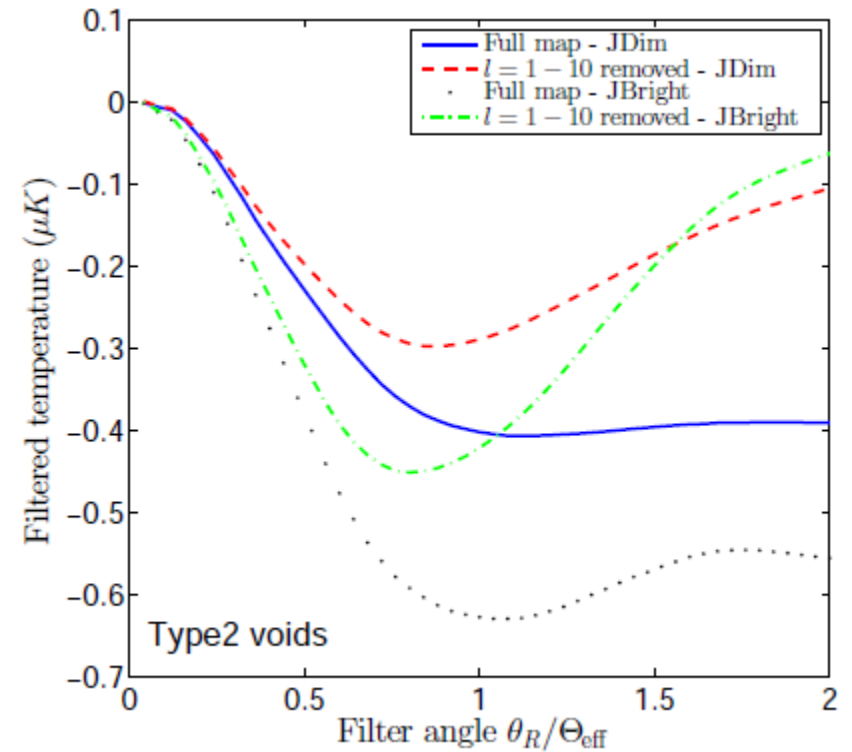
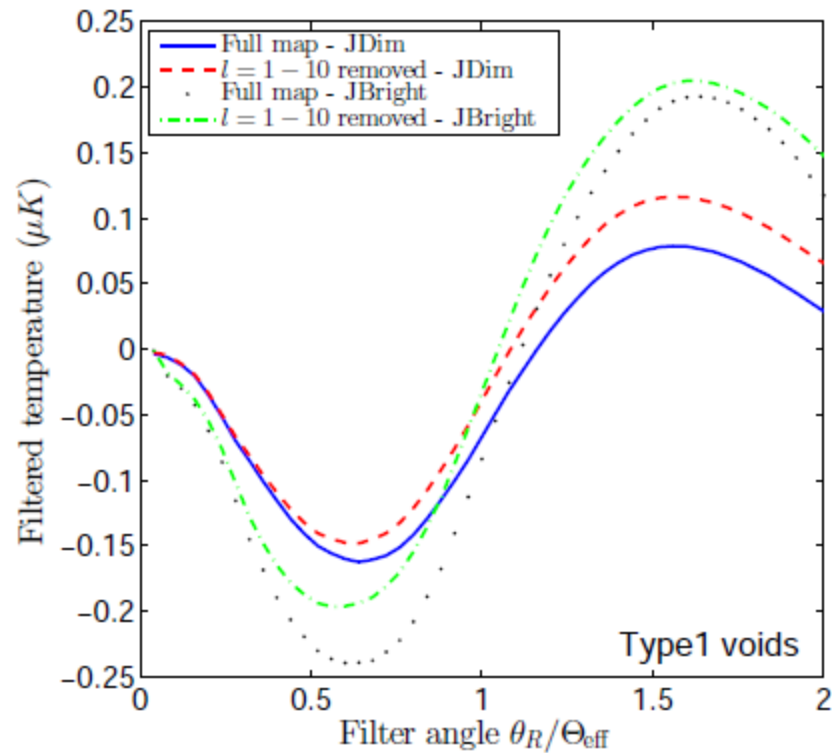




