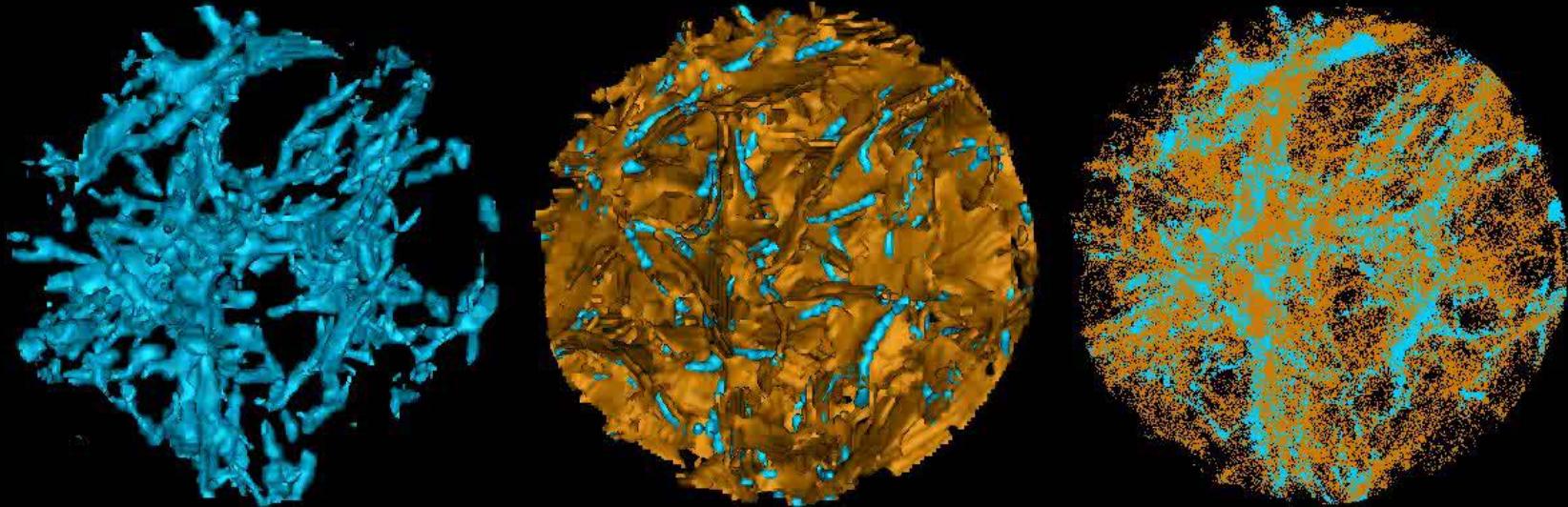


Much Ado: About Nothing: a Review of Voids

Rien van de Weygaert, IAU308-Tallinn, June 2014

The Cosmic Web



MMF/Nexus
Cautun et al. 2013, 2014

Stochastic Spatial Pattern

- Clusters,
- Filaments &
- Walls
- around
- Voids

in which matter & galaxies
have agglomerated
through gravity

Outline

- Why voids are interesting ...
- Void (pre)history
- Void characteristics
- Void formation
- Void dynamics
- Void detection:
 - (M)WVF - Multiscale Watershed Void Finder
- Void hierarchy & Void merger trees
- Voids & Cosmology: void shapes, void outflows & supervoids
- Void galaxies & void substructure

Voids:

Inventory & (Pre)history

REVIEW ARTICLE

Giant voids in the Universe

Ya. B. Zeldovich*, J. Einasto^{†‡} & S. F. Shandarin*

* Institute of Applied Mathematics, Moscow A-47, 125047, USSR

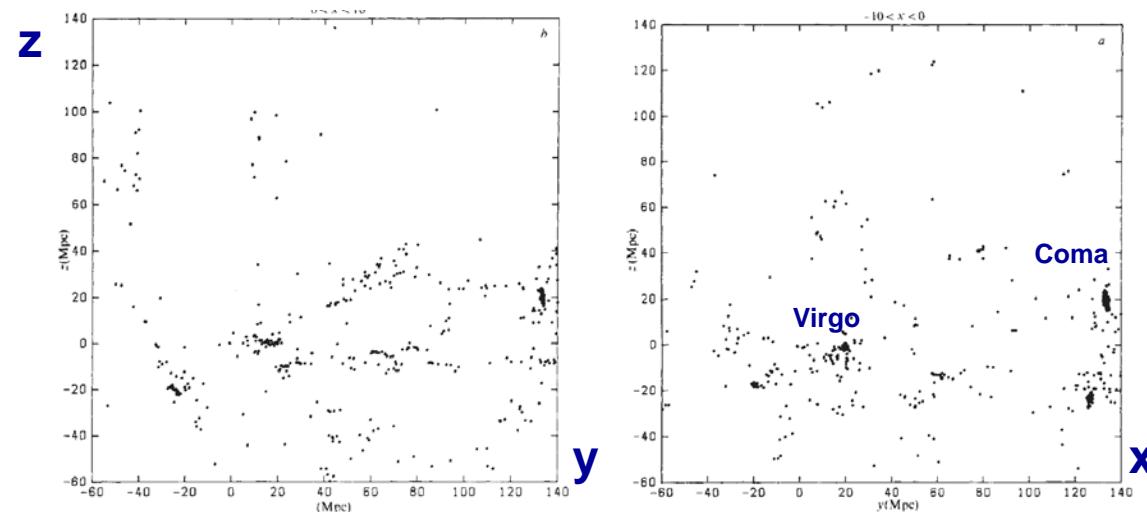
† Tartu Astrophysical Observatory, 202444 Estonia, USSR

‡ European Southern Observatory, 8046 Garching, FRG

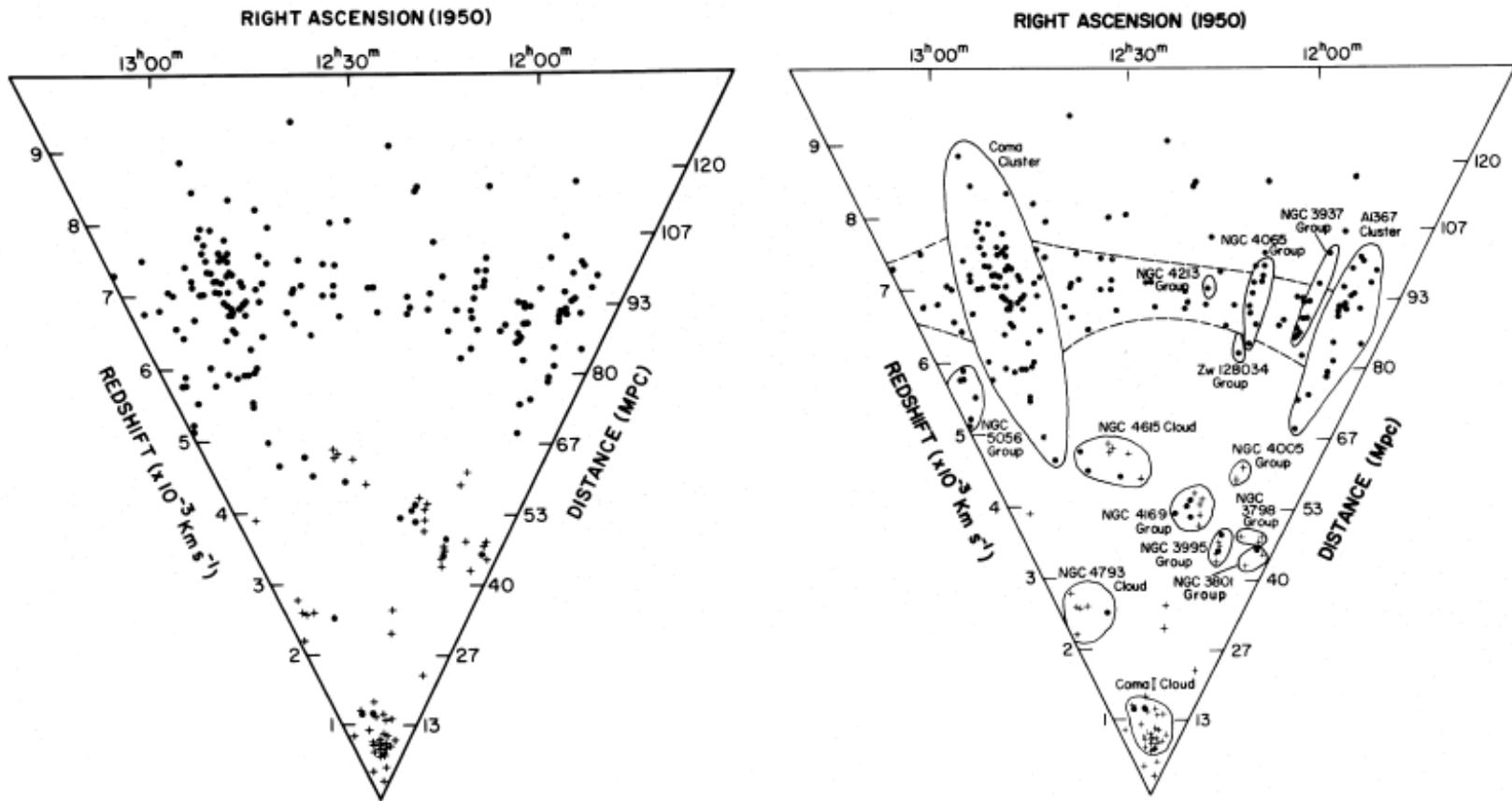
Recent observations indicate that most galaxies are concentrated in superclusters consisting of galaxies, and clusters of galaxies, aligned along strings. Giant volumes exist between superclusters which are almost empty of visible objects. Theories of galaxy formation predict the formation of non-spherical superclusters and giant voids. Large-scale structure changes very slowly, so the currently observed structure reflects the whole history of galaxy formation and structural evolution.

Zeldovich, Einasto & Shandarin 1982:

First linking of observationally visible void regions and the theory of cosmic structure formation.



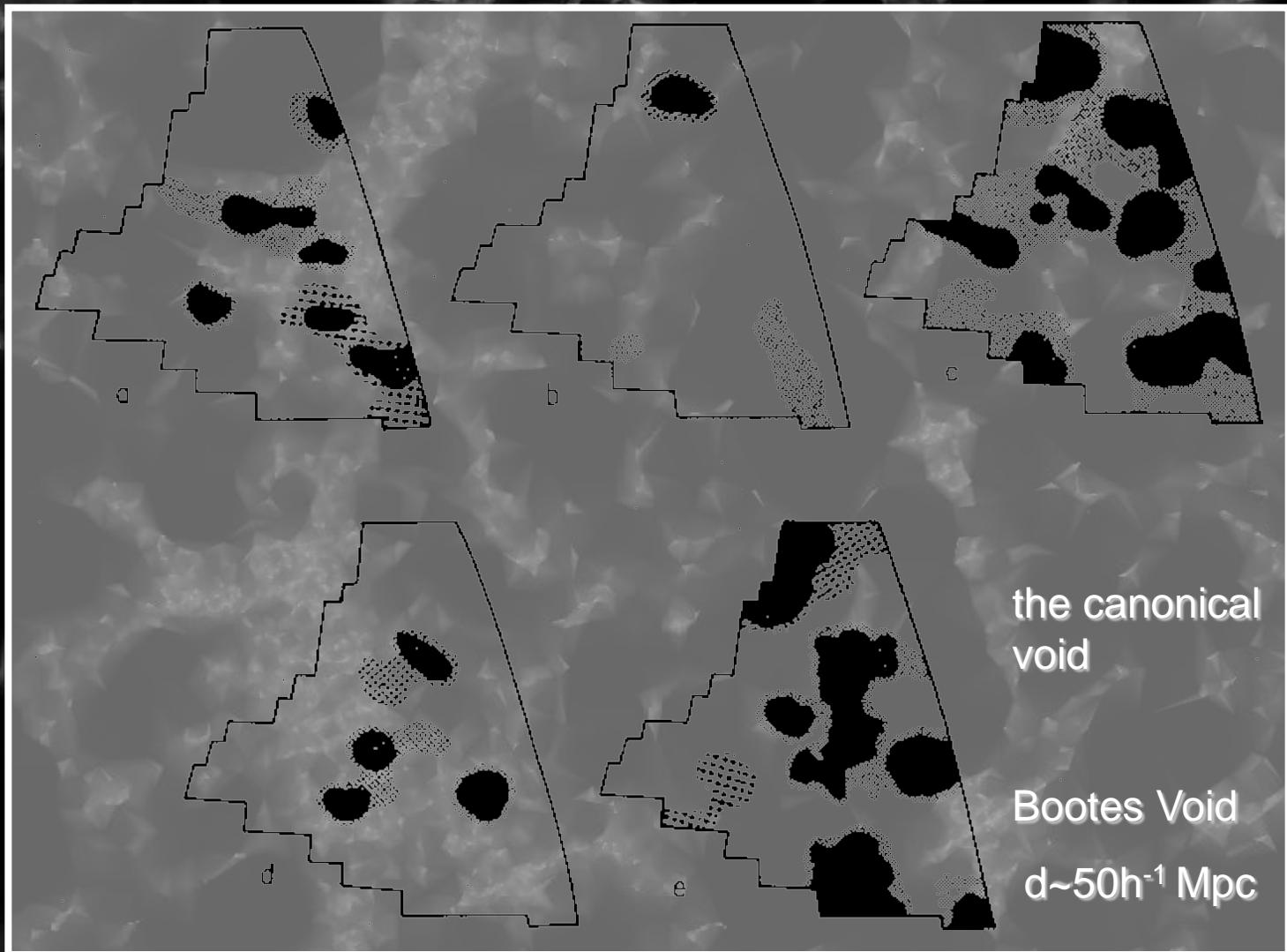
First Voids



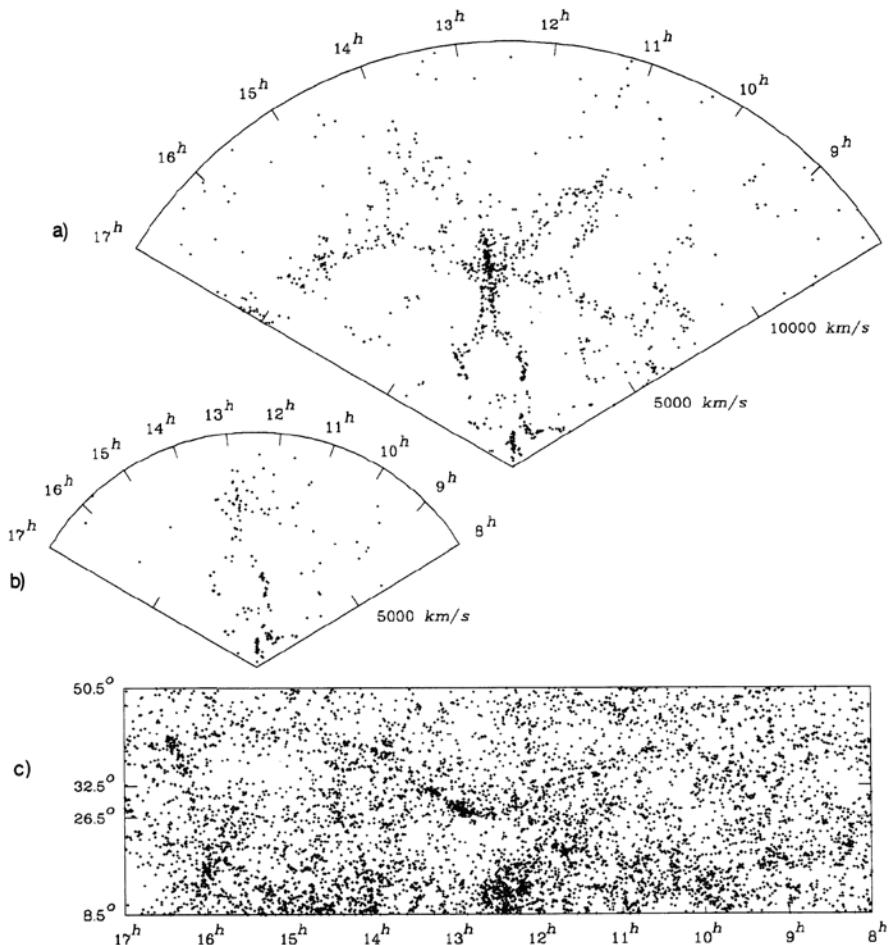
Gregory & Thompson 1978:
redshift survey of Coma/A1367 supercluster region revealed existence of large
near-empty regions of space.

Bootes Void

Kirshner, Oemler, Schectman, Schechter (KOSS) 1981, 1987



Voids & the Cosmic Web

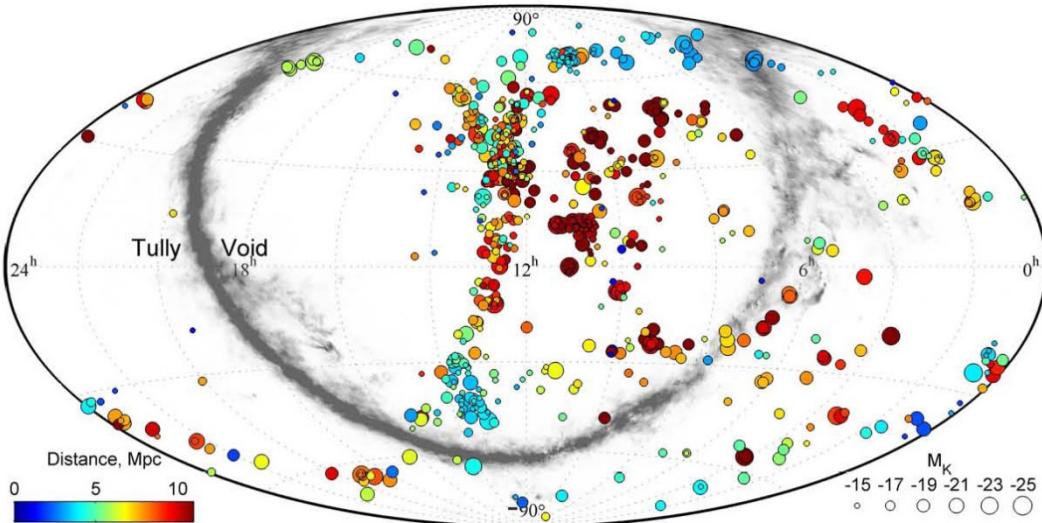
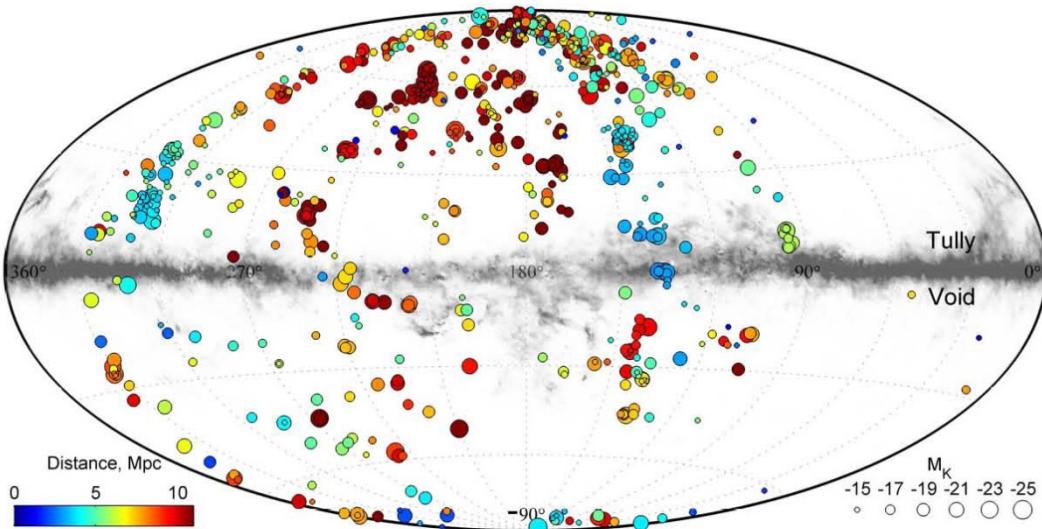


deLapparent, Geller & Huchra, 1986:

“a slice of the Universe”

Voids appear to be an integral part of a complex weblike arrangement of galaxies

local Cosmic Web: Local Void



Karachentsev et al.

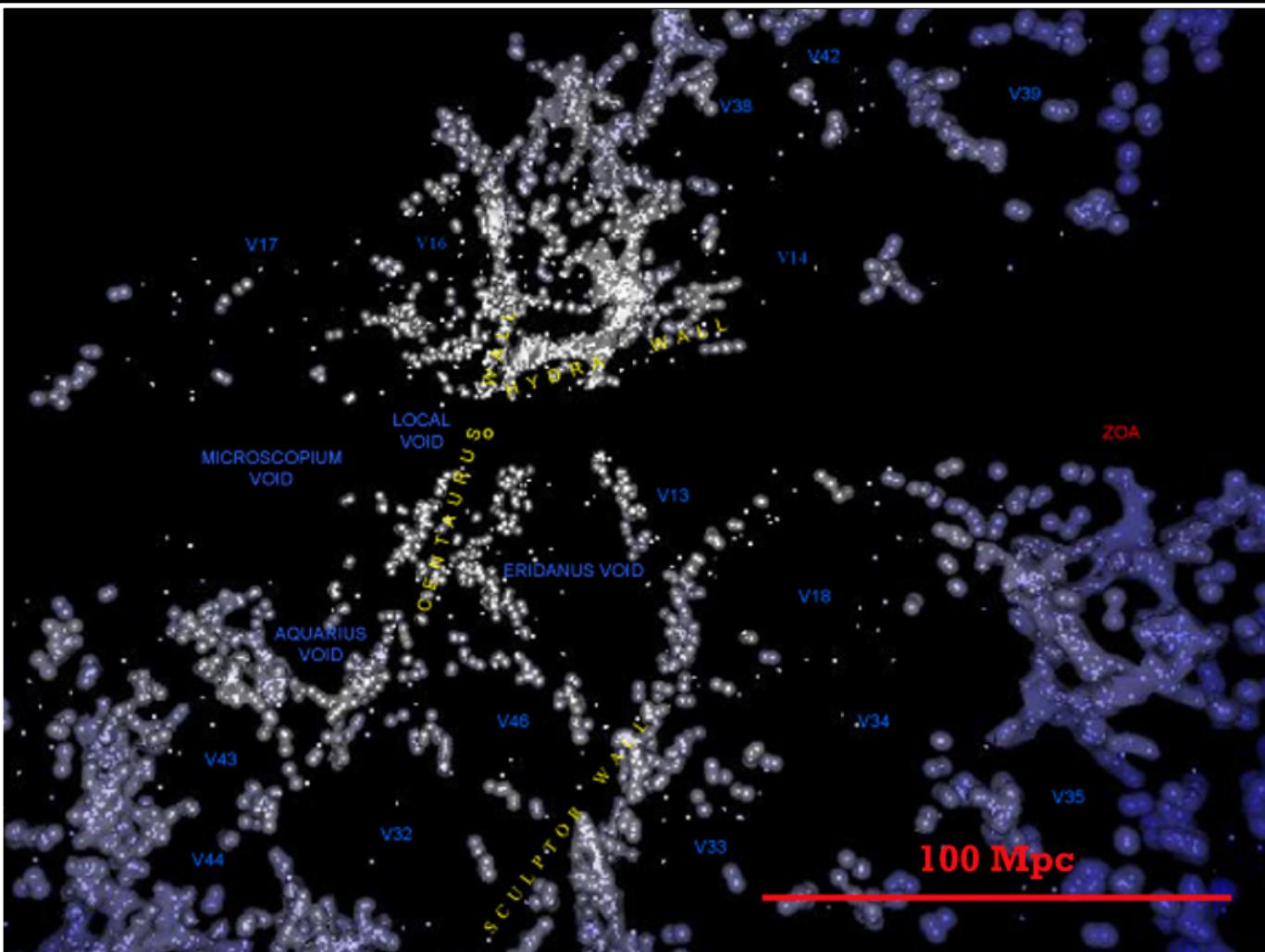
LV catalog:
galaxies within 10 Mpc reveal
beautifully the magnificent

Local Void – Tully Void

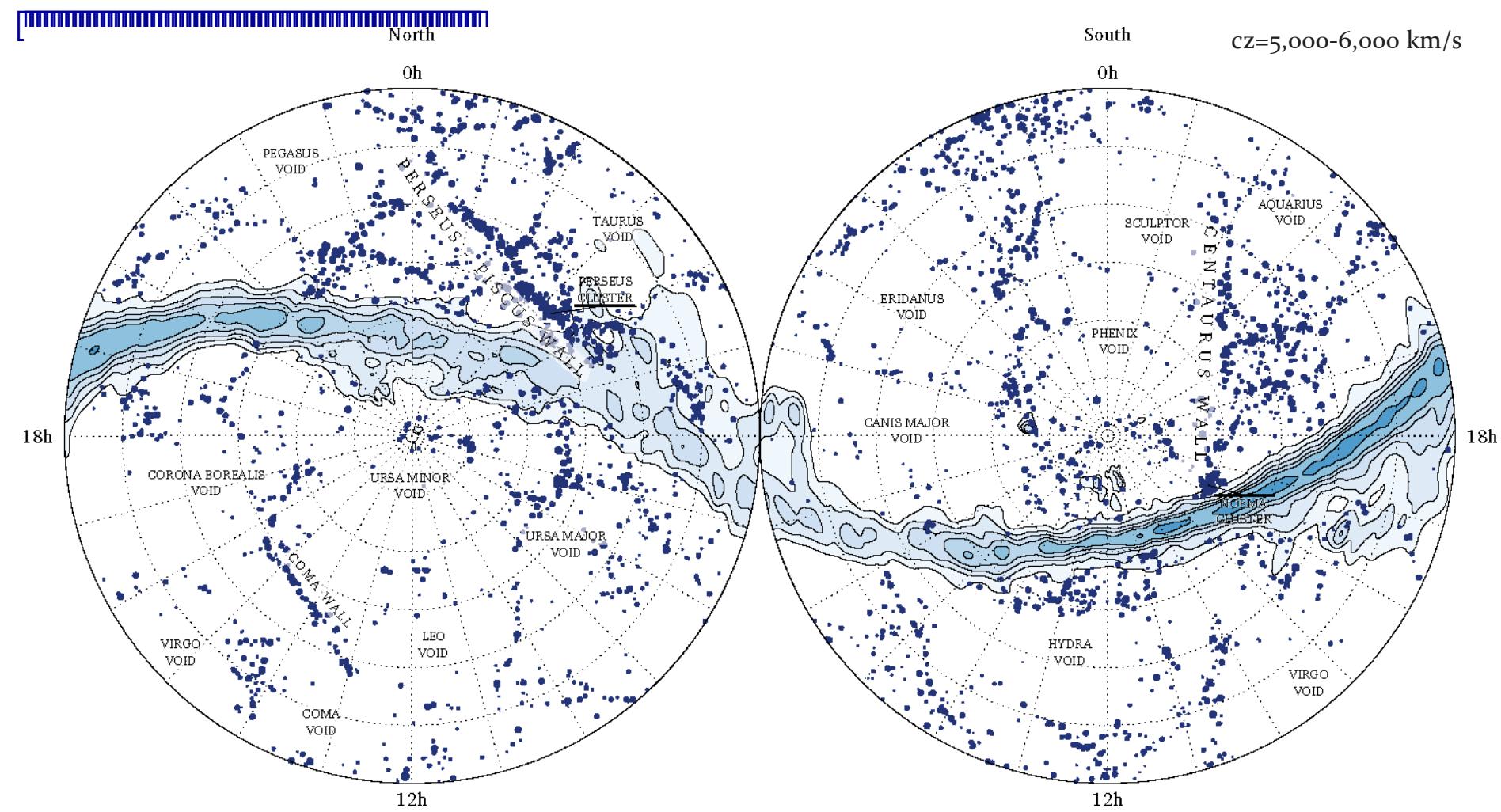
Fairall Void Catalogue

6dF based catalogue

of 526
visually identified
voids.

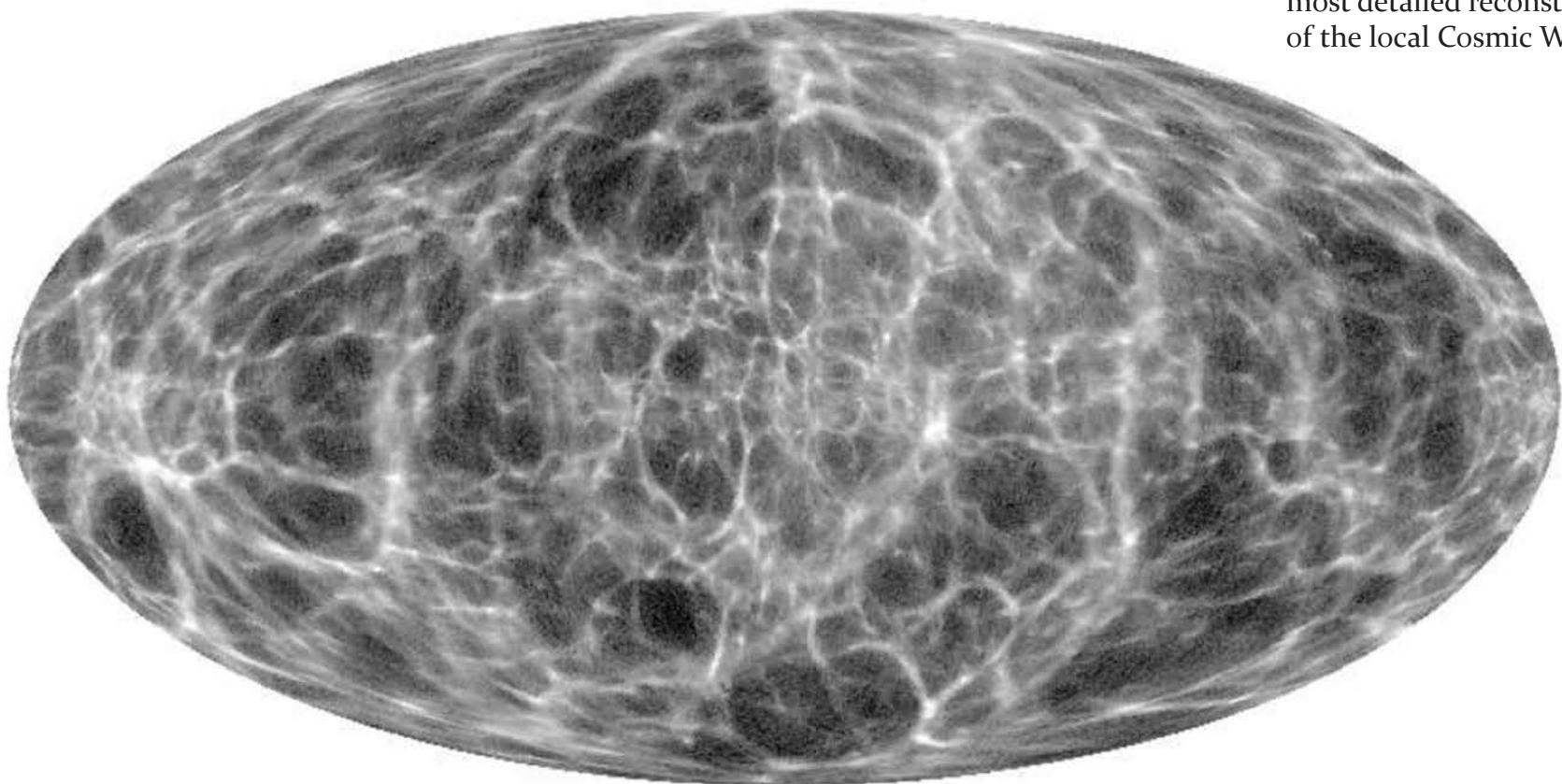


local Cosmic Web: 2MASS



Courtesy: Johan Hidding

local Cosmic Web: 2MASS



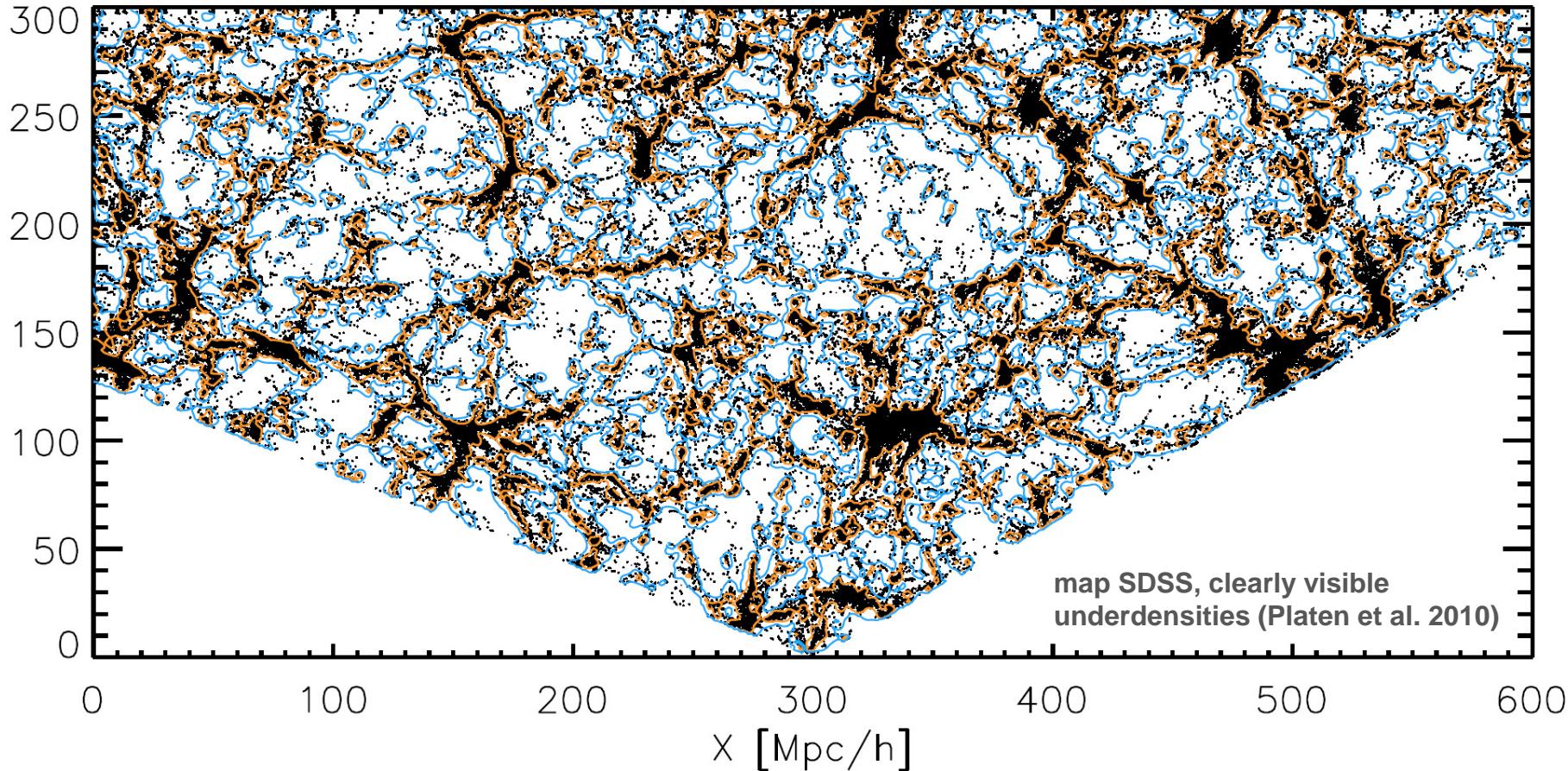
most detailed reconstruction
of the local Cosmic Web

1.0

6.0

Courtesy: Francisco Kitaura

SDSS Galaxy Survey

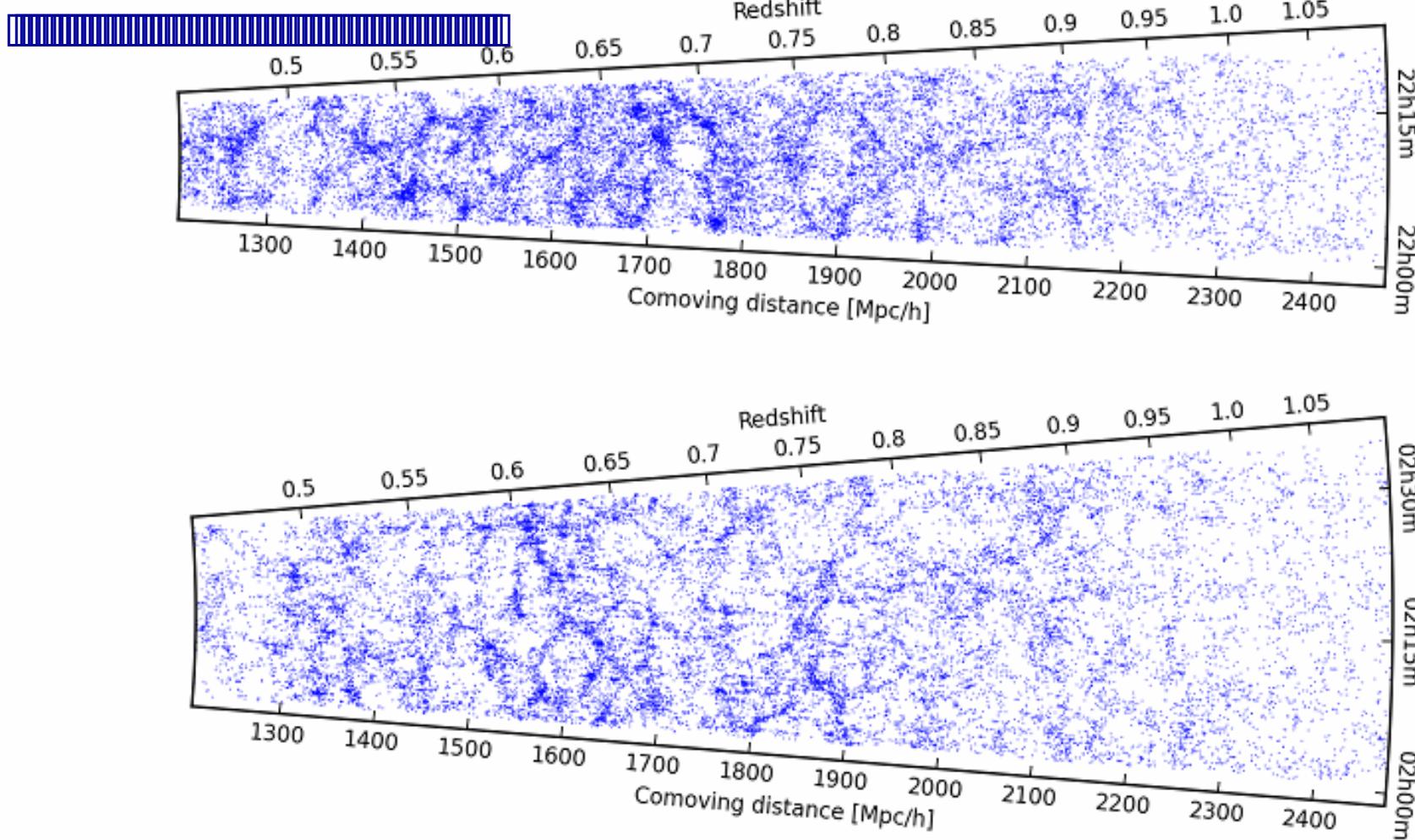


with the advent of large galaxy redshift surveys

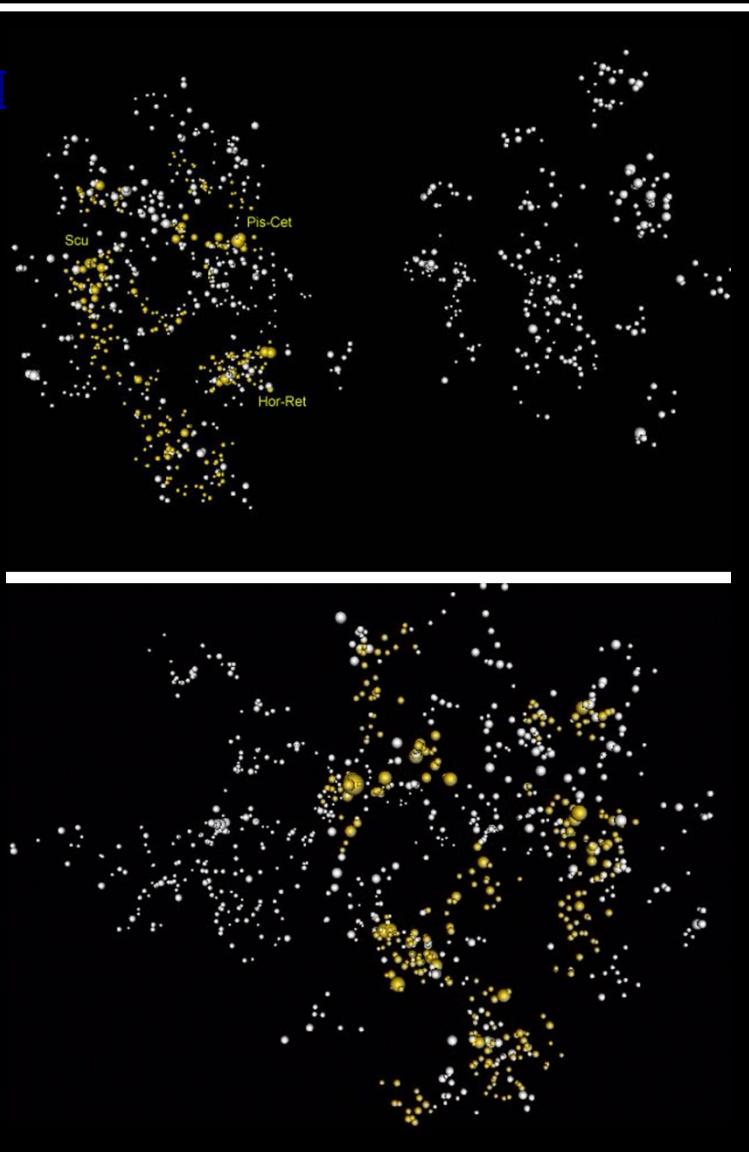
– LCRS, 2dFGRS, SDSS, 2MRS –

voids have been recognized as one of the quintessential components of the Cosmic Web

high-z Cosmic Web: VIPERS



Voids & Clusters

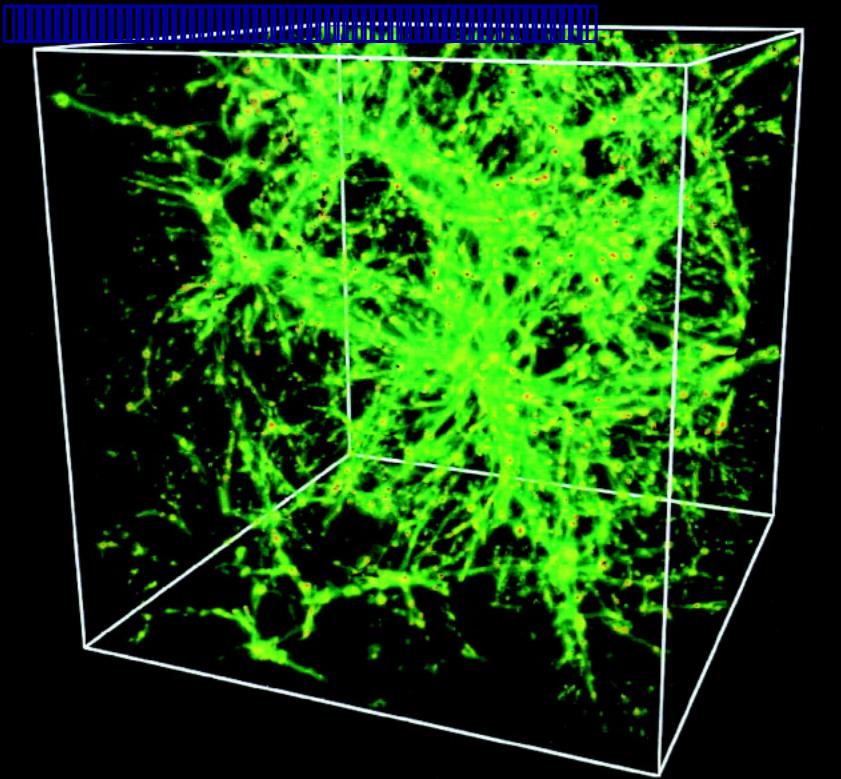


Einasto, Saar et al. (1990s)

- Superclustering in Abell/APM clusters catalog
- Finding of characteristic scale ~140 Mpc, corresponding to large voids in the cluster distribution

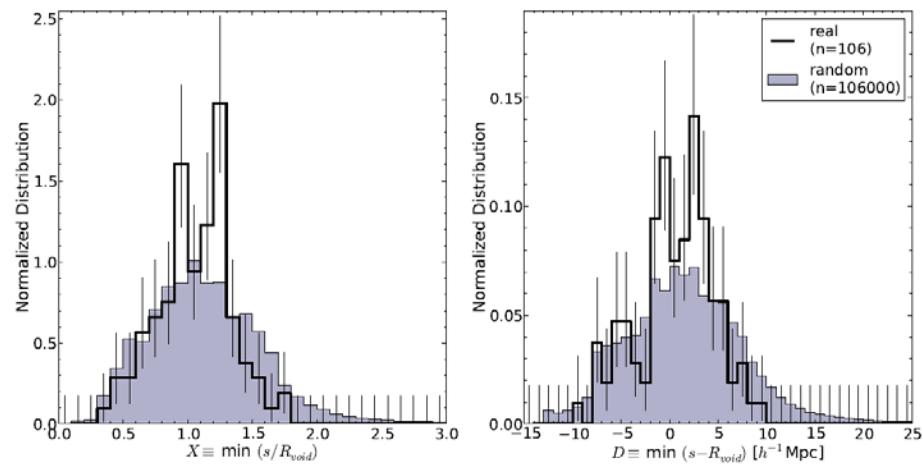
Reflex II cluster catalog (Bohringer et al.) reveals same population of voids in cluster distribution (see talk by Collins).

Voids & the Gaseous Web



Tejos et al. 2012:

HI Ly α absorption systems clearly delineate voids



Voids:
why interesting ?

Voids: Cosmological Significance

- **Cosmic Structure:**

**prominent components of the Cosmic Web,
structurally & dynamically of major influence,
instrumental in spatial organization of the Megaparsec Universe.**

- **Cosmological probe:**

**Voids contain significant amount of information on global
cosmological parameters:**

- void substructure: test GR/modified gravity
- void outflow: dark matter
- void shapes: dark energy
- supervoids: existence; CMB - ISW

- **Galaxy Formation & Evolution:**

**pristine low-density environment ideal for studying galaxy
formation and the effects of cosmic environment on the
evolution of galaxies.**

Voids:

Formation & Structure

Void Formation

Void Evolution

an illustration

cosmology:

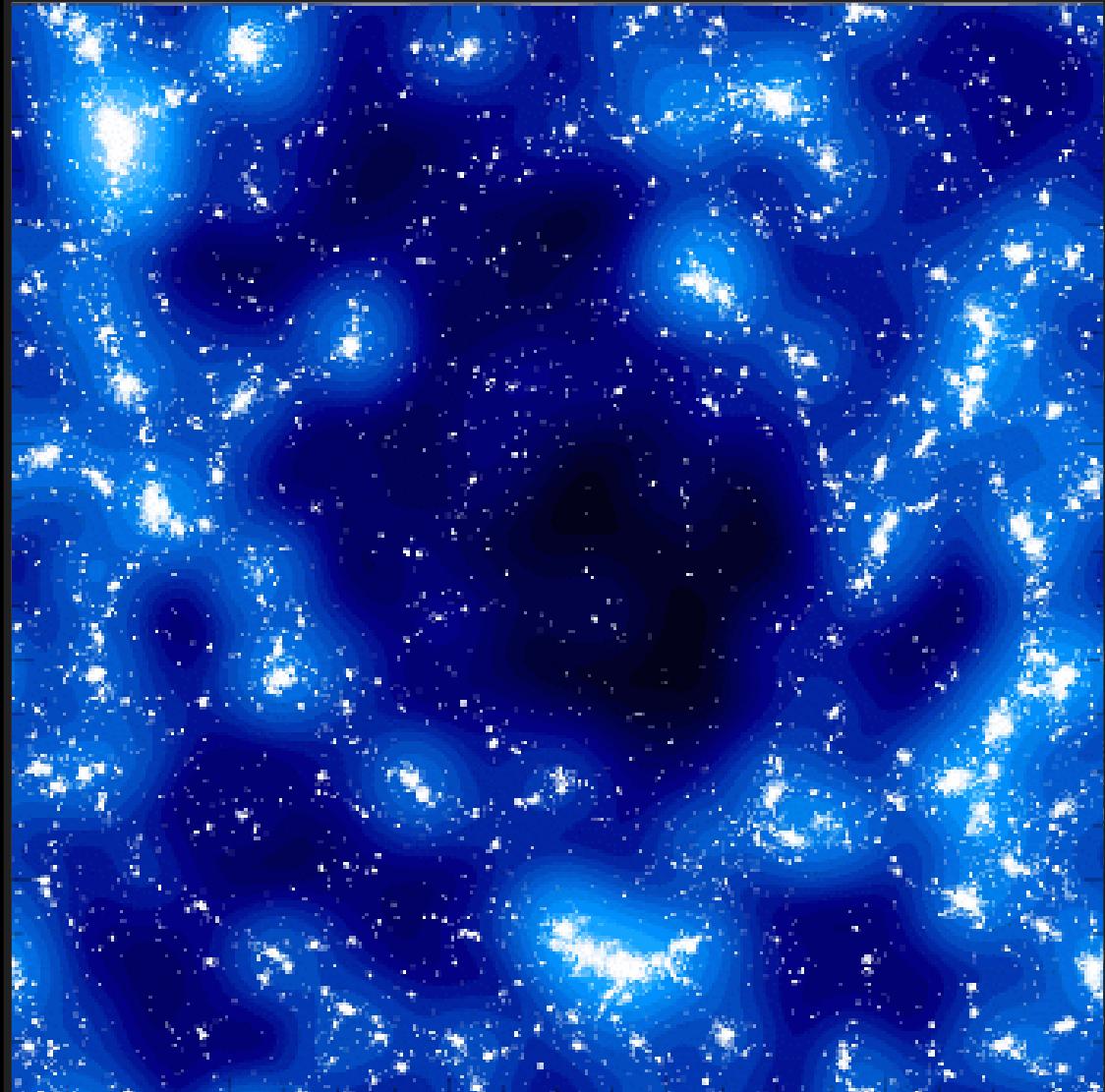
$$\Omega_m = 1.0; \quad H_0 = 70 \text{ km / s / Mpc}$$

initial conditions:

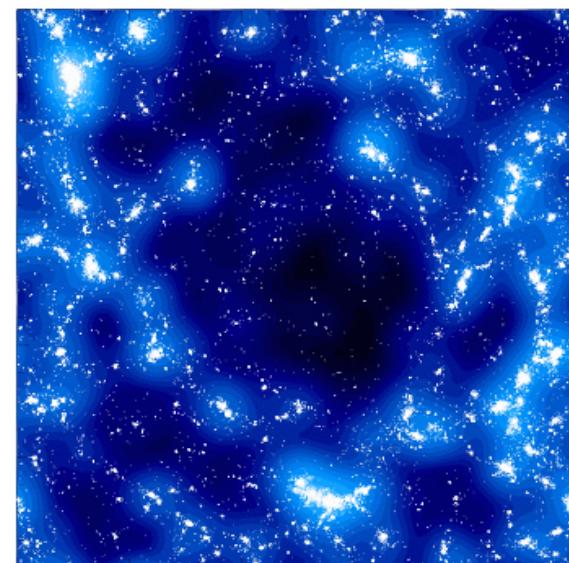
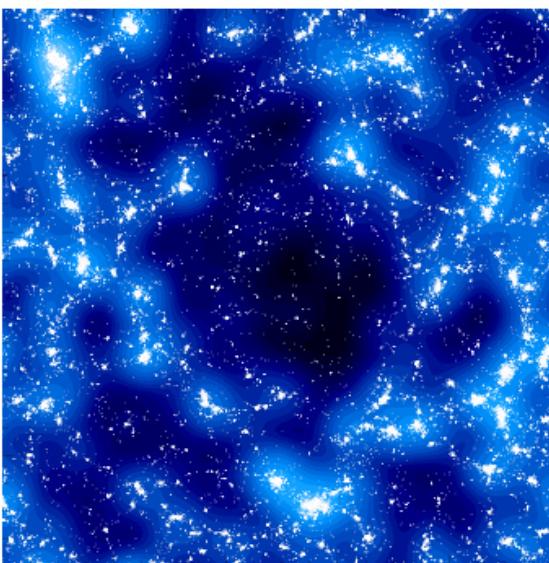
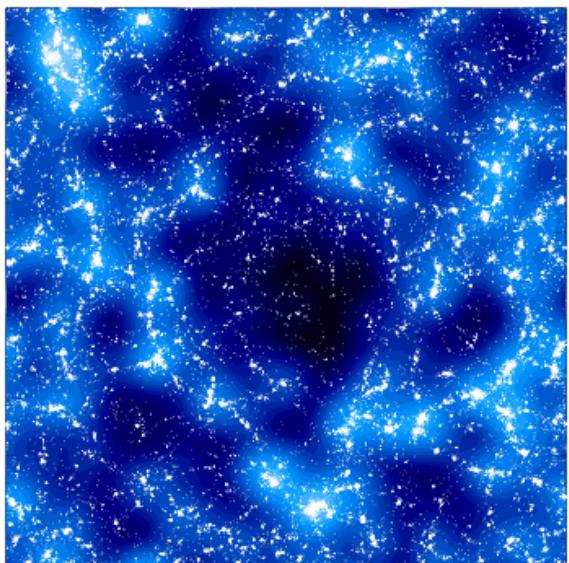
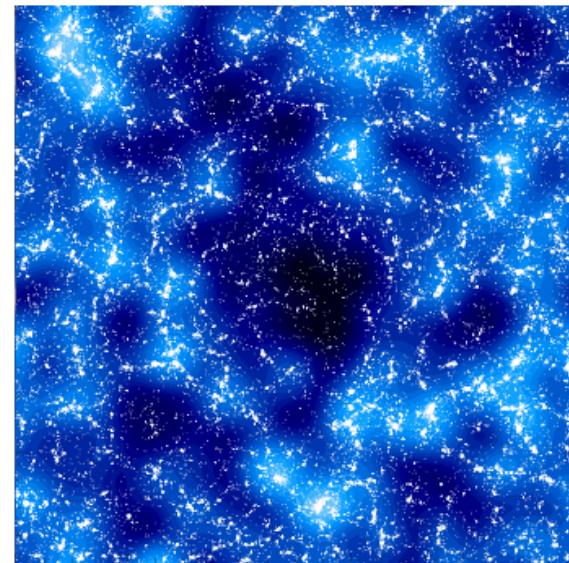
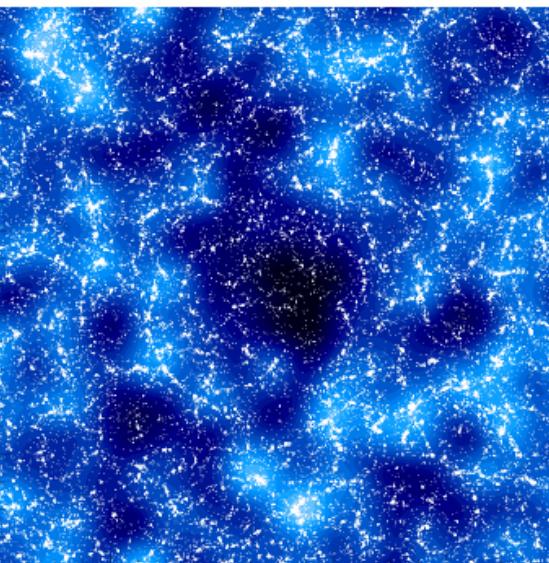
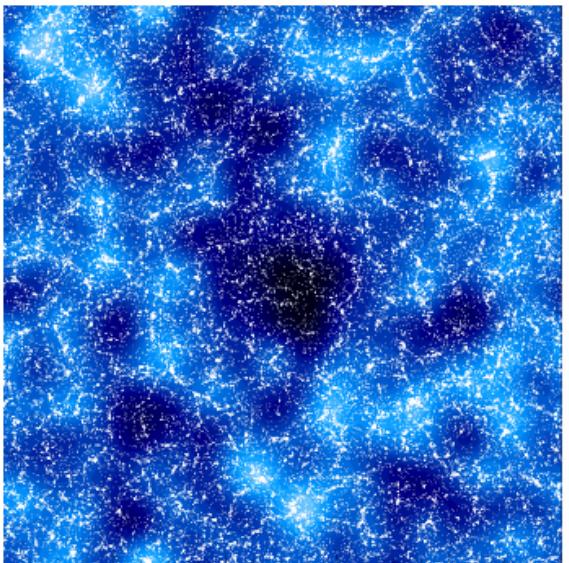
underdensity, Gaussian field

$$R_G \sim 4h^{-1}\text{Mpc}$$

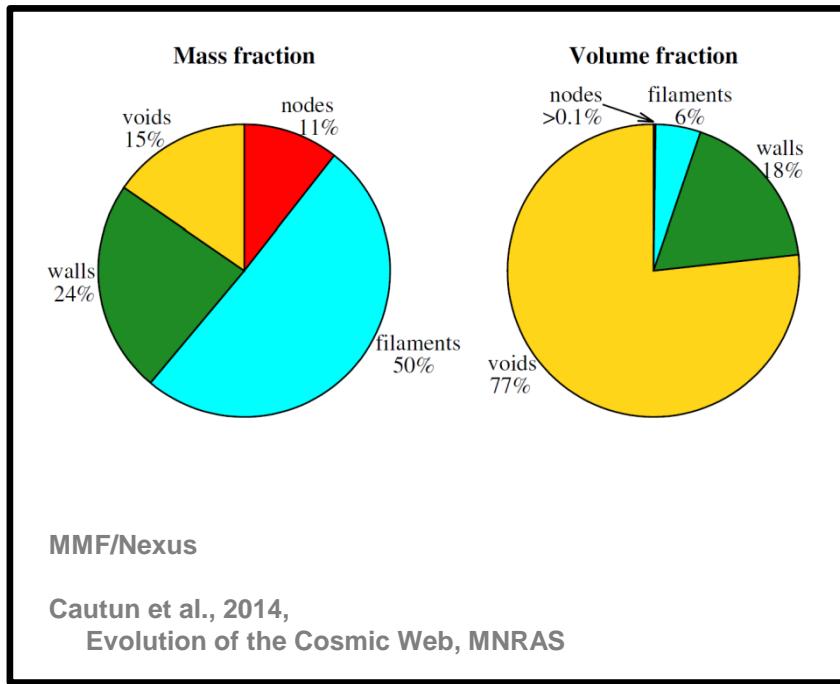
$$P(k) \propto k^{-0.5}$$



Void Formation



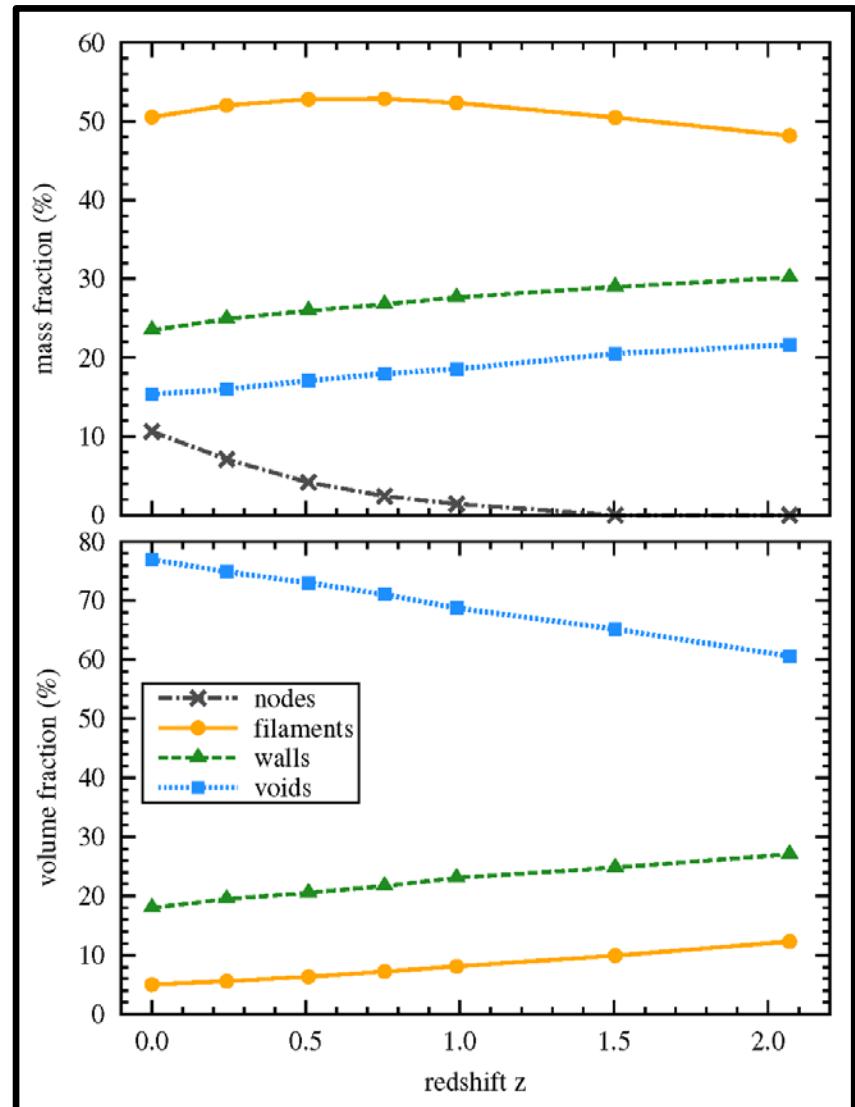
Cosmic Web & Voids



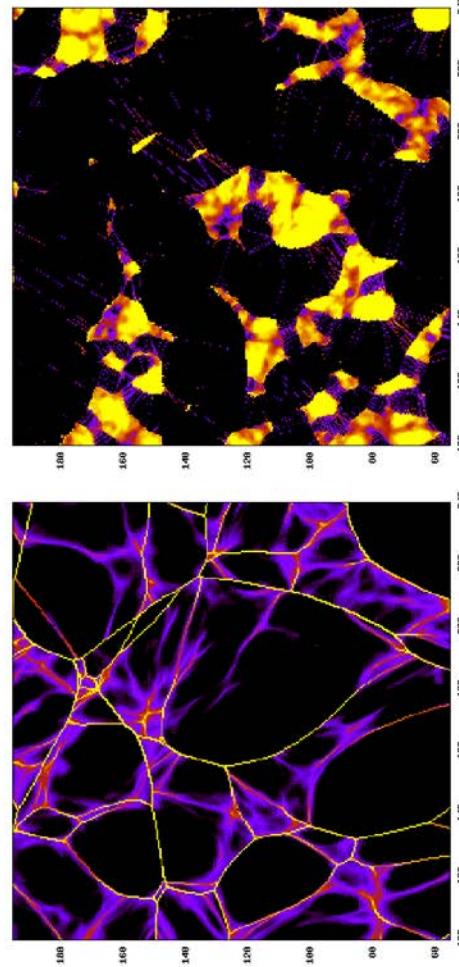
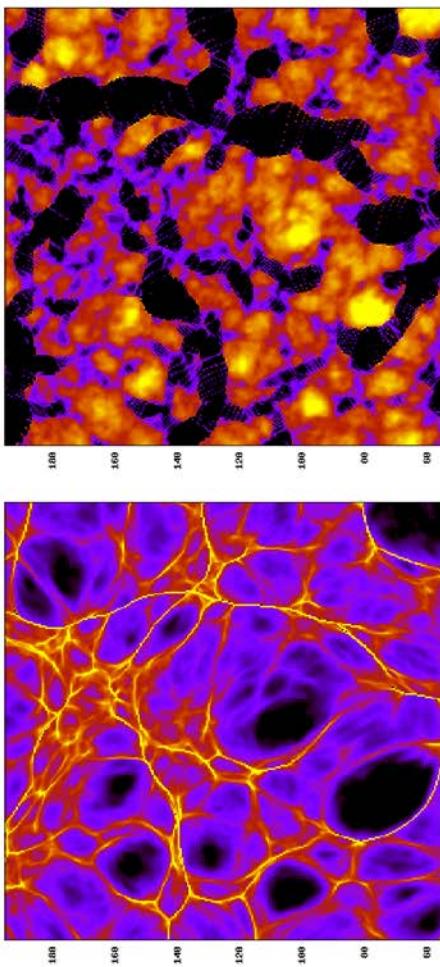
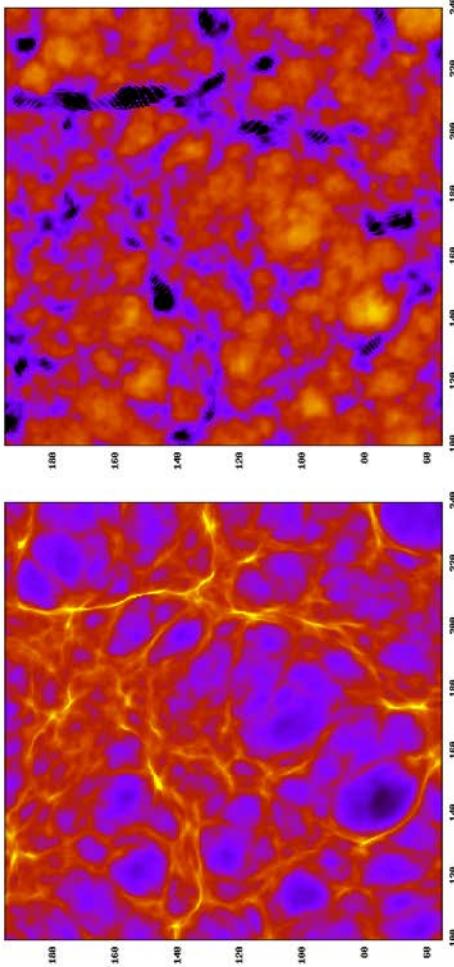
Voids: - occupy most of cosmic volume: 77%
- of mass, only: 15%

Void evolution:

- volume fraction increases with time
(void expansion)
- mass fraction decreases with time
(void evacuation)



Void Formation: Lagrangian View

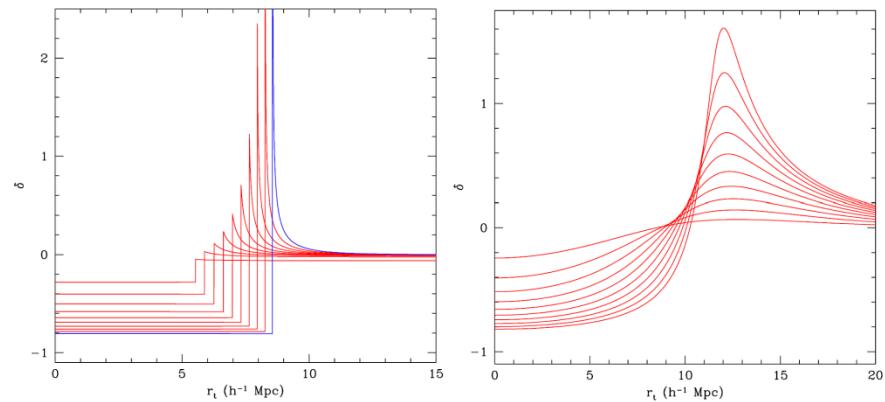


Lagrangian evolution:
void regions shrink

Eulerian Evolution:
void regions expand

Void Evolution: The Perfect Sphere, Tophat as well

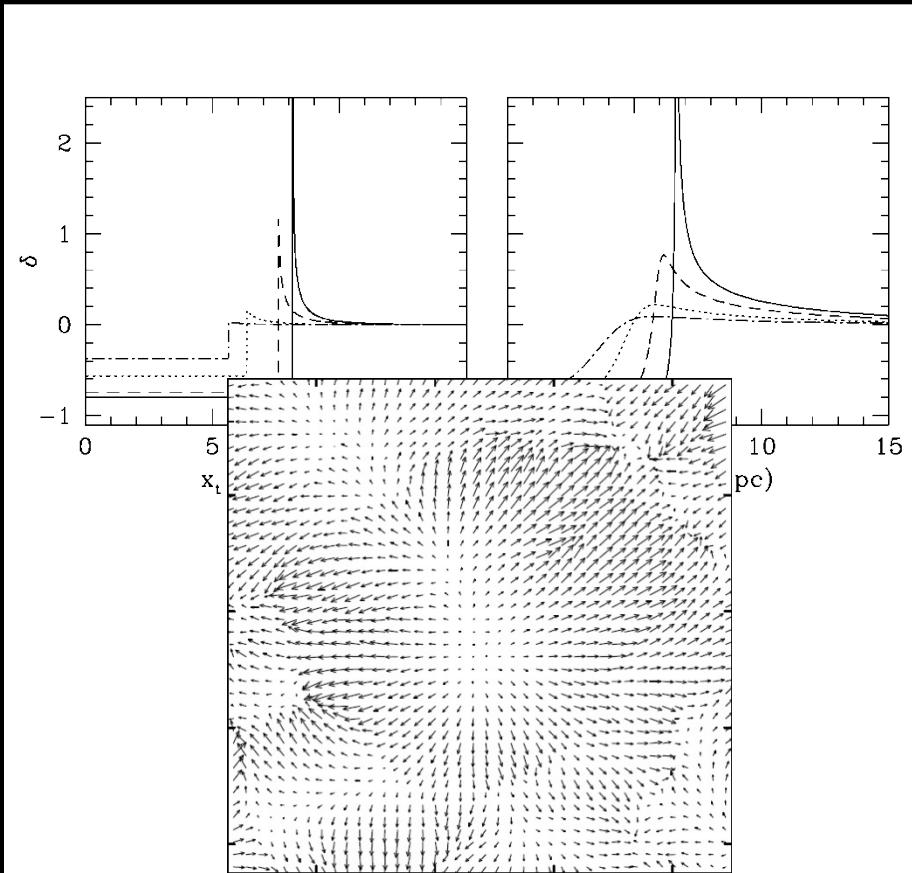
- ``Bubble Theorem''
Voids become increasingly spherical, due to anisotropic outward directed force
- Tophat Configuration
Any initial configuration tends towards “bucket” shape
- Density Ridge
Except for gentlest initial density profiles, a ridge forms



Superhubble Expansion

- Superhubble Expansion tending towards “bucket” shape, the void outflow is one with uniform velocity divergence

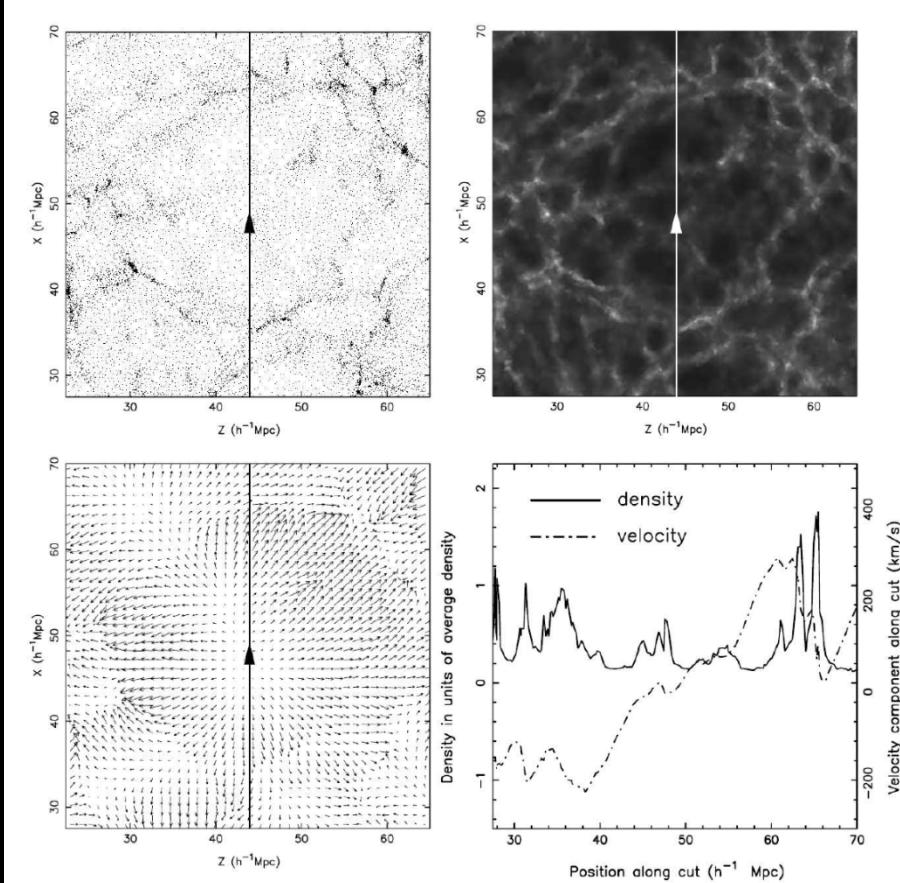
$$\theta = \frac{1}{H} (\nabla \cdot \vec{v}) \quad \Rightarrow \quad \theta_{\max} = 1.5 \Omega^{0.6}$$



Superhubble Expansion

- Superhubble Expansion tending towards “bucket” shape, the void outflow is one with uniform velocity divergence

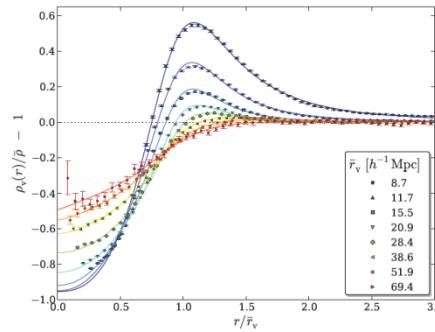
$$\theta = \frac{1}{H} (\nabla \cdot \vec{v}) \quad \Rightarrow \quad \theta_{\max} = 1.5 \Omega^{0.6}$$



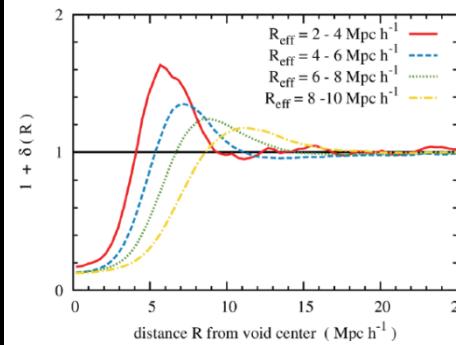
Void Density Profile

- Tophat Configuration
Any initial configuration tends towards “bucket” shape
- Density Ridge
Except for gentlest initial density profiles, a ridge forms

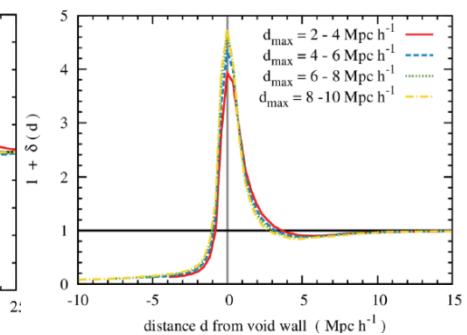
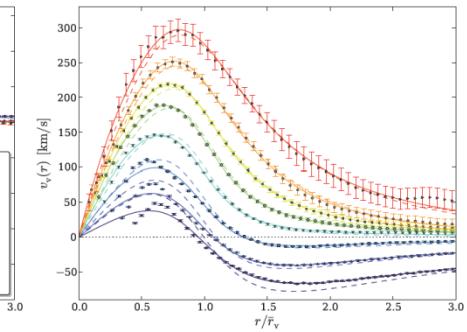
Is there a universal void density profile ?