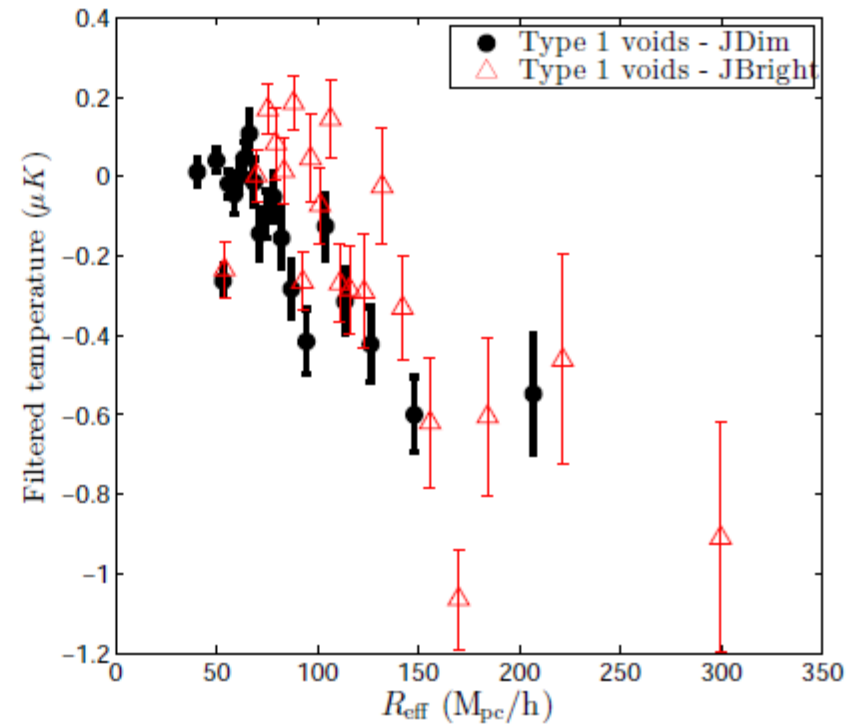
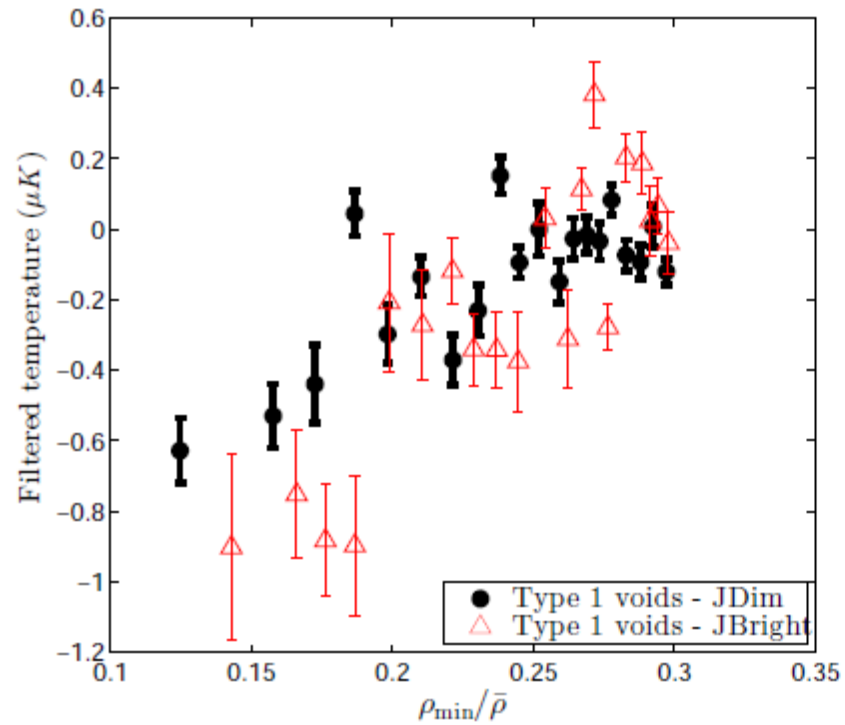
# Jubilee et al. (us)