e.g. Hamaus et al. 2014

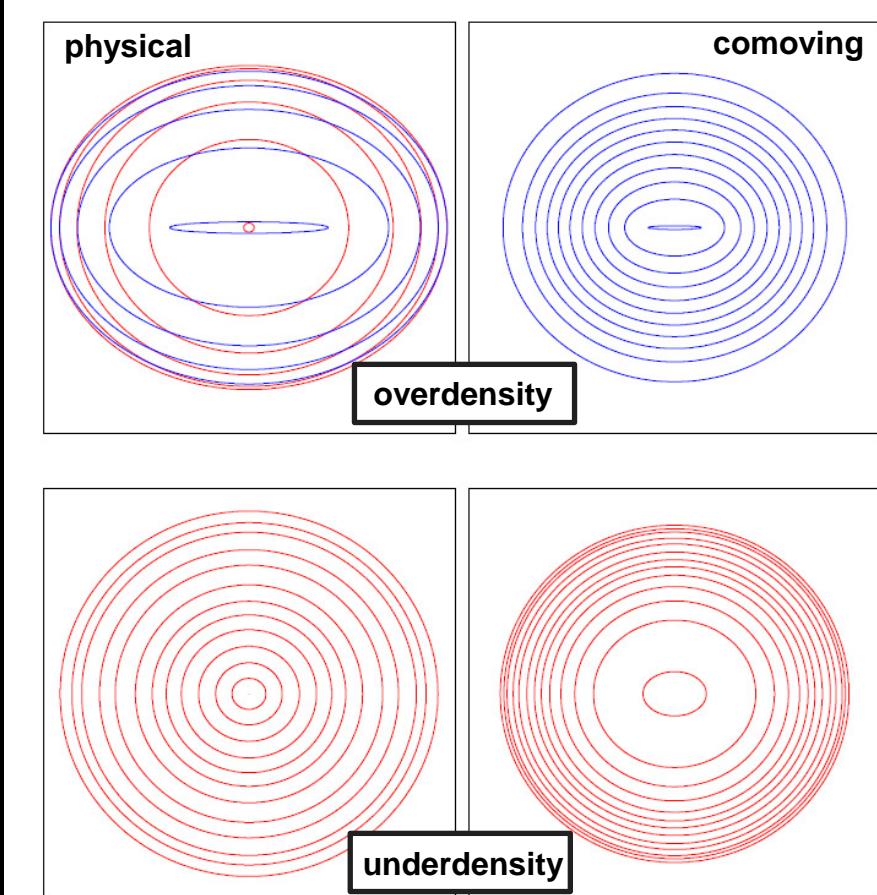
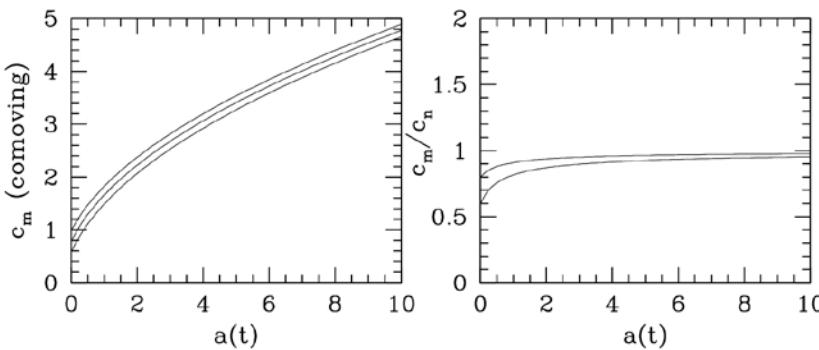


Cautun et al. 2014



Bubble Theorem

- **Bubble Theorem (Icke 1984)**
Isolated voids tend to become more spherical as they expand and evolve



Void Shapes & Environment

- Bubble Theorem Revisited:

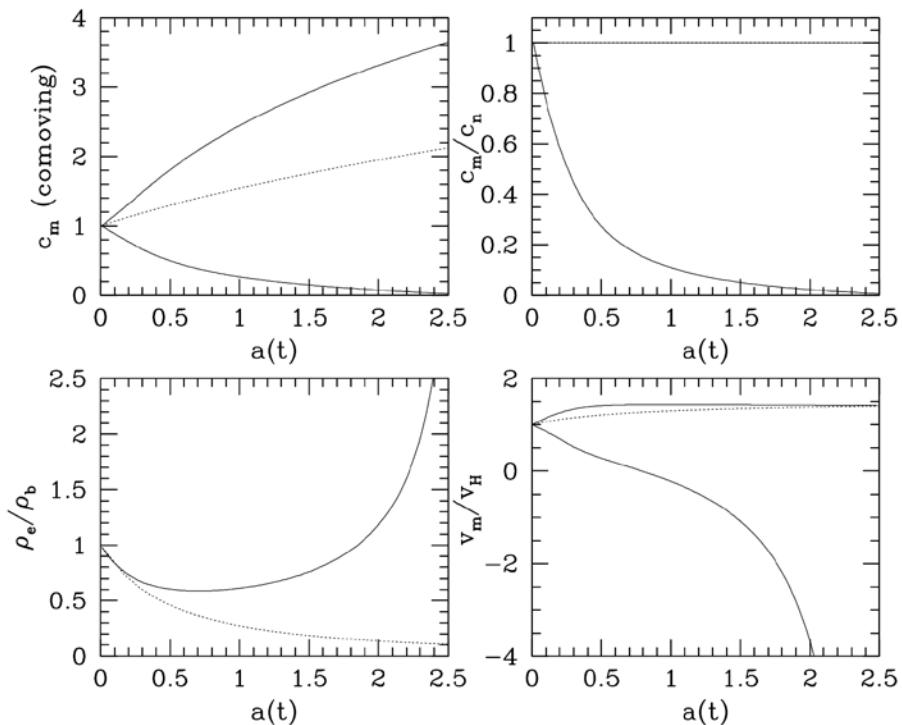
voids will not be spherical:

- Voids never isolated: run into neighbours
- Void evolution largely dominated by large scale (tidal) environment:

voids always represent restricted density fluctuation: $|\bar{\rho}| < 1$

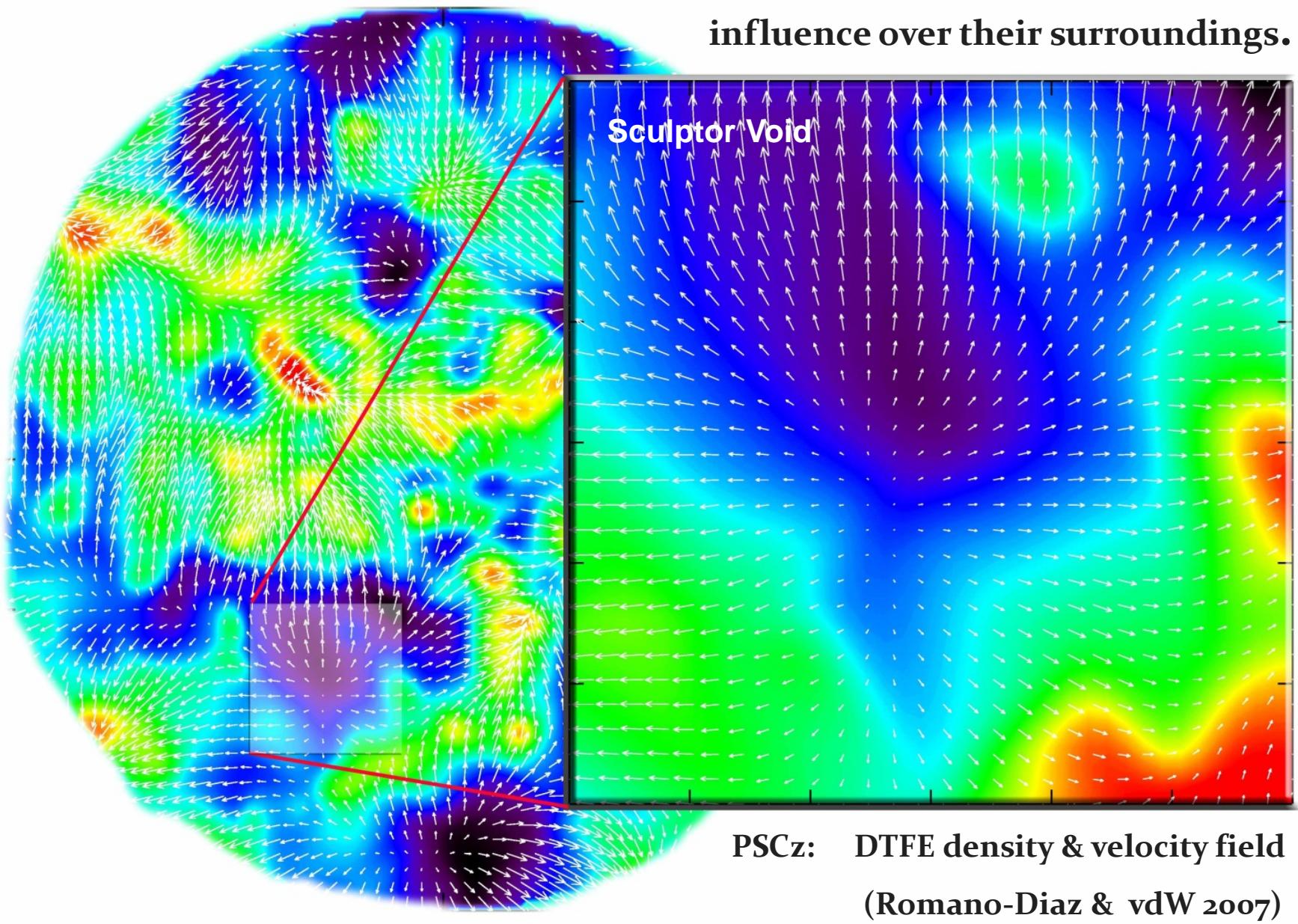
Evolution homogeneous ellipsoidal void in external tidal field T_{kl}

$$\frac{d^2 R_m}{dt^2} = -2\pi G \left[\alpha_m \rho_e + \left(\frac{2}{3} - \alpha_m \right) \rho_u \right] R_m - T_{mm} R_m$$

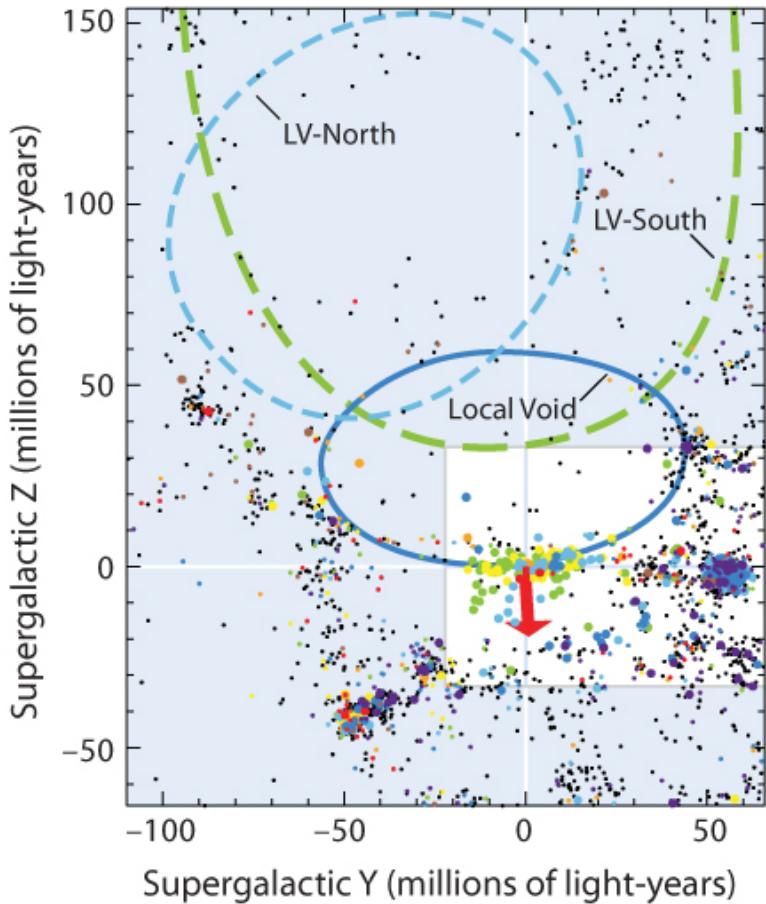


Void Dynamics

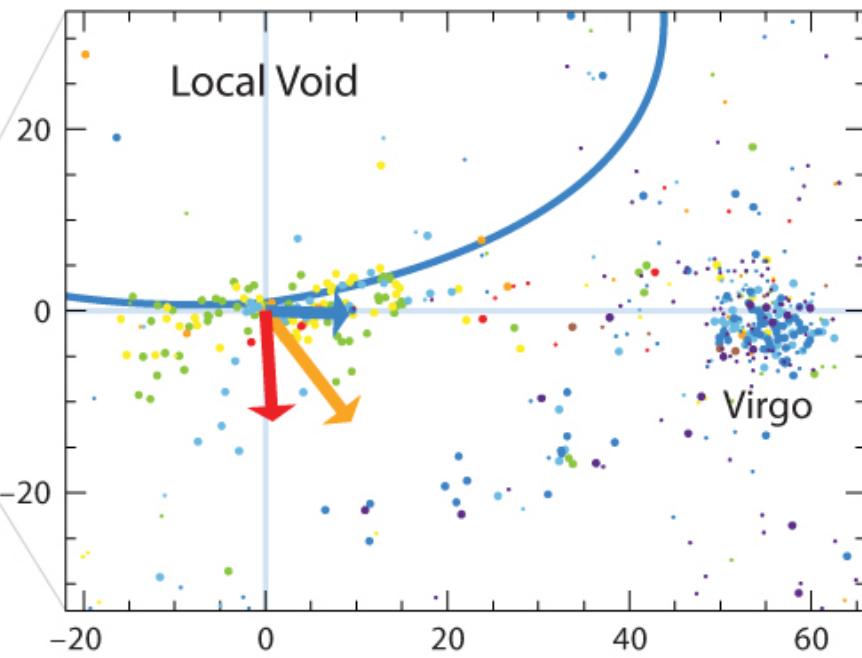
Voids exert a repulsing dynamical influence over their surroundings.



Push of the Local Void

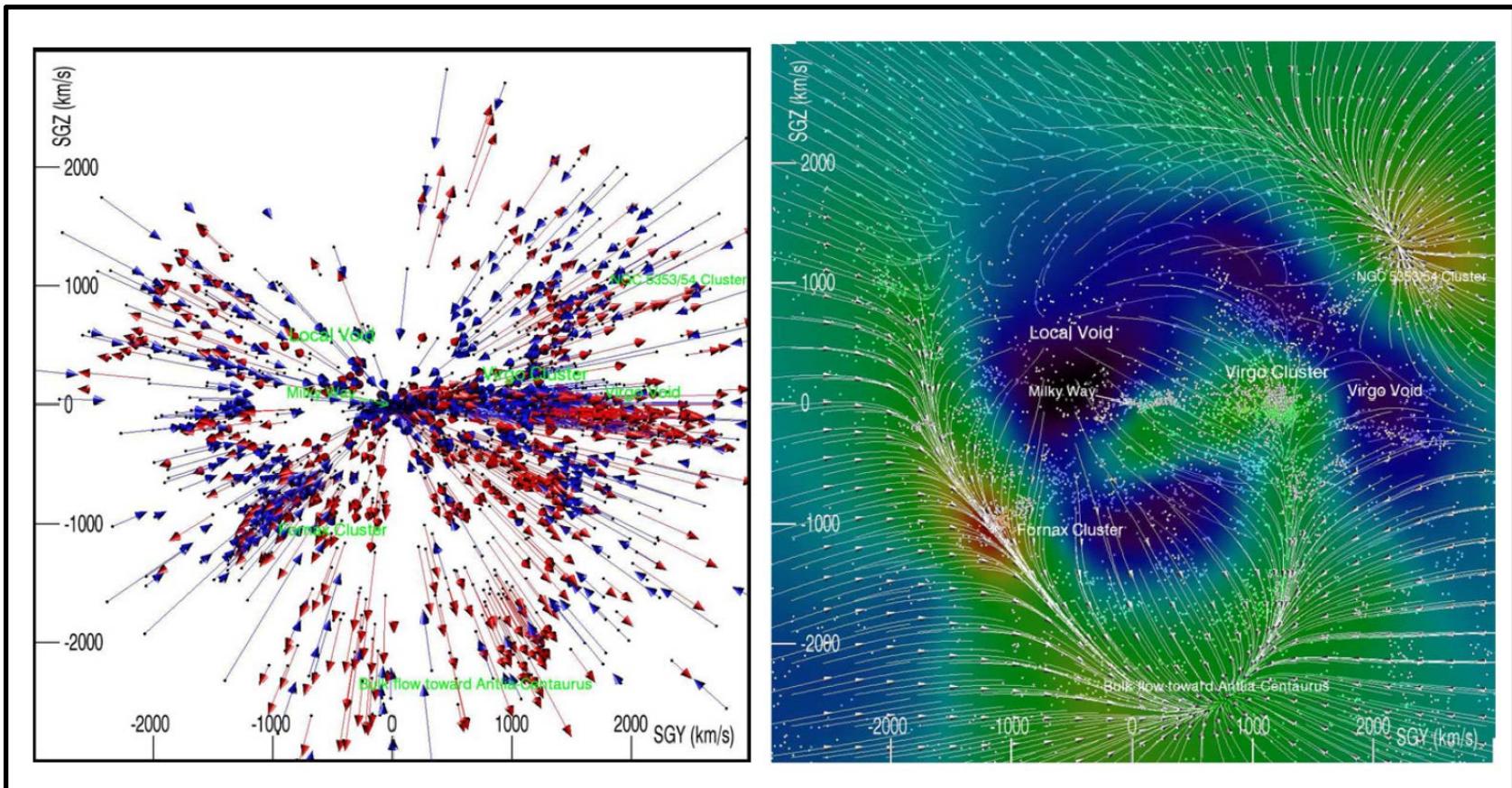


Our motion with the respect to galaxies in the Local Supercluster *Tully et al. 2008, ApJ, 676, 184*



Tully et al. 2008:
Local Void pushes with ~ 260 km/s against our local neighbourhood

Local Void Outflow



Courtois et al. 2013
Local void expansion in Cosmoflows-2

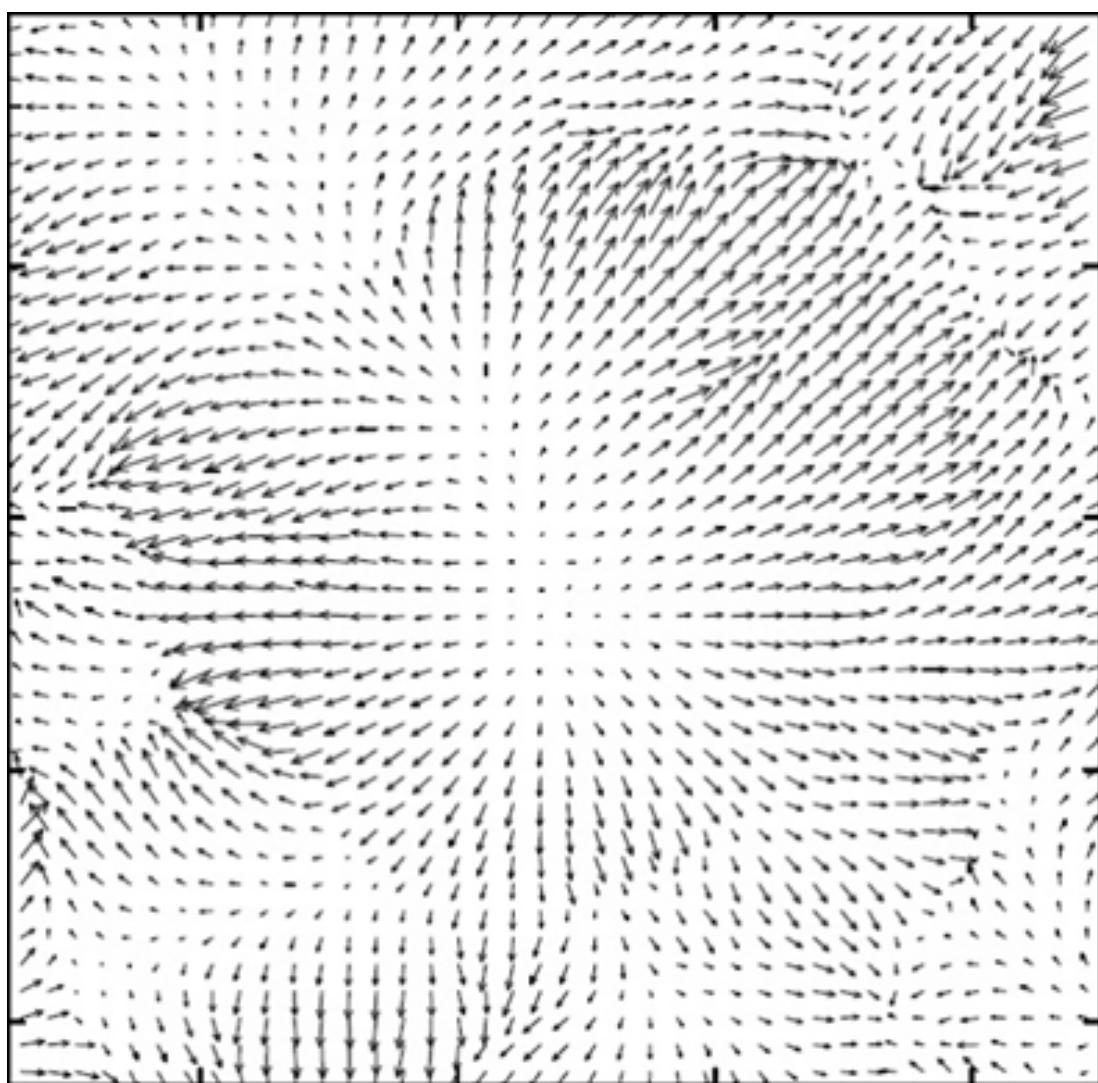
Nonlinear Void Dynamics

Formation of Cosmic Web
nonlinear process:

$$\nabla \cdot \vec{v} = \frac{3}{2} Ha f(\Omega_m) \left\{ 1 - (1 + \delta(\vec{x}))^{2/3} \right\}$$

$$\nabla \cdot \vec{v} = -f(\Omega_m) \frac{\delta(\vec{x})}{1 + 0.18\delta(\vec{x})}$$

$$\theta = \frac{1}{H} (\nabla \cdot \vec{v}) \quad \Rightarrow \quad \theta_{\max} = 1.5 \Omega^{0.6}$$