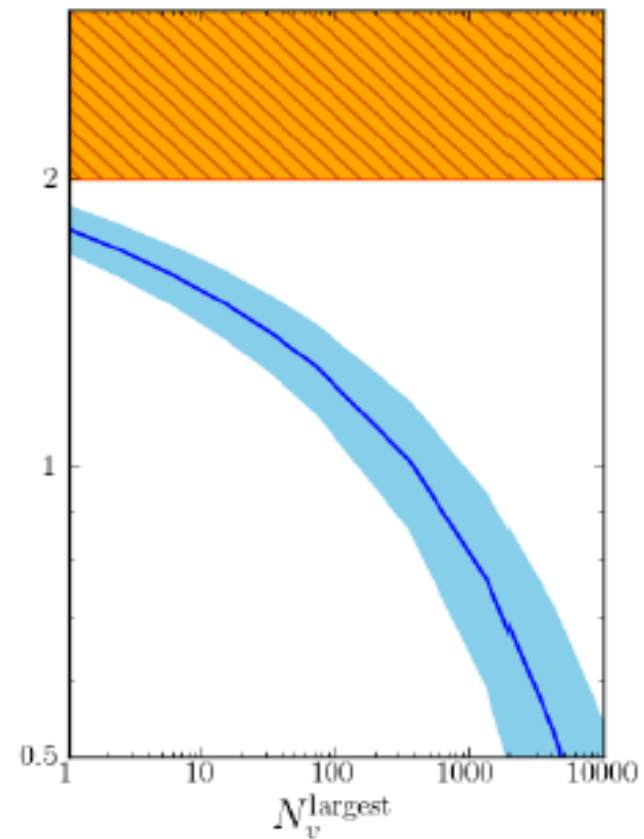
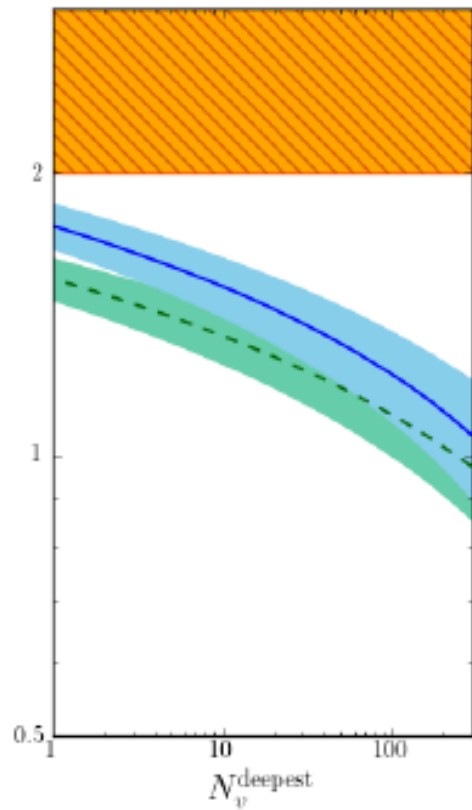
# Jubilee et al. (expectations)



# Jubilee et al. (most extreme)



# Theory (most extreme)



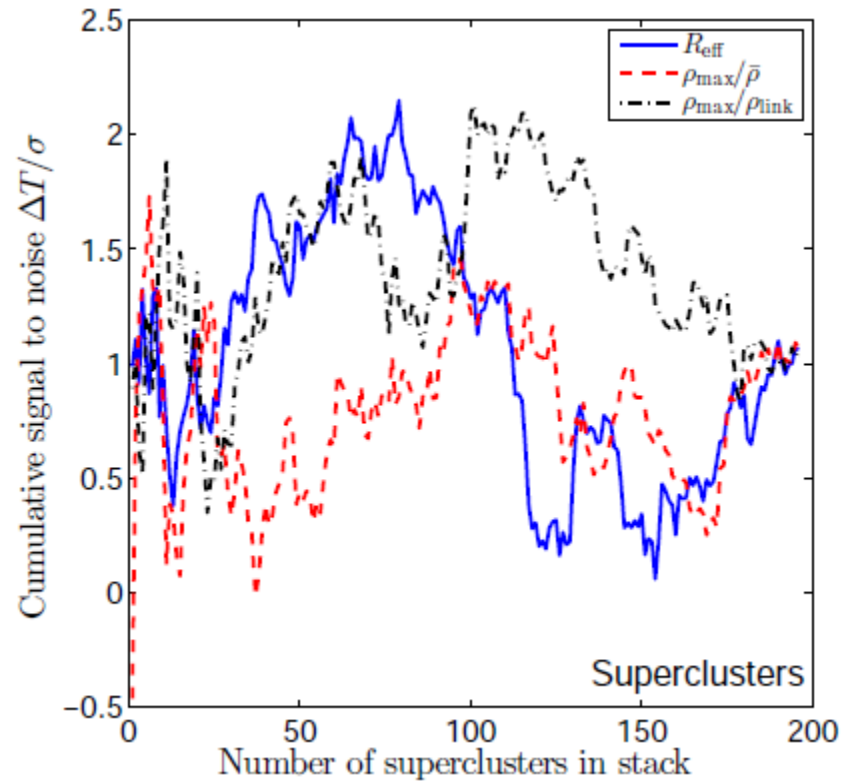
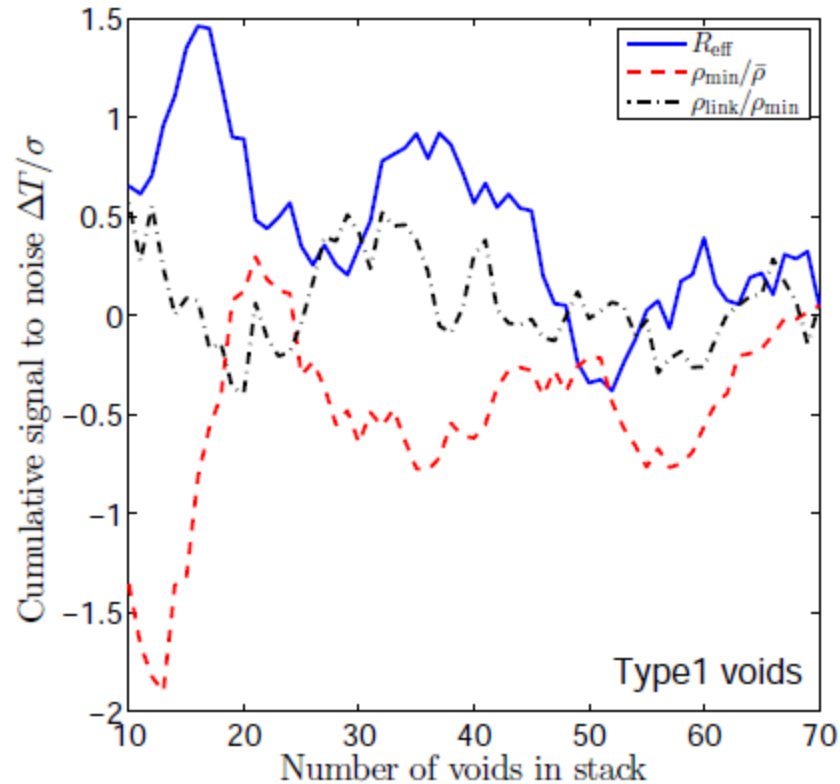
What about the real  
world?

Nothing!

# Nothing!

(as expected)

# Really, nothing!





# A Supervoid Imprinting the Cold Spot in the Cosmic Microwave Background

Fabio Finelli, Juan Garcia-Bellido, Andras Kovacs, Francesco Paci, Istvan Szapudi