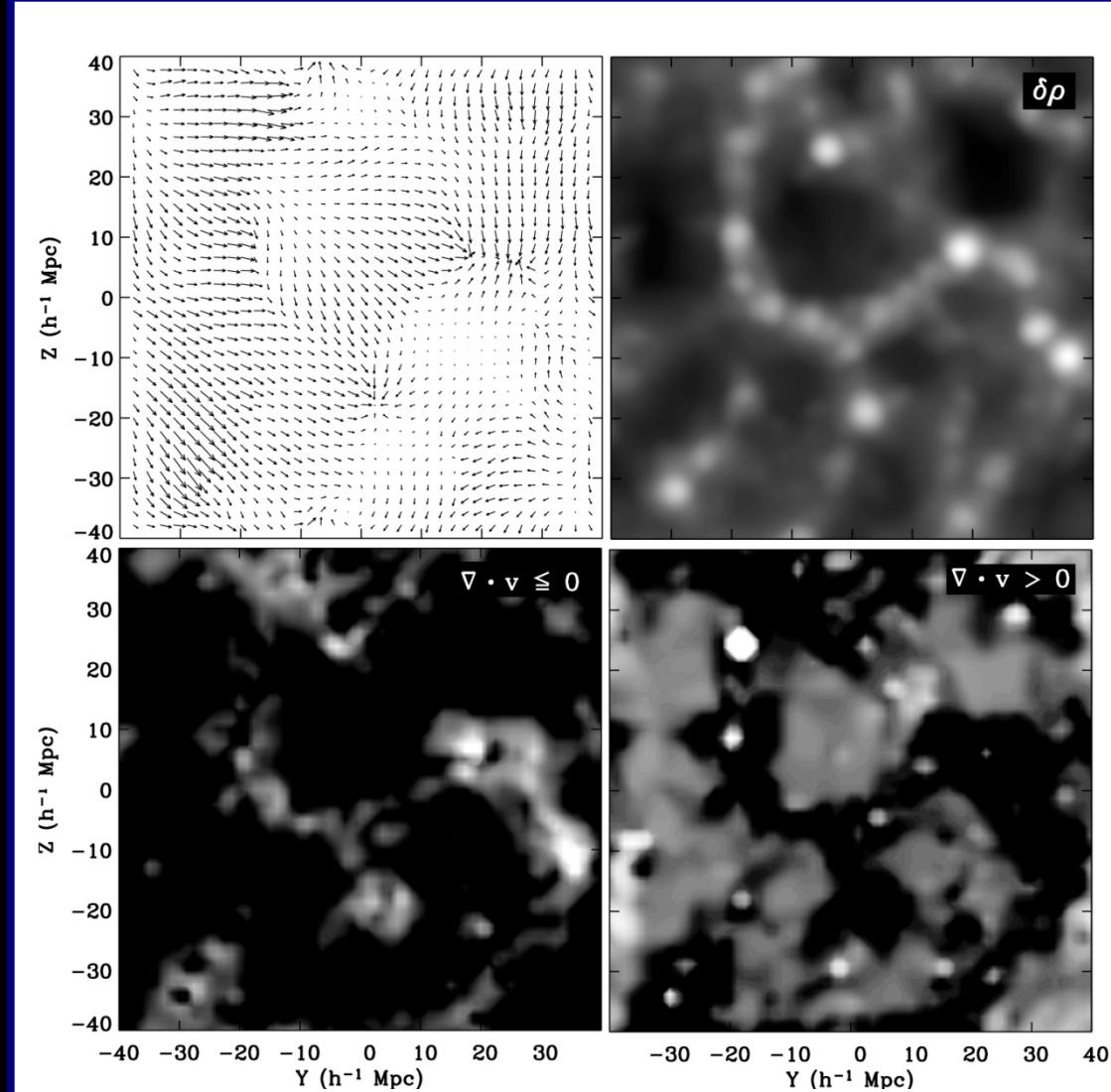
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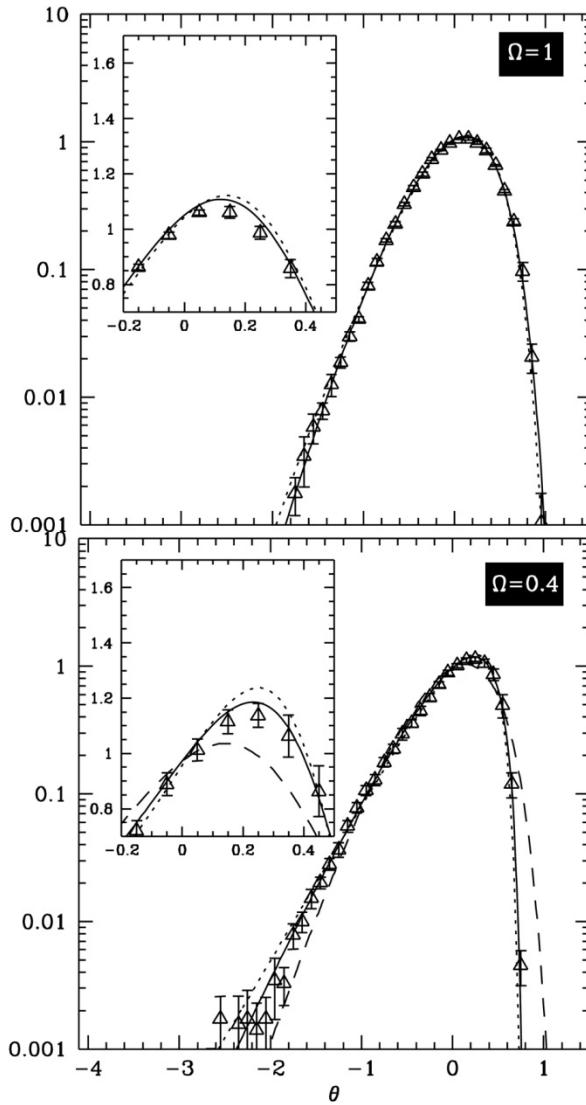
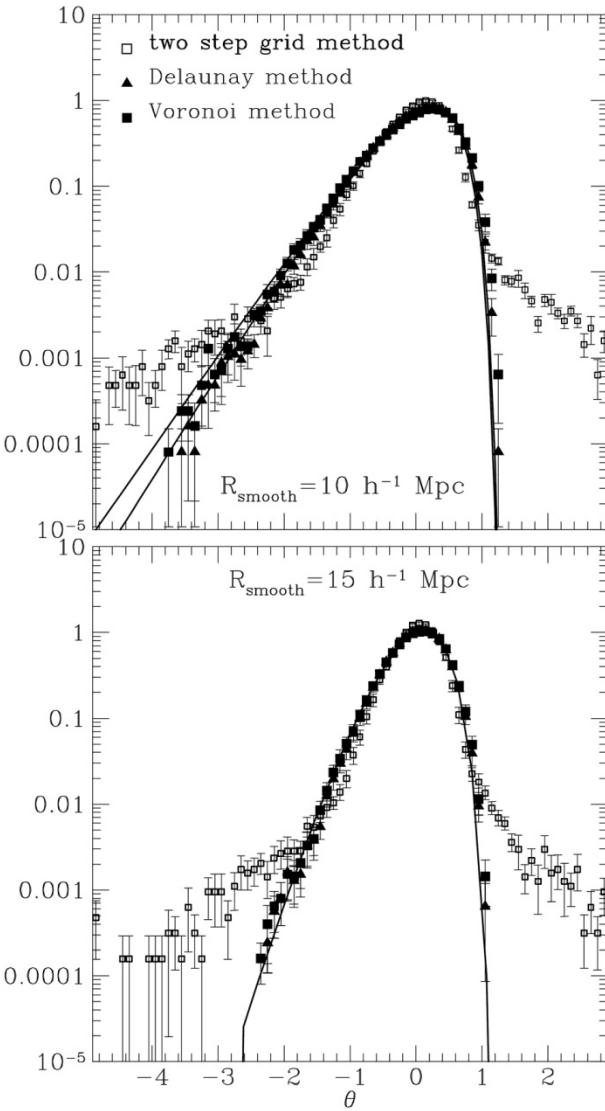
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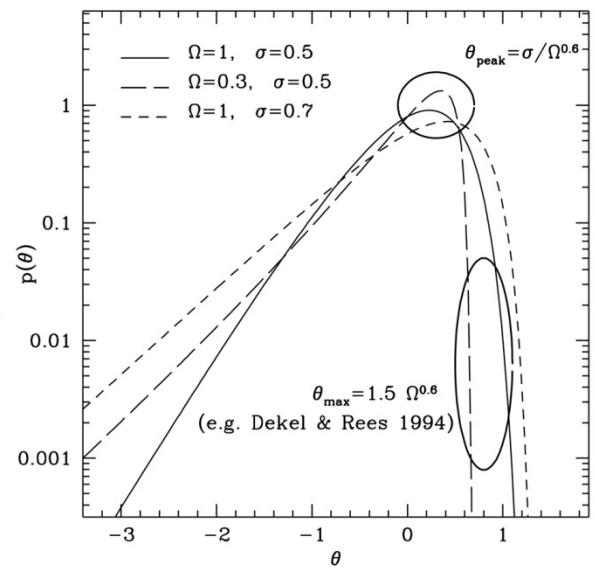
$$\theta = \frac{1}{H} (\nabla \cdot \vec{v}) \quad \Rightarrow \quad \theta_{\max} = 1.5 \Omega^{0.6}$$



DTFE Velocity Field Divergence



Regions of positive velocity divergence:
Outflow velocities mostly corresponding to underdense regions.



Void Identification

the Watershed Void Finder

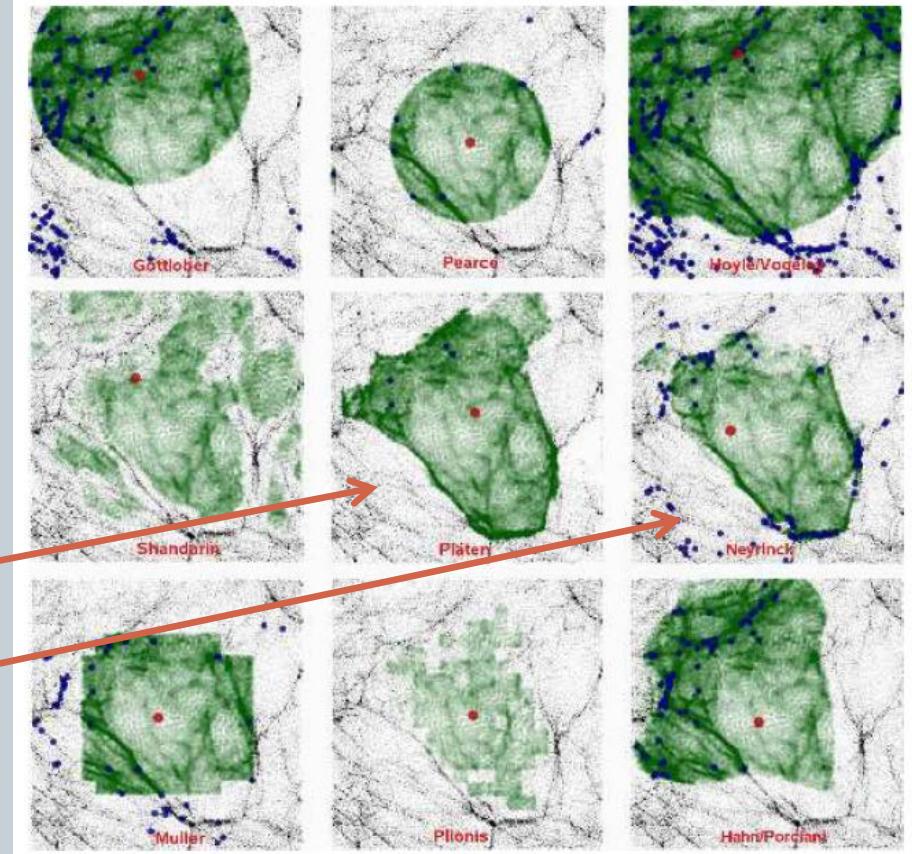
No exact definition of a void!
→ broad range and variety of
void detection techniques

Watershed void finders:

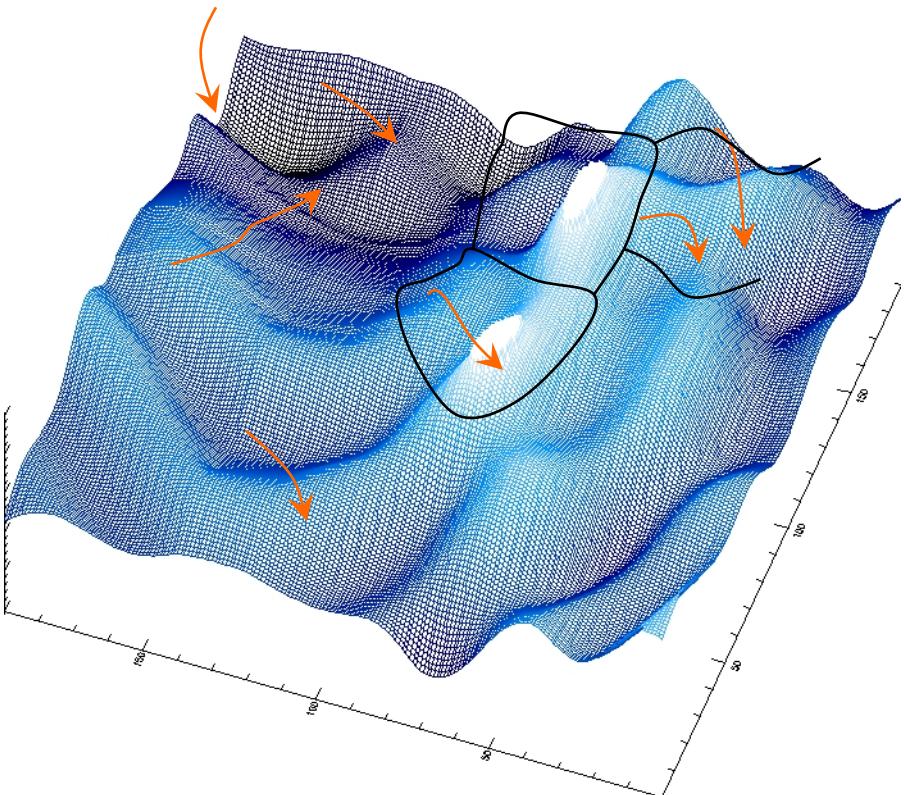
- closely follows real geometry cosmic web
- no assumptions geometry void
- no user defined parameters

→ Watershed Void Finder (WVF)
by *Platen et al.*, 2007.

ZOBOV
by *Neyrinck*, 2008



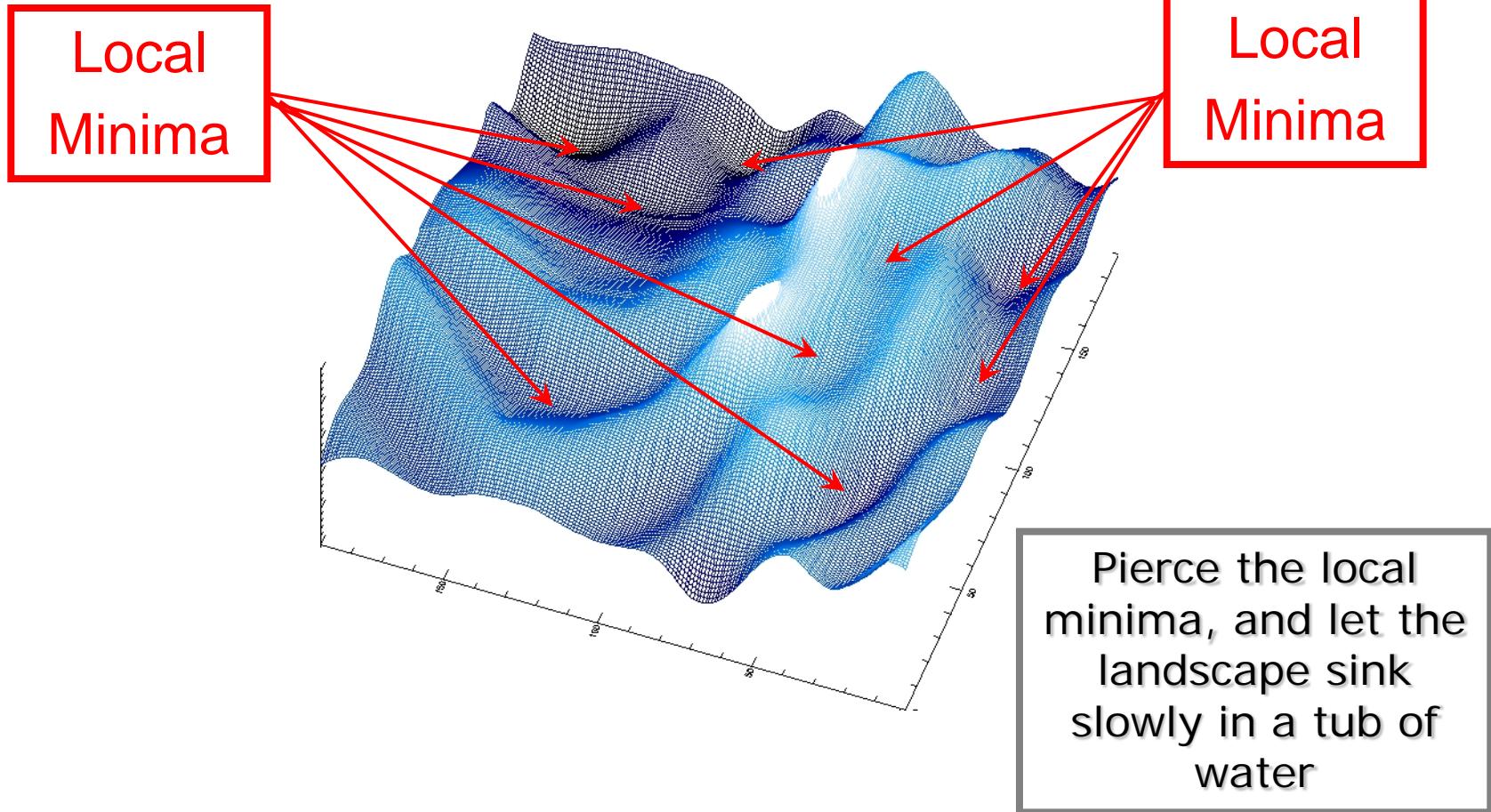
Colberg et al., 2008
Void Comparison Project



Following the water-flow into the distinct catchment basins.

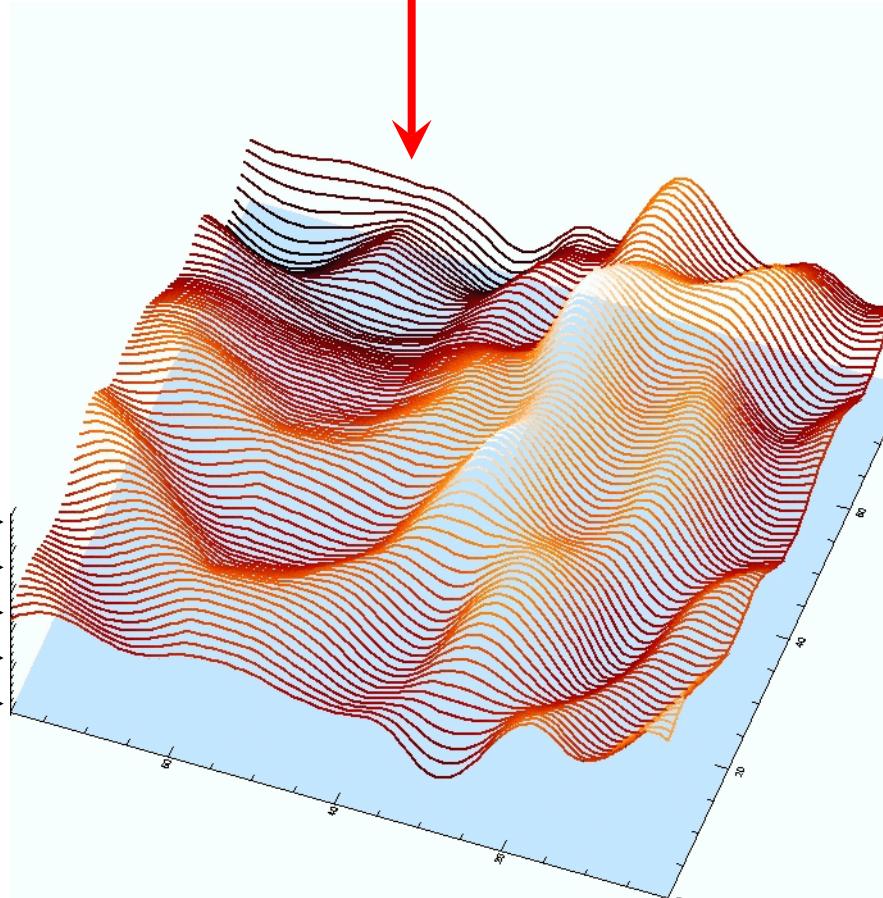
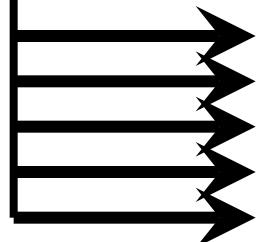
Each basin belonging to one individual minima defines one region

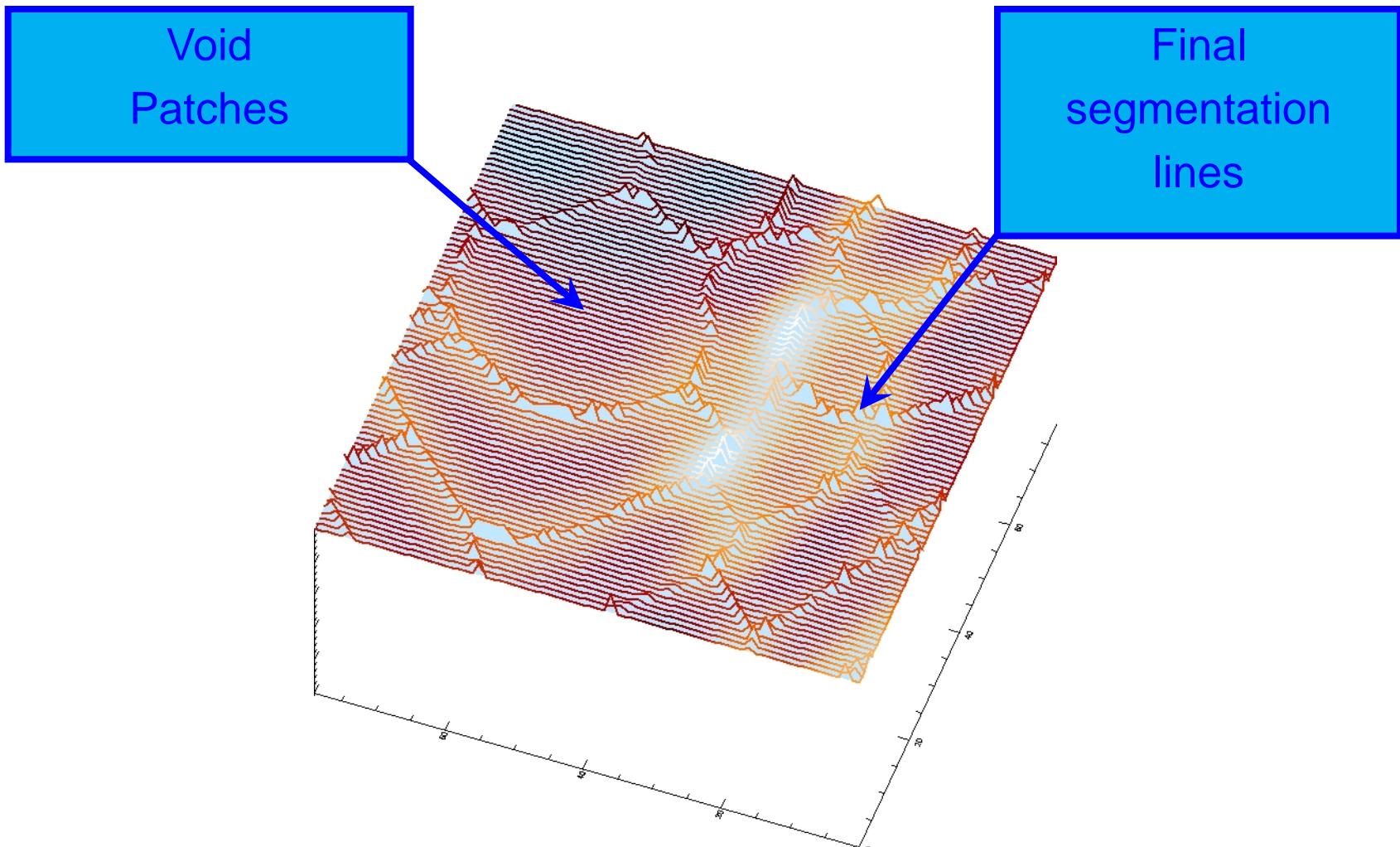
Surface of Density Field



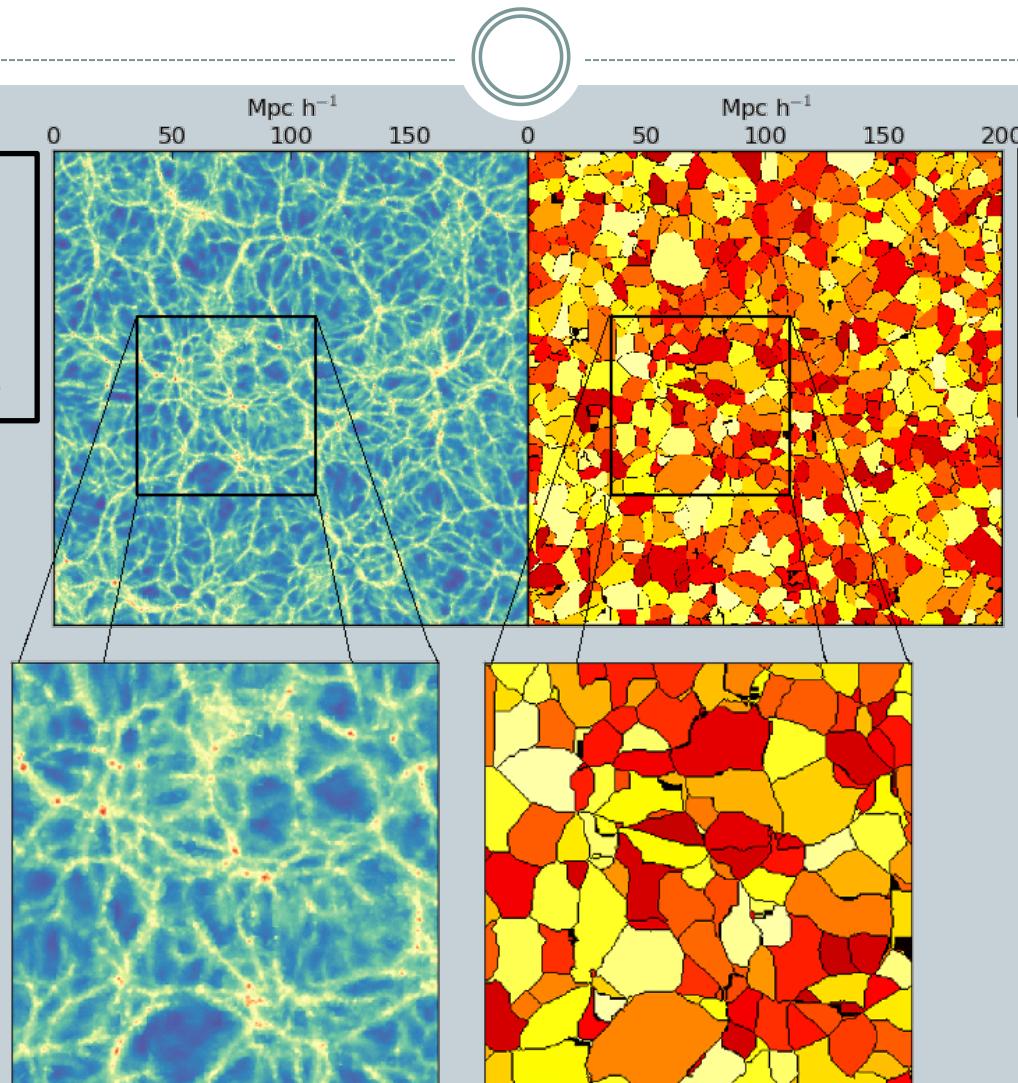
Every time two different flooding basins meet we draw a dividing wall