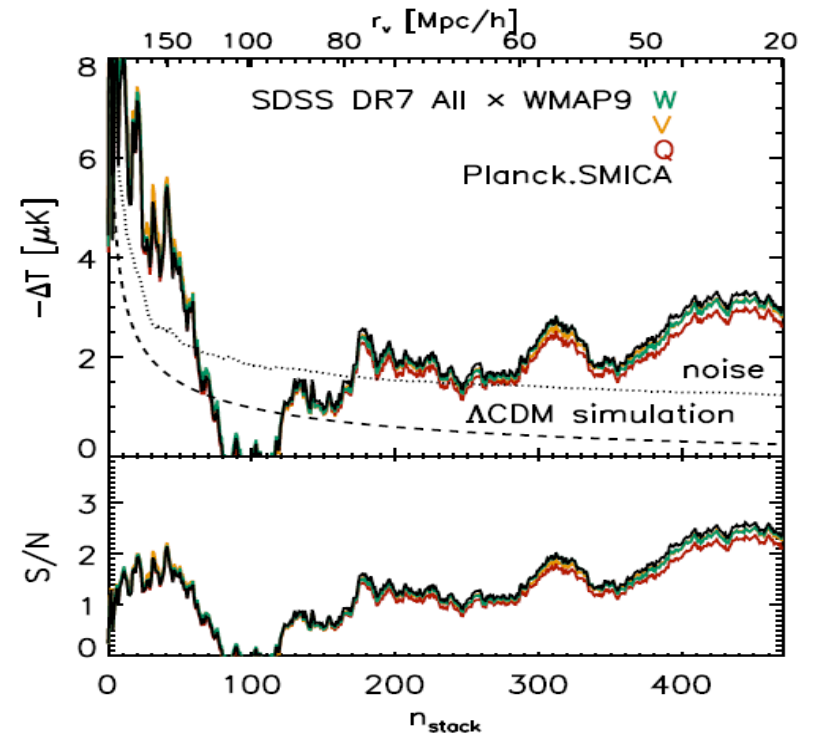
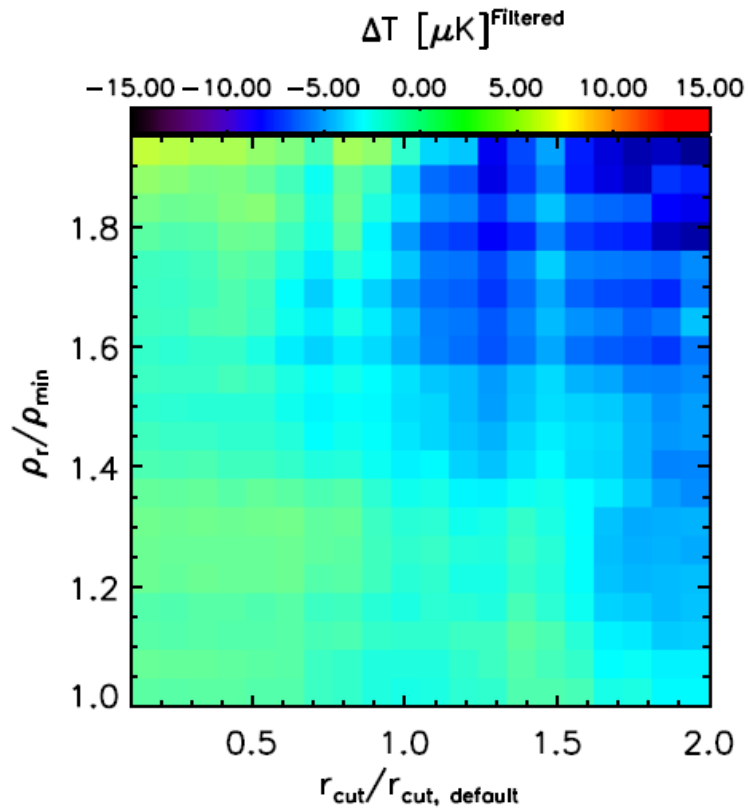
(Submitted on 7 May 2014 (v1), last revised 8 May 2014 (this version, v2))

The Cold Spot is one of the anomalies in the Cosmic Microwave Background, and could be of primordial origin, or caused by a foreground structure. The recently constructed WISE-2MASS all-sky infrared catalogue has a projected underdensity in the direction of the Cold Spot with an angular size 10's of degrees, and as deep as  $\delta \simeq -0.12$  in the center. We show that a spherically symmetric Lemaitre-Tolman-Bondi (LTB) void model can simultaneously fit the underdensity in the WISE-2MASS catalogue and the Cold Spot as observed by both the Wilkinson Anisotropy Probe and Planck satellites. Such an LTB supervoid gives a perfect explanation, via a Rees-Sciama effect, of the Cold Spot anomaly, and is strongly preferred (using a Bayesian analysis) over the null hypothesis (statistical fluctuation) or a texture model. When the galaxy bias, measured from the large-scale angular power spectrum, is taken into account, a simultaneous three-parameter fit for the void model and the temperature profile gives  $z_0 = 0.16 \pm 0.04$  for the mean redshift of the supervoid,  $r_0 = 195 \pm 35 h^{-1}$  Mpc for its size, and  $\bar{\delta} = -0.10 \pm 0.03$  for the top-hat-projected average depth of the void. These parameters are in excellent agreement with the results of \cite{SzapudiEtal2014}, who used additional photometric redshifts from Pan-STARRS1 for direct tomographic imaging of the void.

What I want you to take from this talk

Knowledge that the ISW effect from voids in LCDM is extremely small (unobservable)

# Cai et al (a reproduction?)



# Cai et al (a reproduction?)