Flooding the Density Field





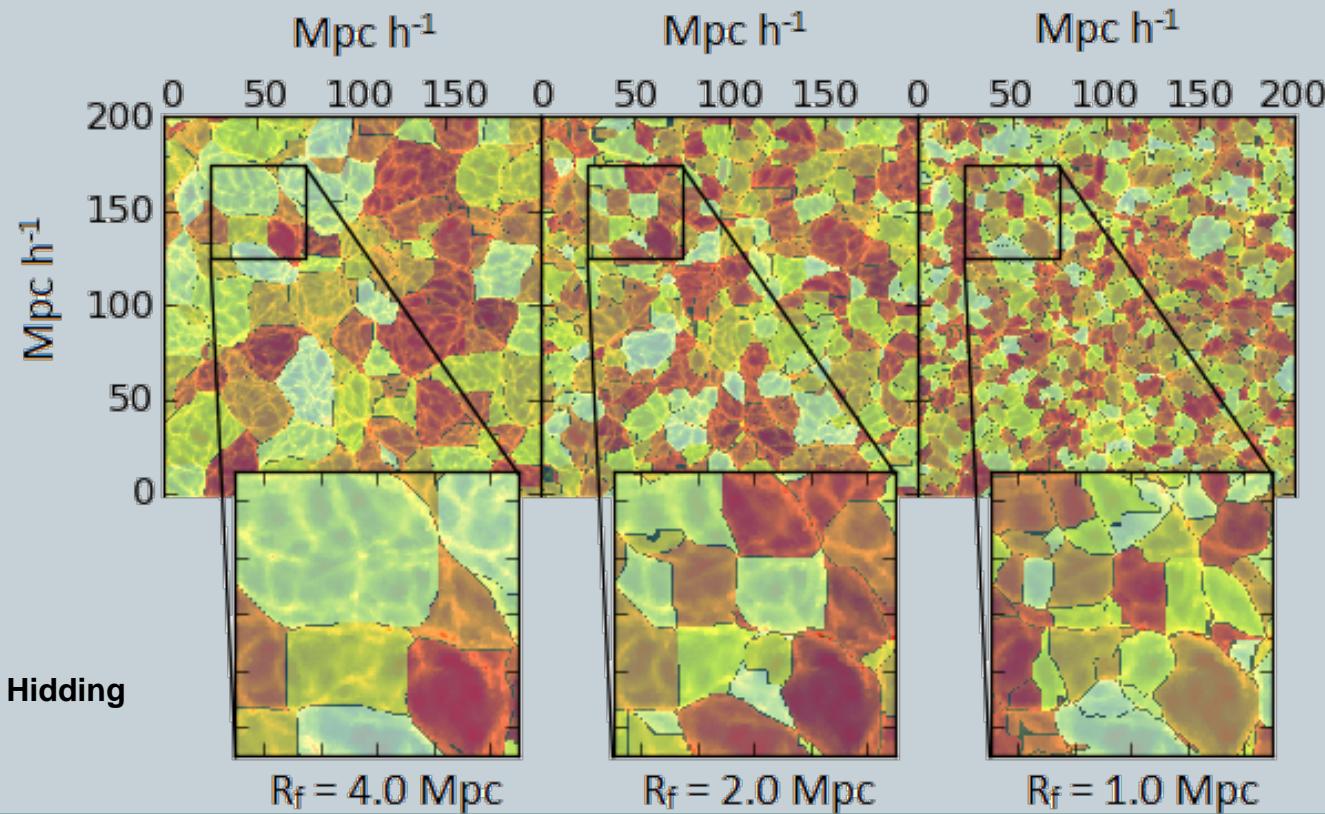
WVF: Watershed Void Finder



The Multiscale Watershed Void Finder



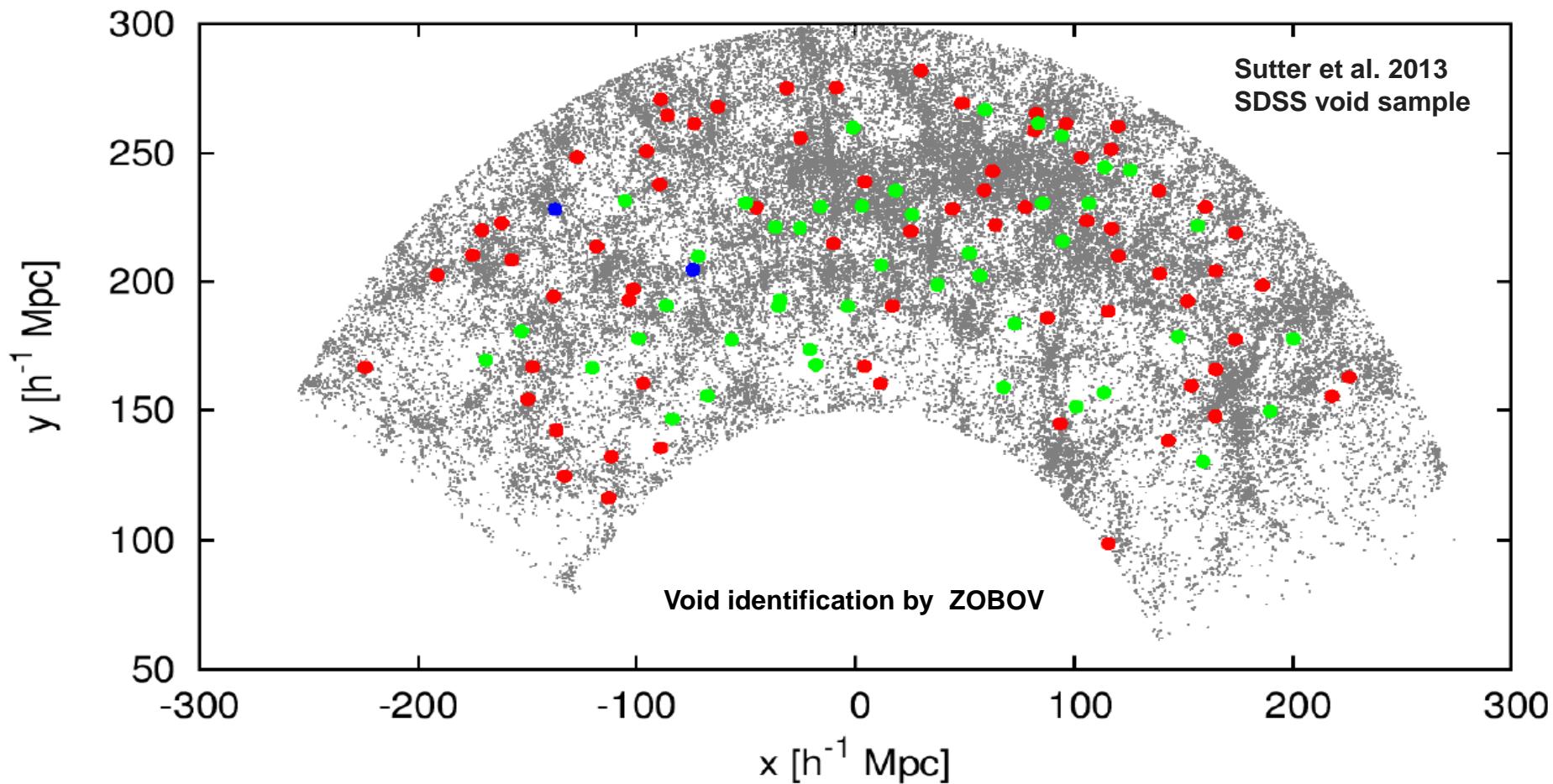
Multiscale void population:
void population as a function of filter radius



Void Catalogues

SDSS Void Sample

Sample: dim2



$R = 5-15 h^{-1} \text{ Mpc}$
 $R = 15-25 h^{-1} \text{ Mpc}$



$R = 25-45 h^{-1} \text{ Mpc}$



Choosing emptiness criterion



Observational void sizes depend strongly on galaxy sample used to define voids

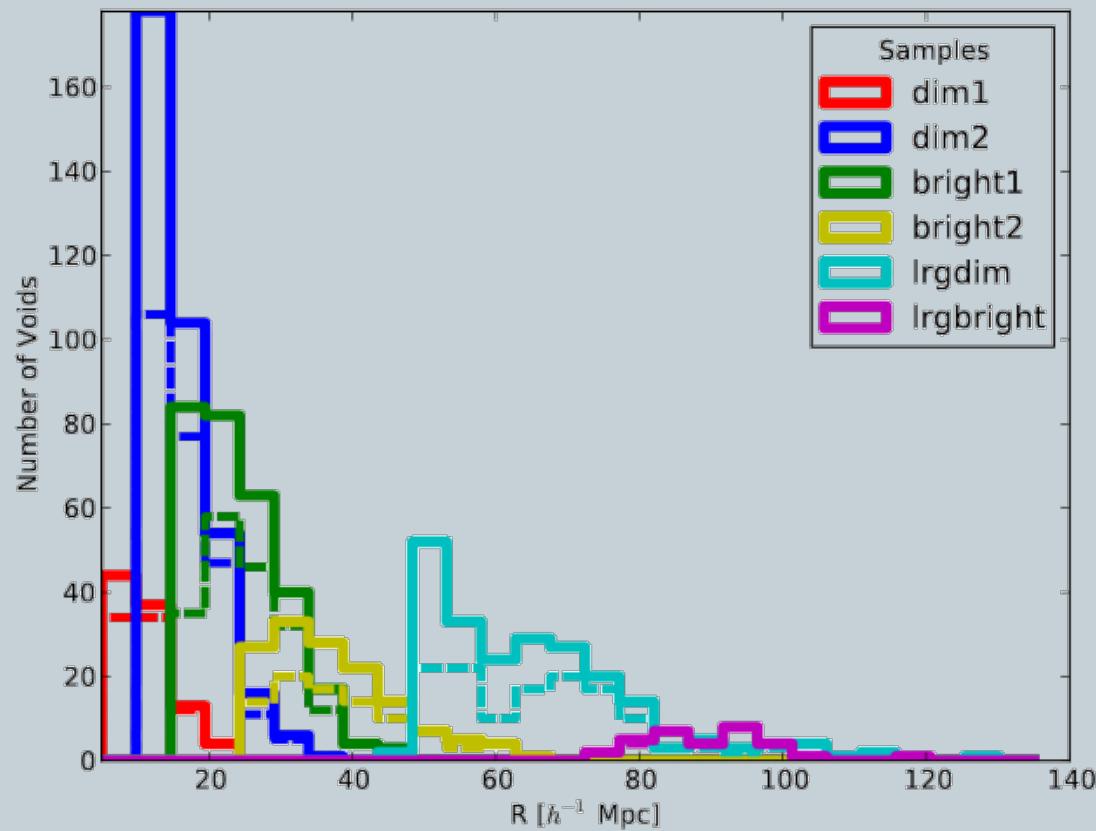
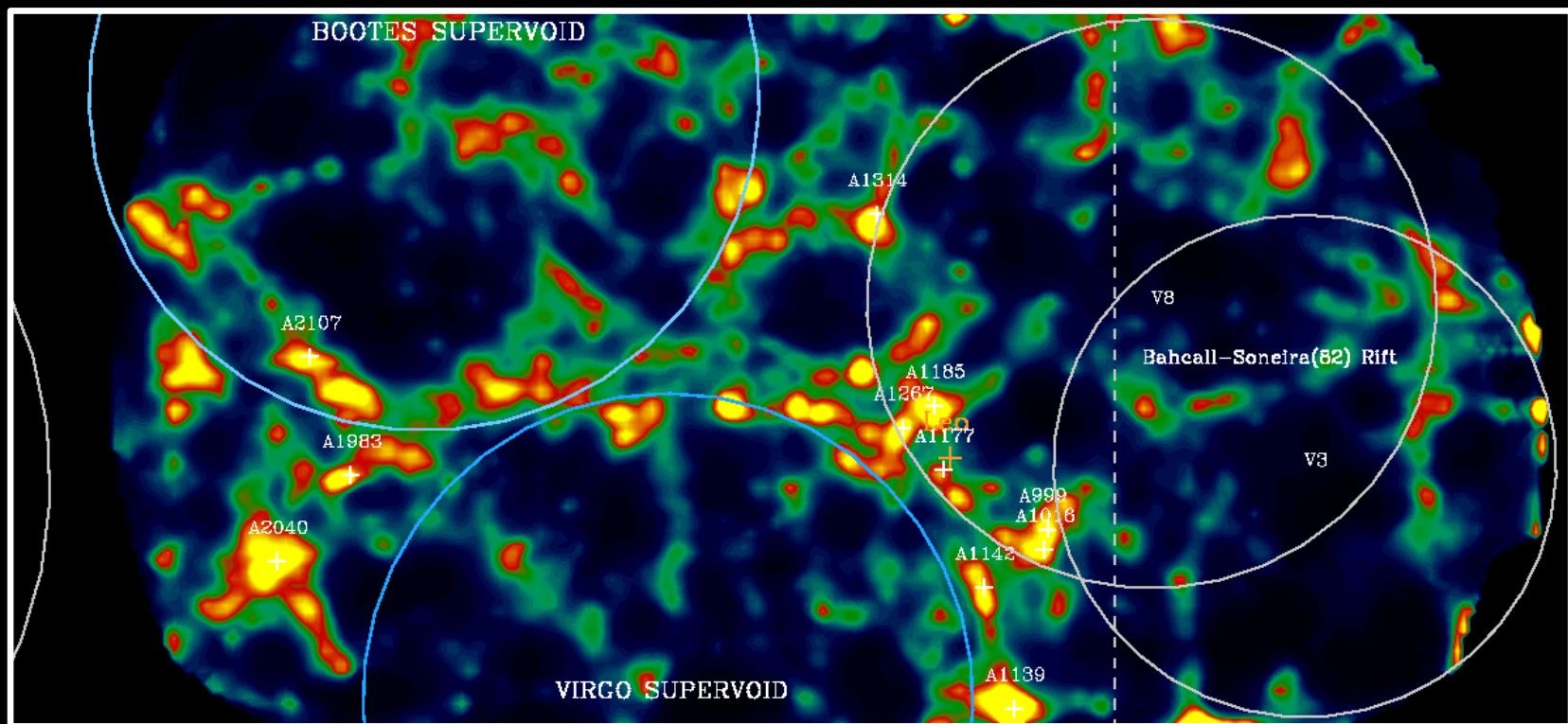


Figure from *Sutter et al., 2013*

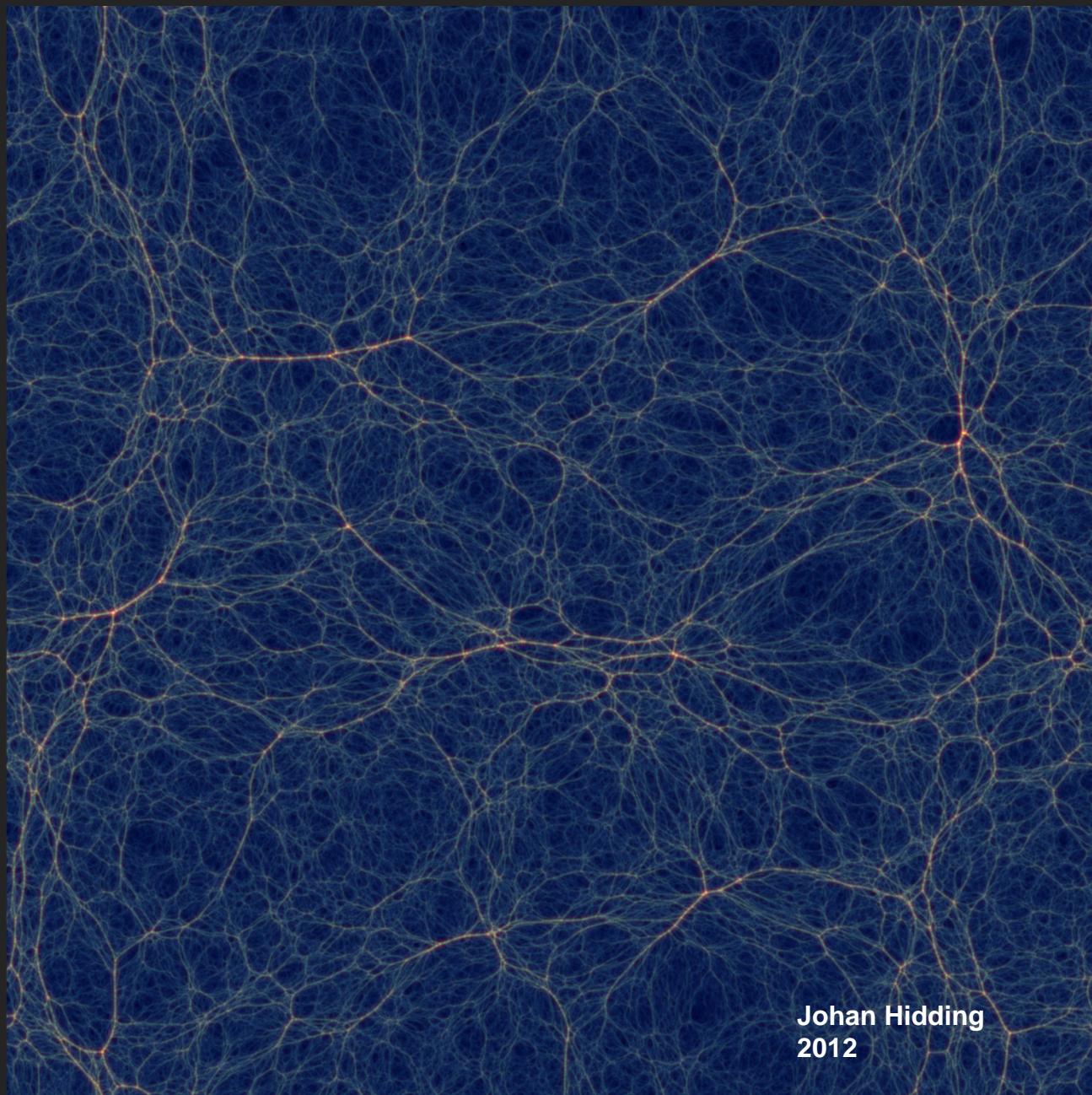
Voids:

Multiscale Infrastructure

Bootes Void: Substructure

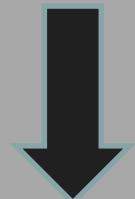


Platen et al. 2009



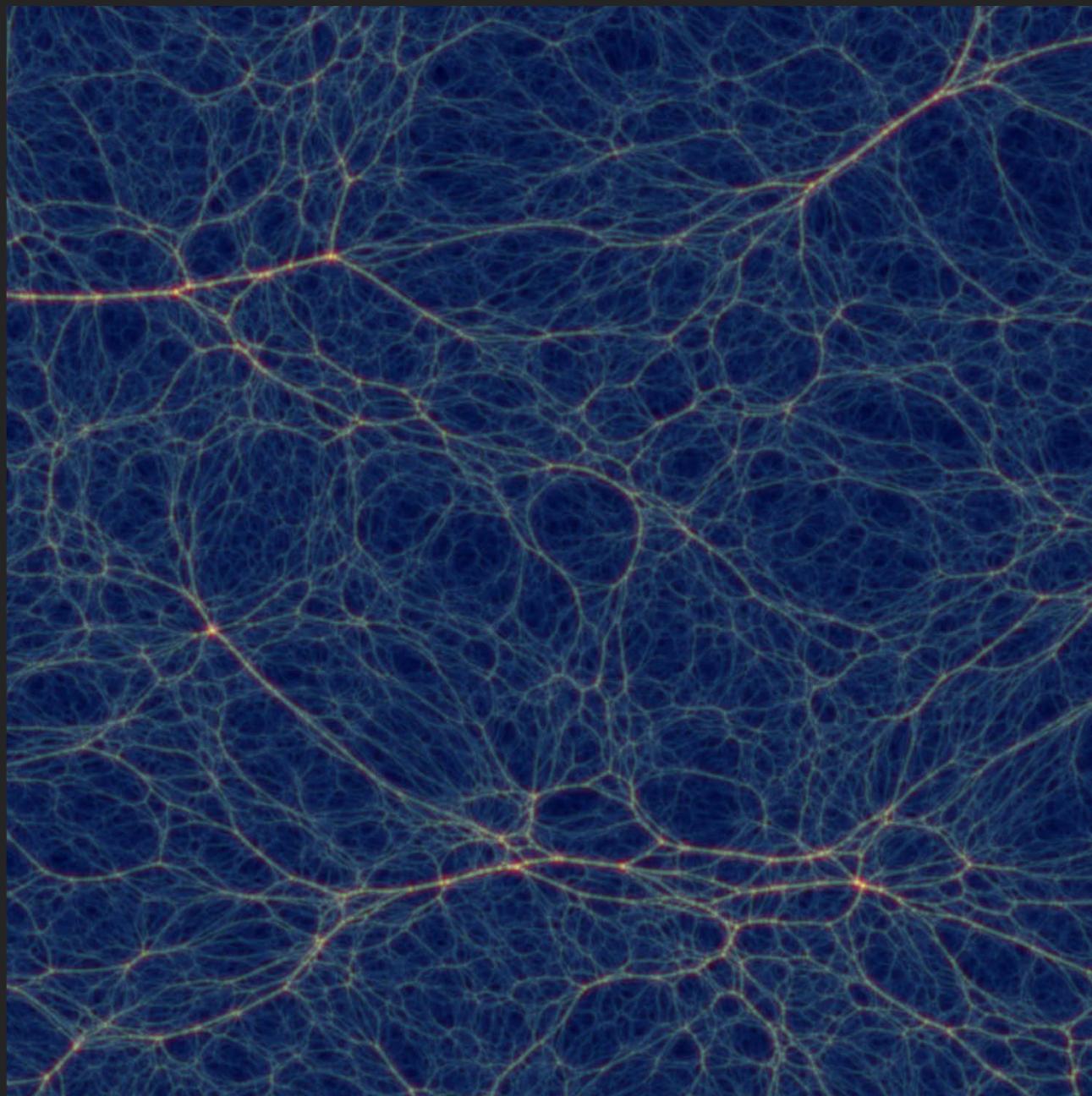
Johan Hidding
2012

Multiscale
Infrastructure
Voids:



manifestation

Hierarchical
Buildup of
Voids

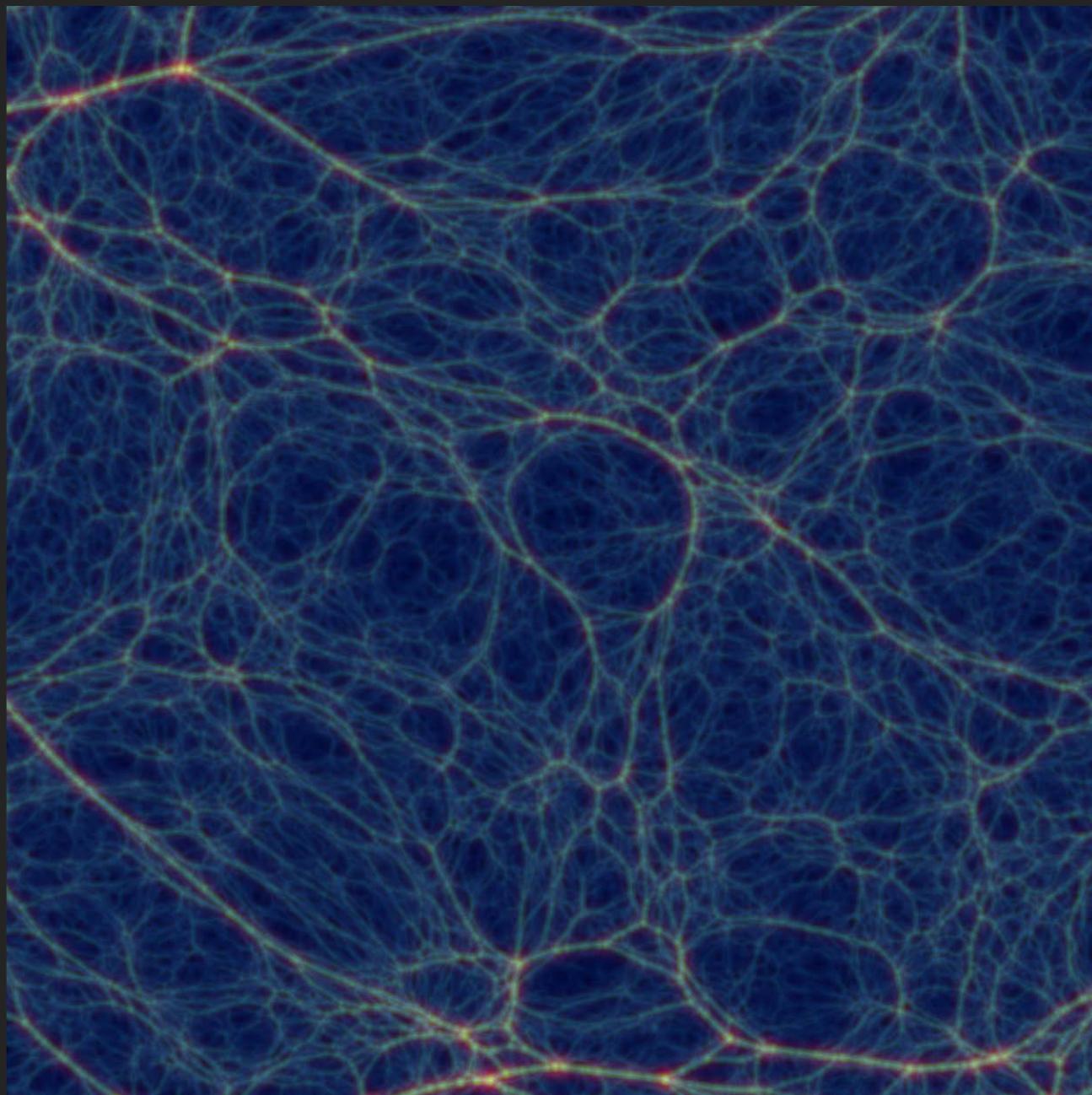


**Multiscale
Infrastructure
Voids:**



manifestation

**Hierarchical
Buildup of
Voids**

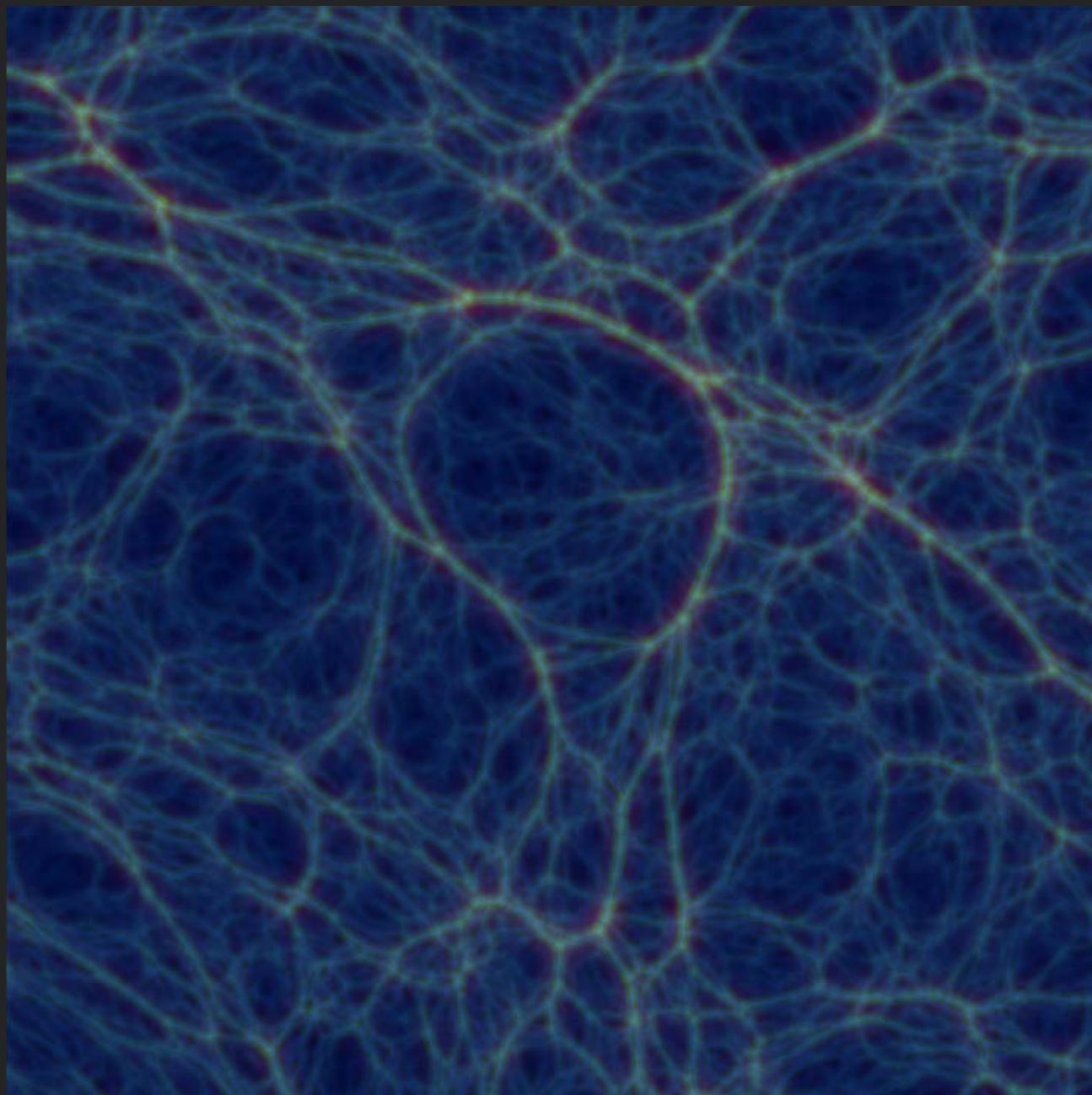


**Multiscale
Infrastructure
Voids:**



manifestation

**Hierarchical
Buildup of
Voids**

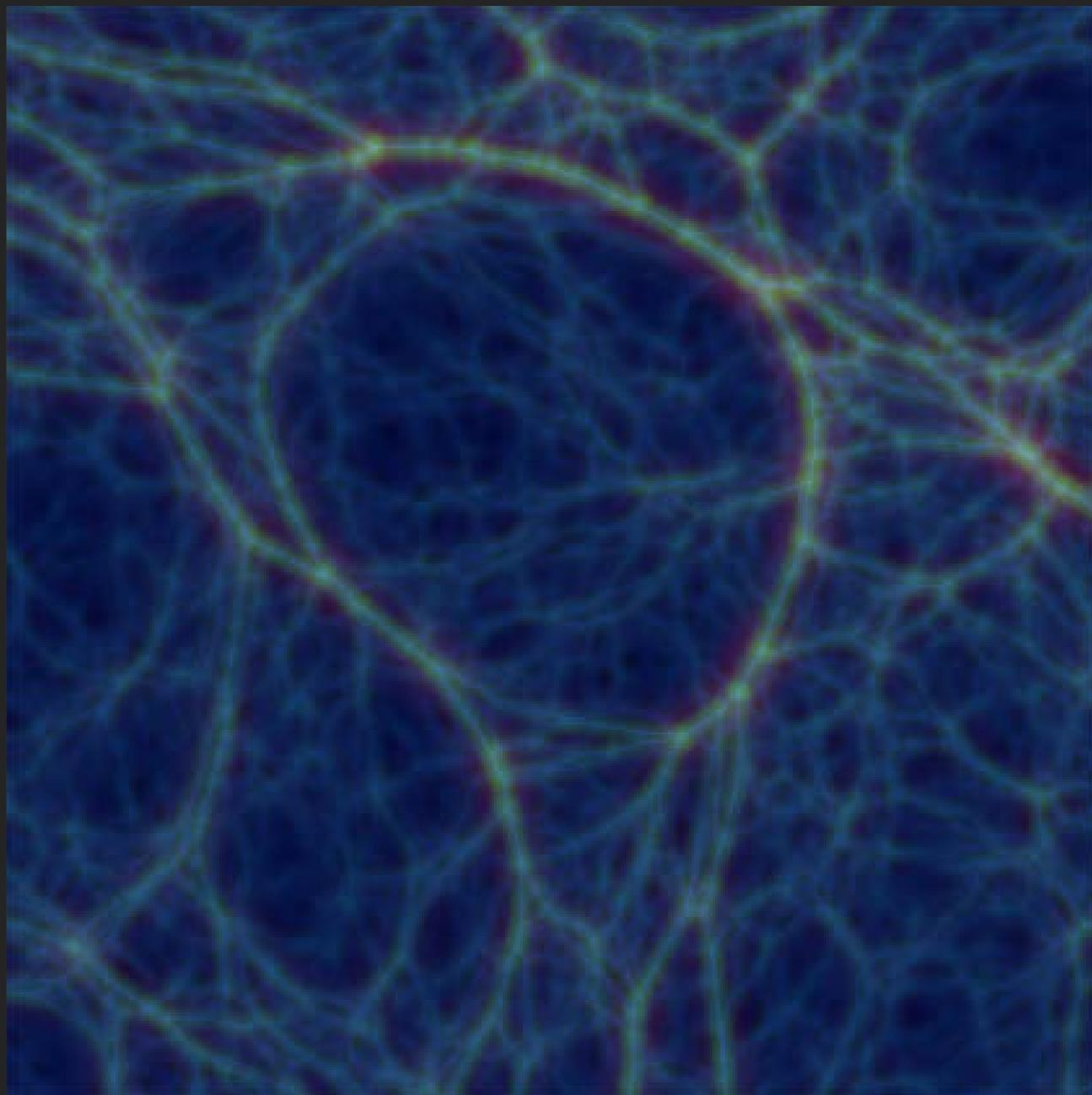


**Multiscale
Infrastructure
Voids:**

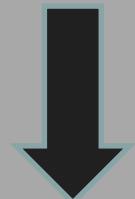


manifestation

**Hierarchical
Buildup of
Voids**

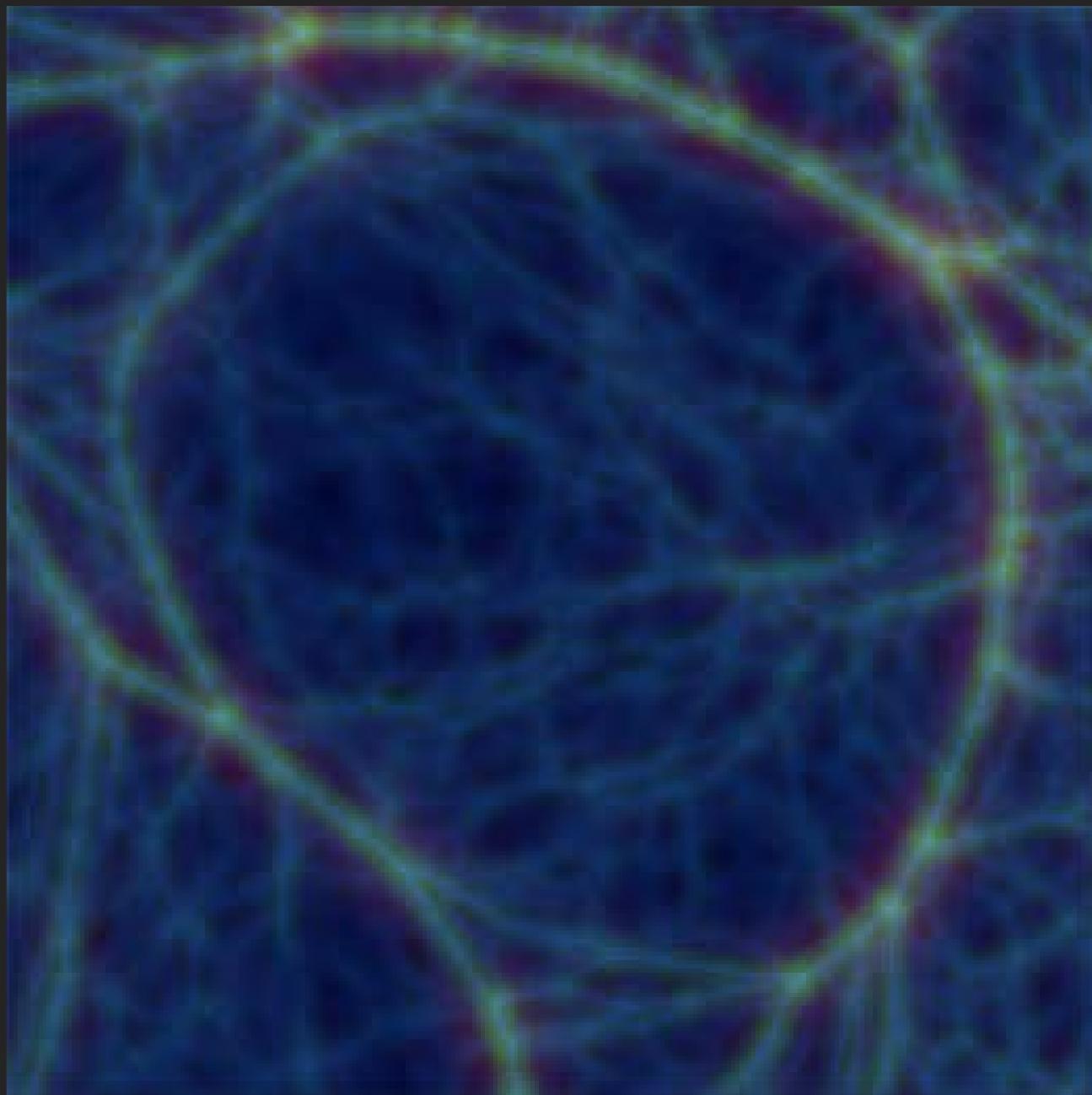


**Multiscale
Infrastructure
Voids:**

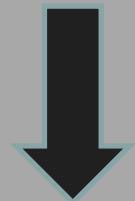


manifestation

**Hierarchical
Buildup of
Voids**

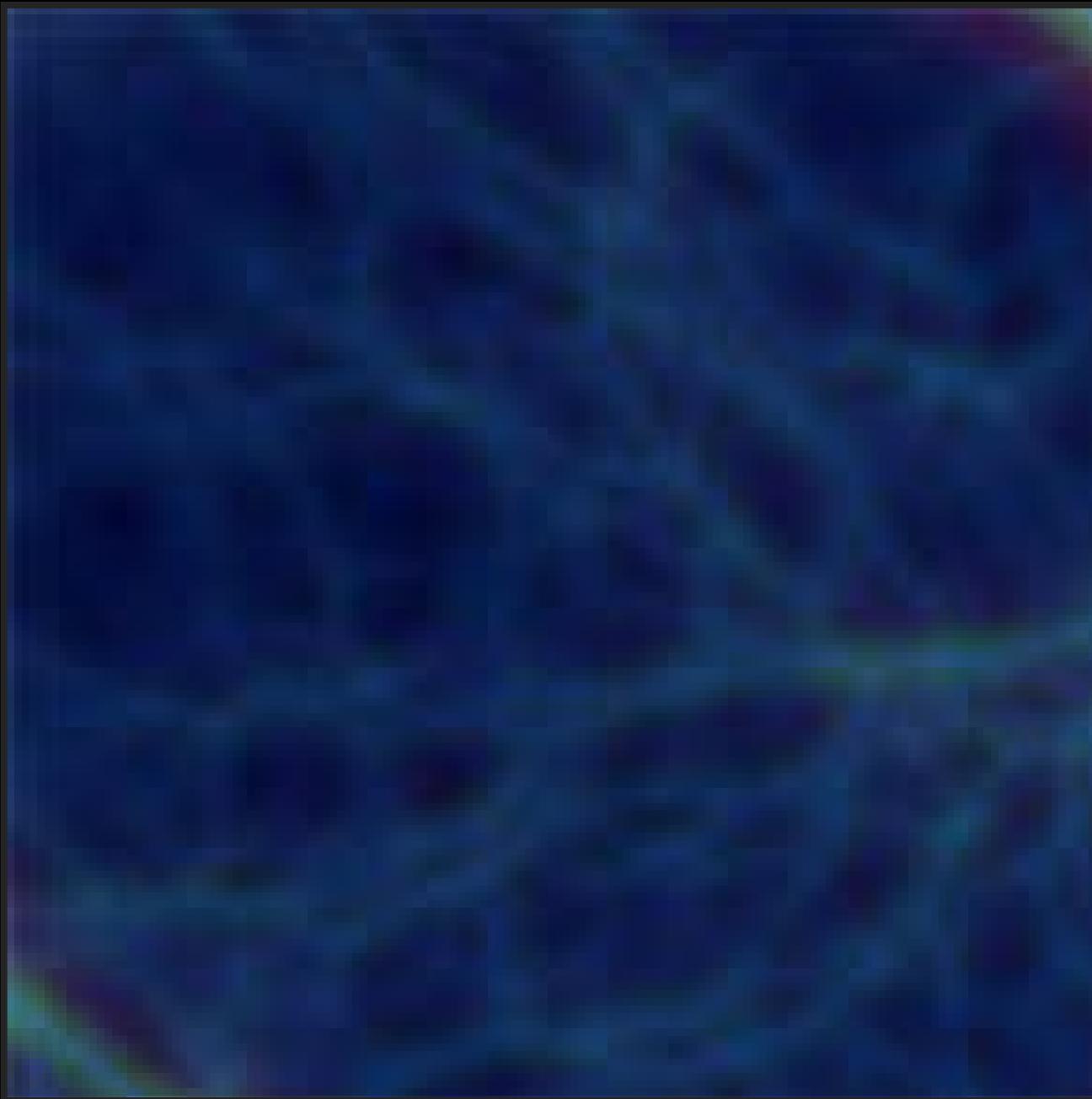


**Multiscale
Infrastructure
Voids:**



manifestation

**Hierarchical
Buildup of
Voids**

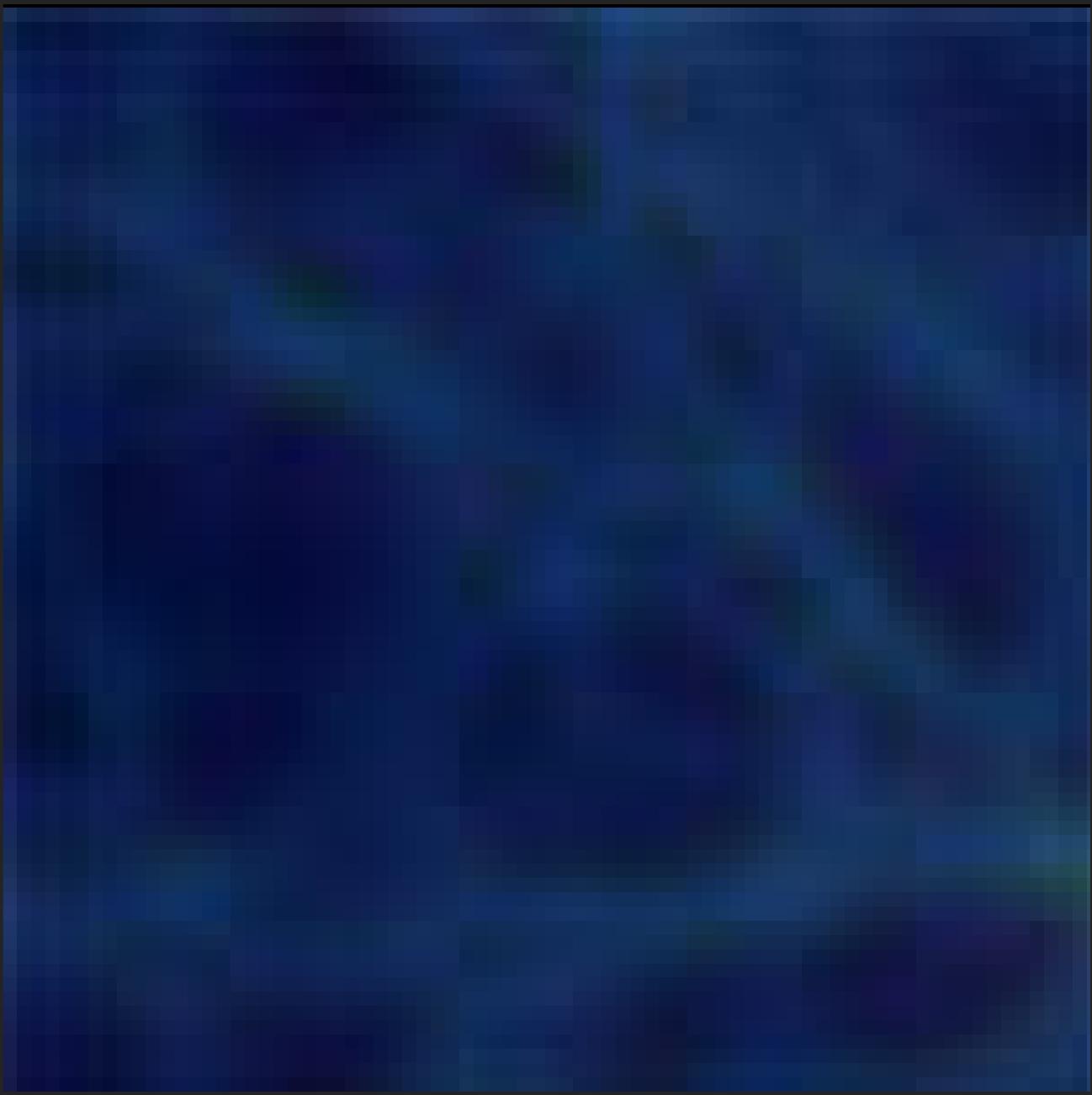


**Multiscale
Infrastructure
Voids:**



manifestation

**Hierarchical
Buildup of
Voids**



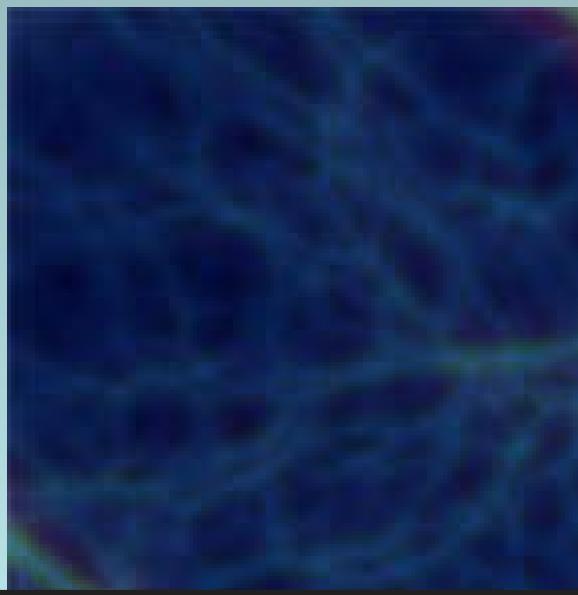
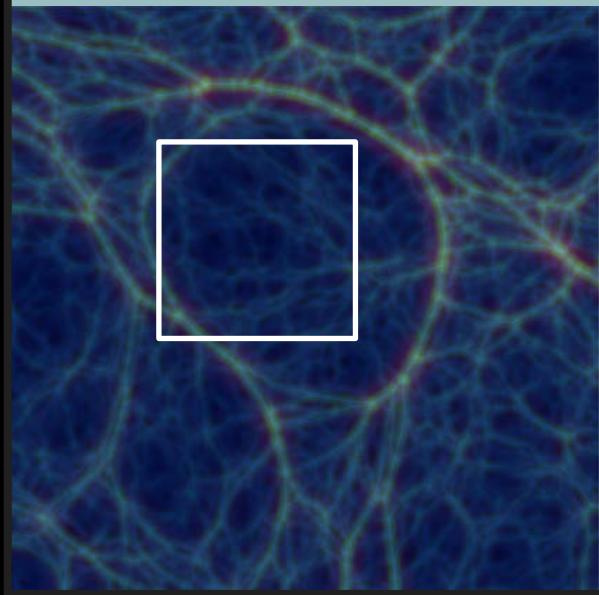
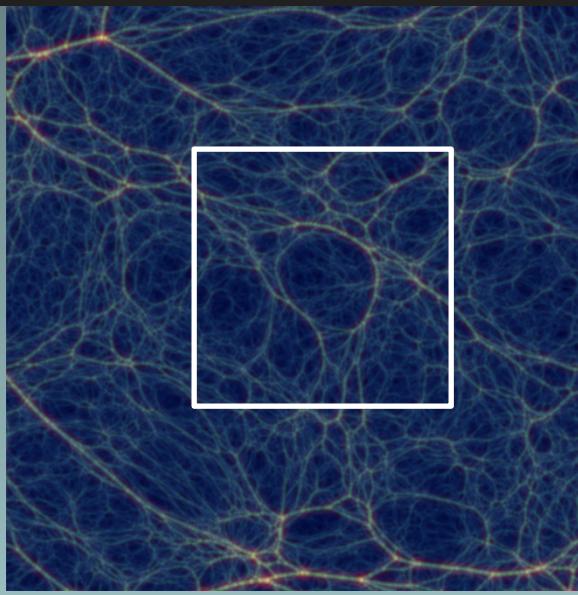
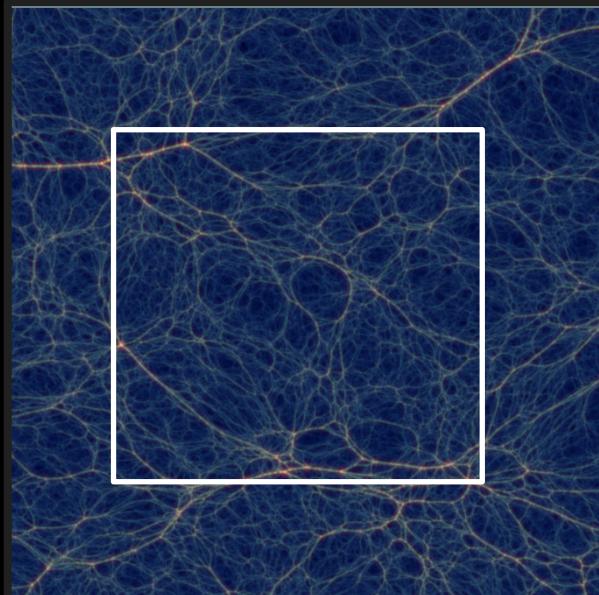
**Multiscale
Infrastructure
Voids:**



manifestation

**Hierarchical
Buildup of
Voids**

Void Substructure



Zoom in: 3 levels

Substructure on
all scales:

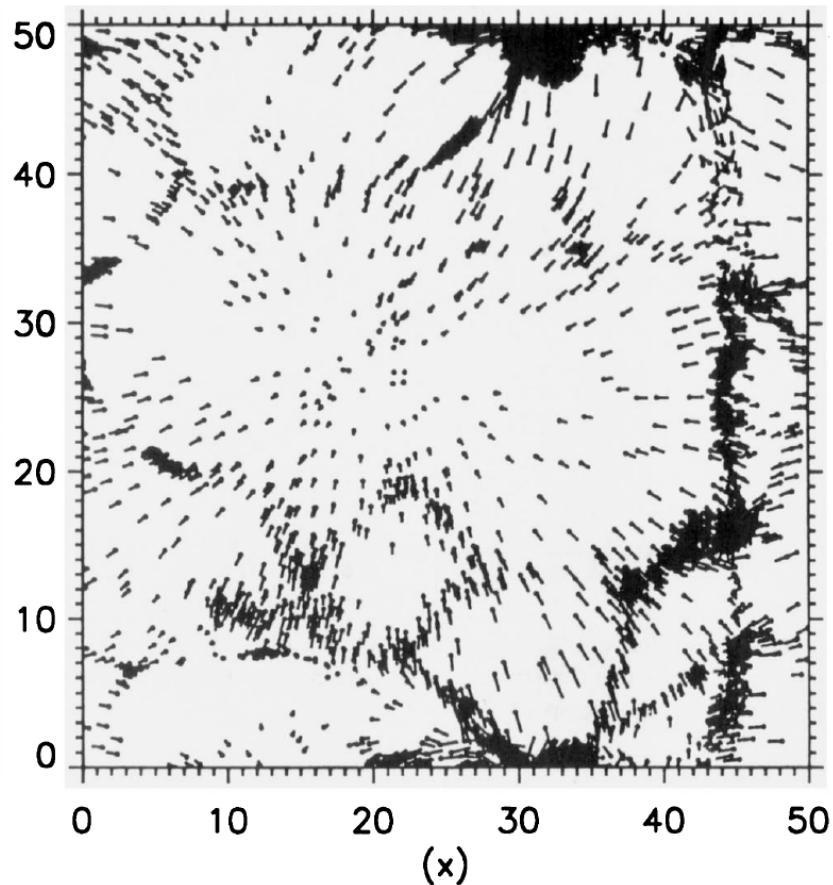
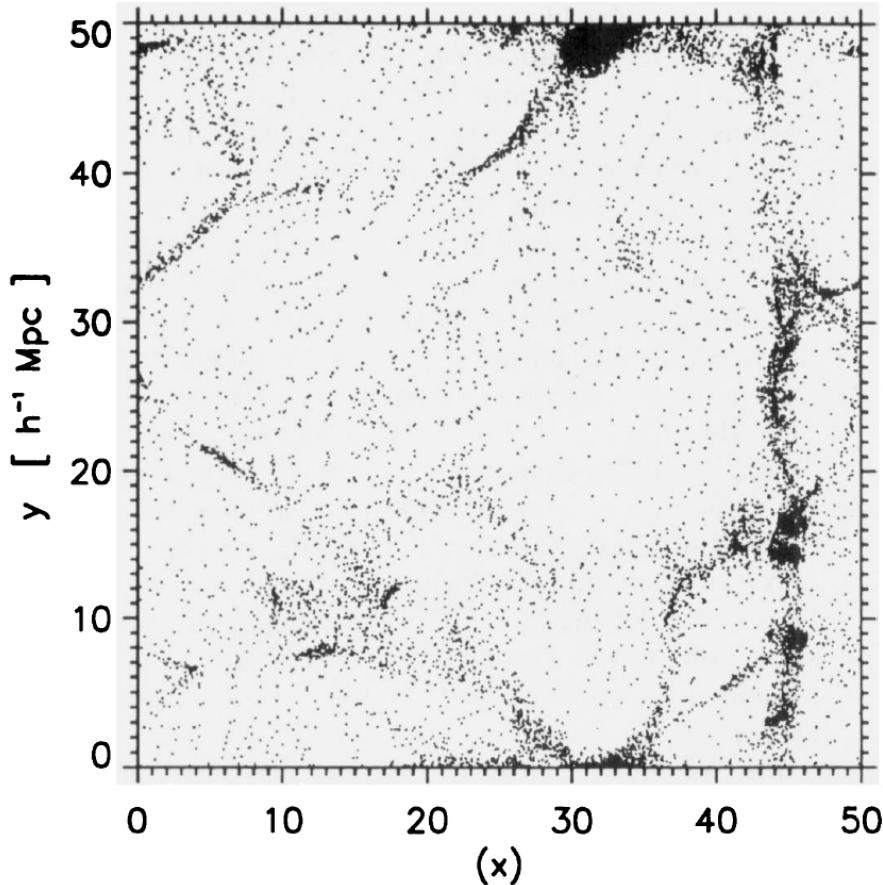
amplitude
diminishing towards
smaller scales

Void Hierarchy

Void Hierarchy

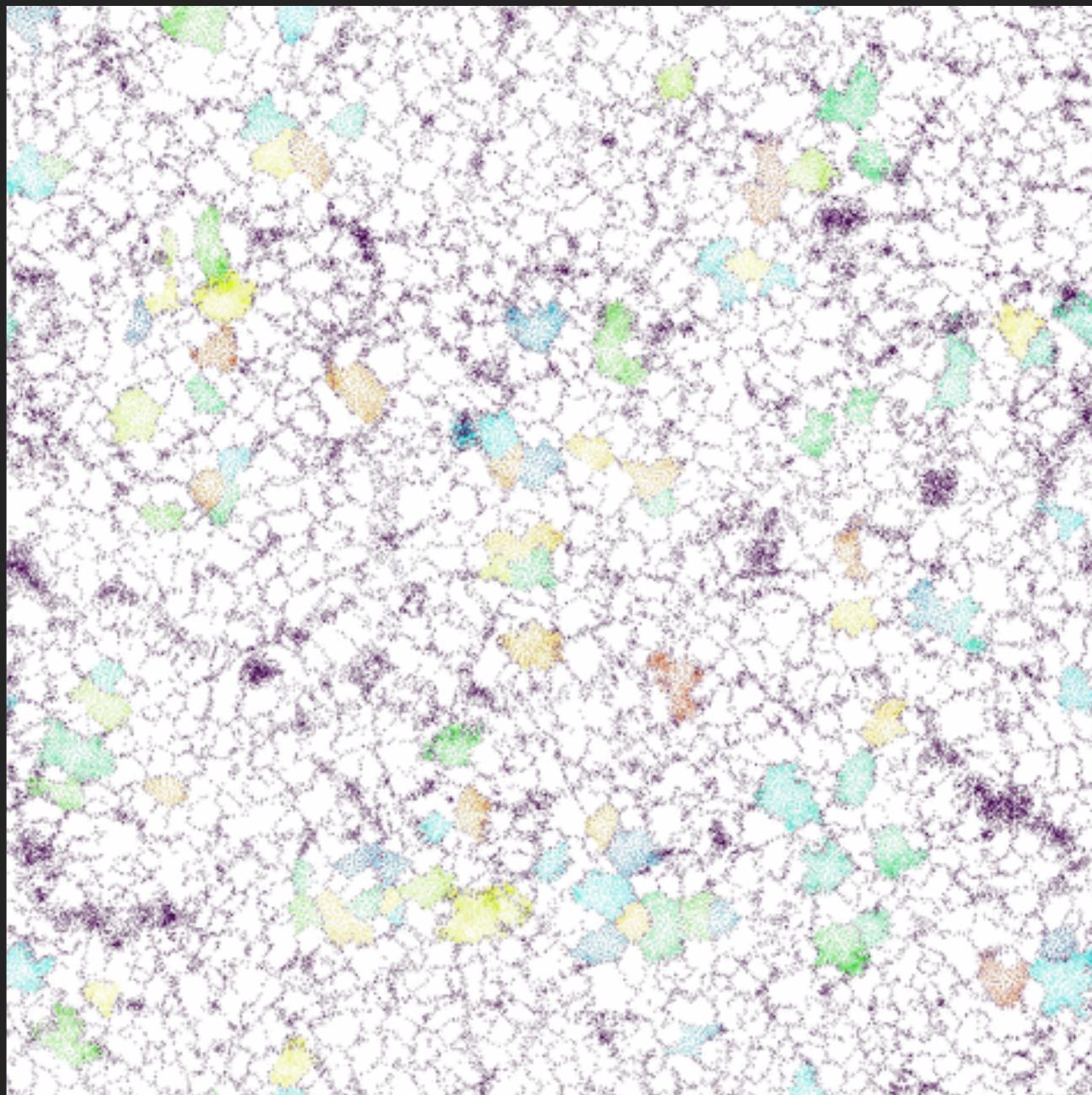
Void formation:
- via gradual merging of voids
- demolition of small voids

Dubinski et al. 1993
Van de Weygaert & van Kampen 1993



Constrained Void Simulations:
Van de Weygaert 1991
vdW & van Kampen 1993

CRF formalism:
Bertschinger 1987
vdW & Bertschinger 1996



Hierarchical Web Evolution:

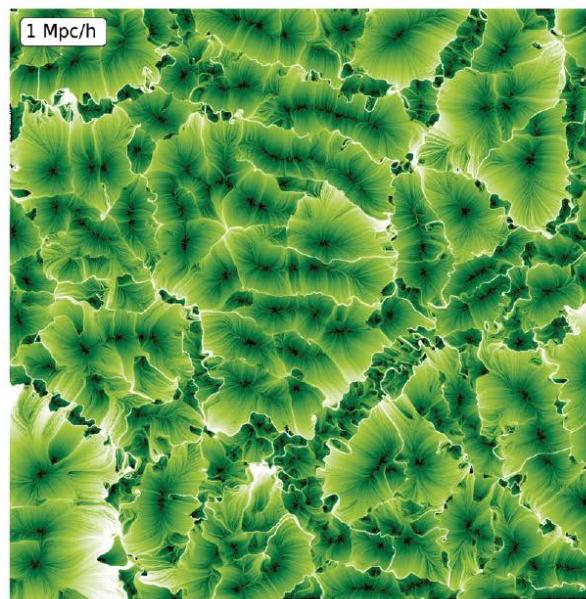
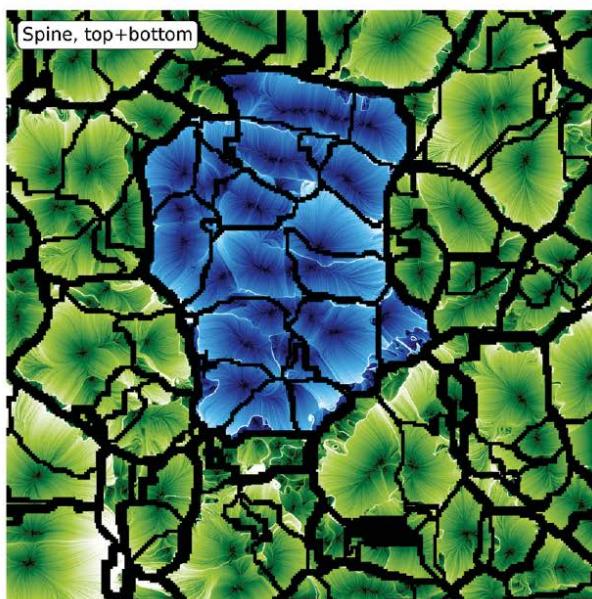
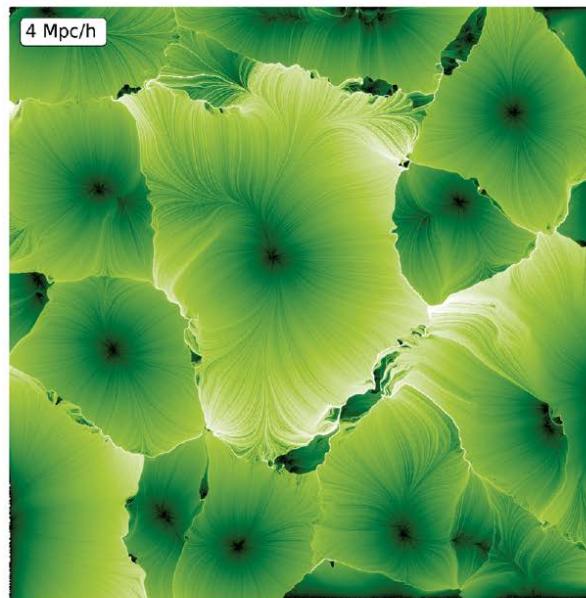
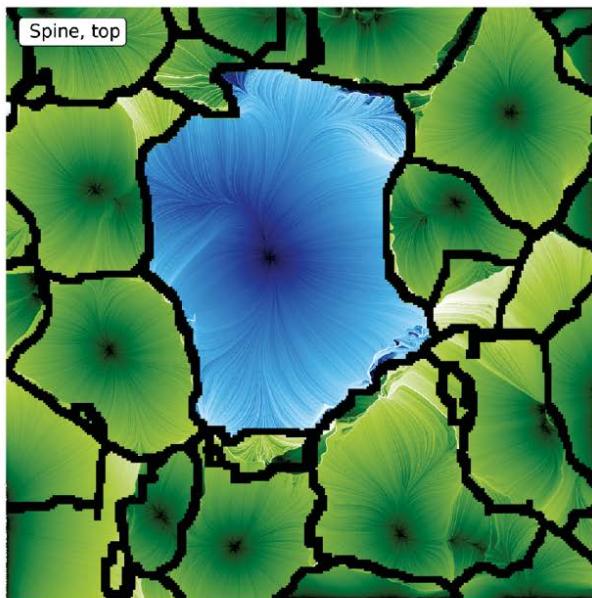
“Lagrangian” view:
development and fate
patterns LSS

Platen & vdW 2004

Hierarchical Web Evolution:

Void hierarchy
expressed in
multiscale structure
velocity outflow

Aragon-Calvo & Szalay 2012



Void Dichotomy

- Voids emerge out of primordial Gaussian density depressions
- Primordial Gaussian Density Field:
symmetry overdense \longleftrightarrow underdense
- Why is void population:
Not dominated by small voids ?
What happened to (PS) hierarchy ?

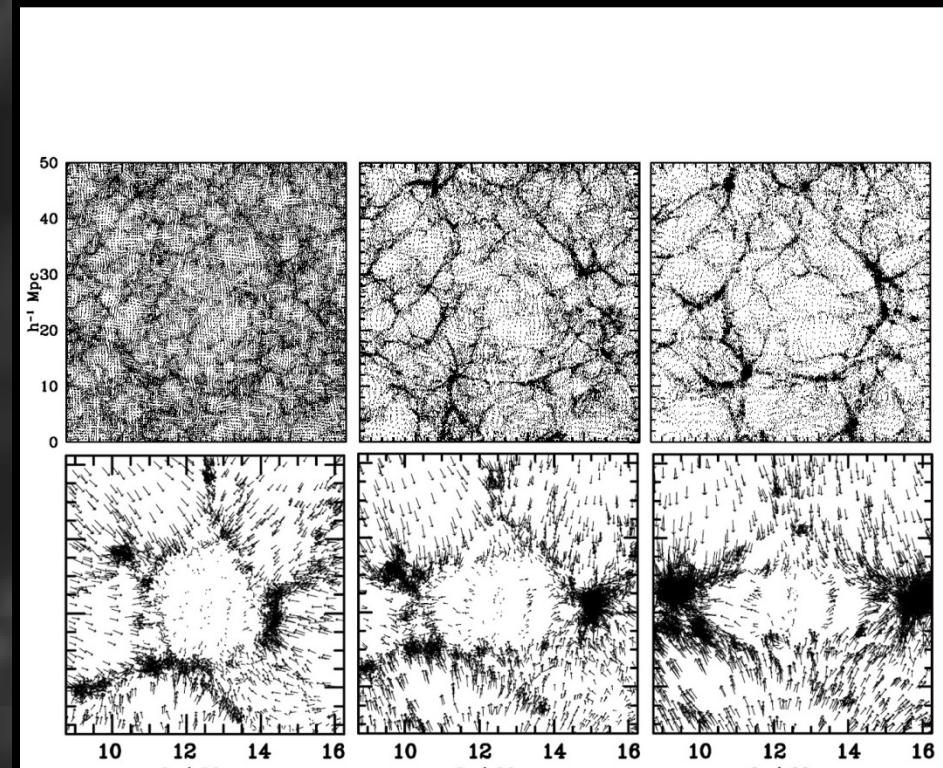
Void Evolution Processes

- **Void Merging**

as voids expand and meet their peers, they merge into ever larger voids ...

- **Void Collapse**

when embedded within an overdense or tidally sheared region (filaments ...), weak voids get squeezed out of existence...

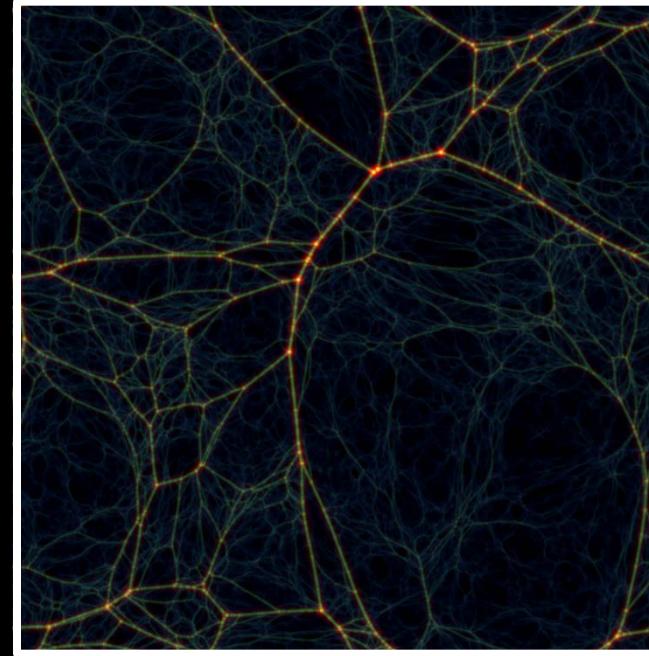
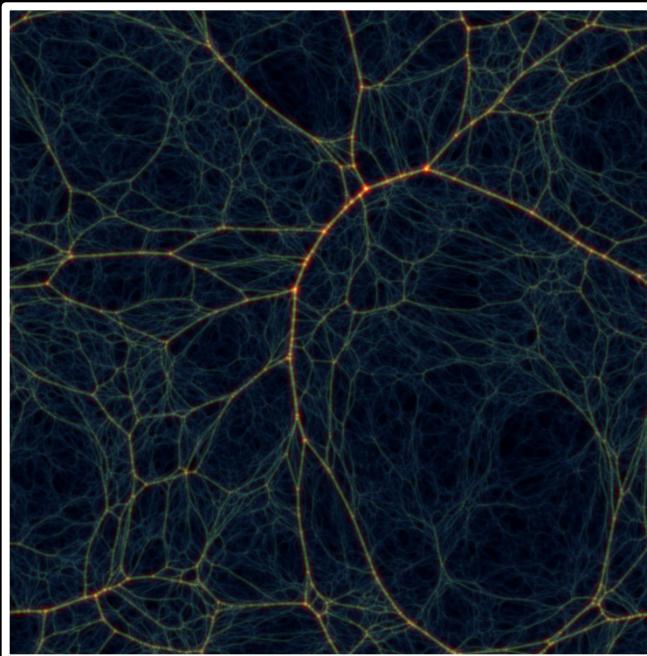
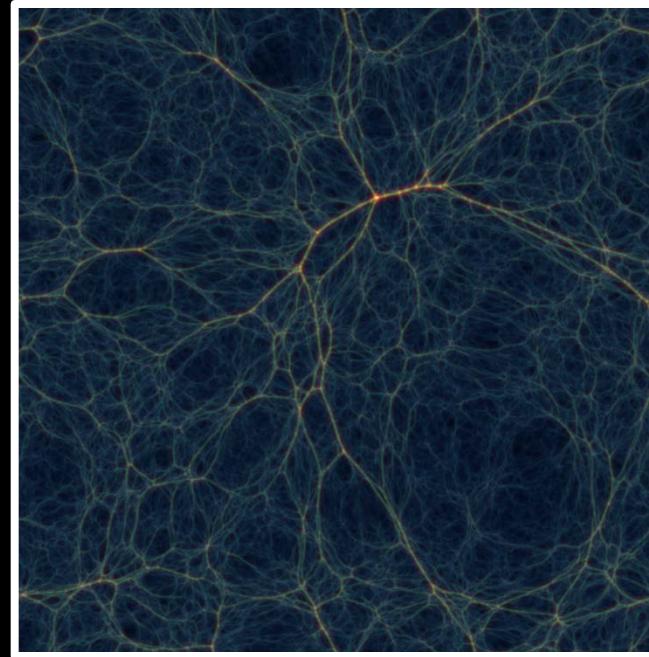
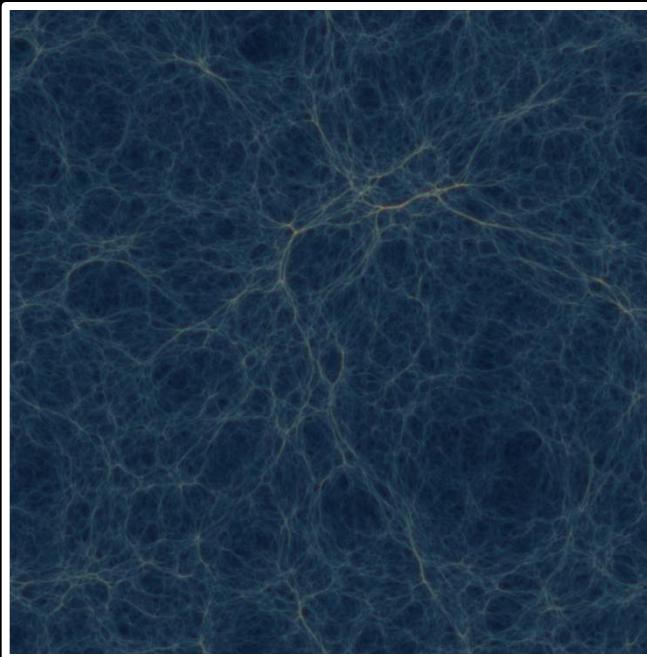


Hierarchical Web Evolution:

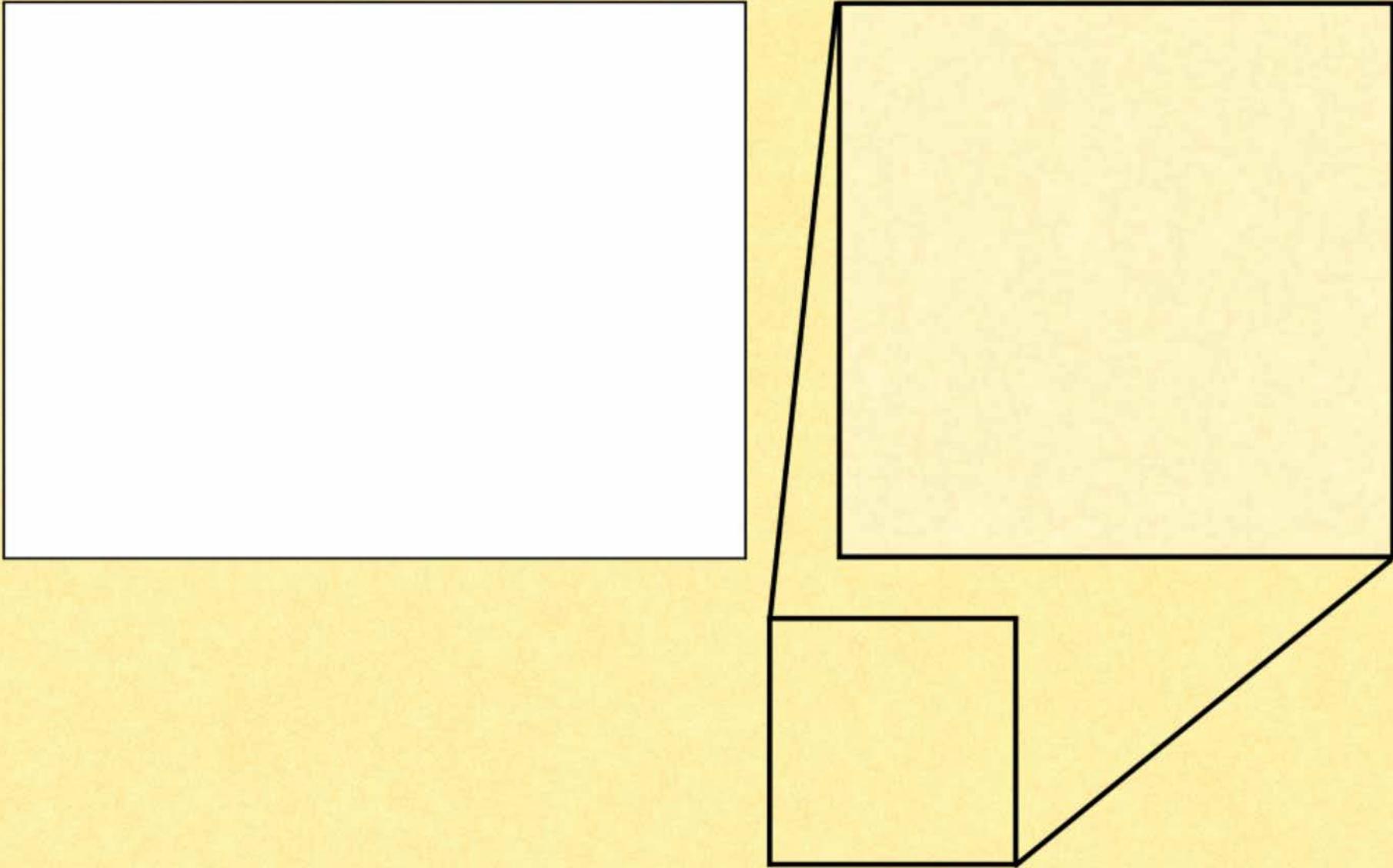
Adhesion simulation
buildup Cosmic Web

Johan Hidding
2012

time
evolution



Johan Hidding
2012



Hierarchical Cosmic Structure Formation

Two-Barrier Excursion Set Formalism

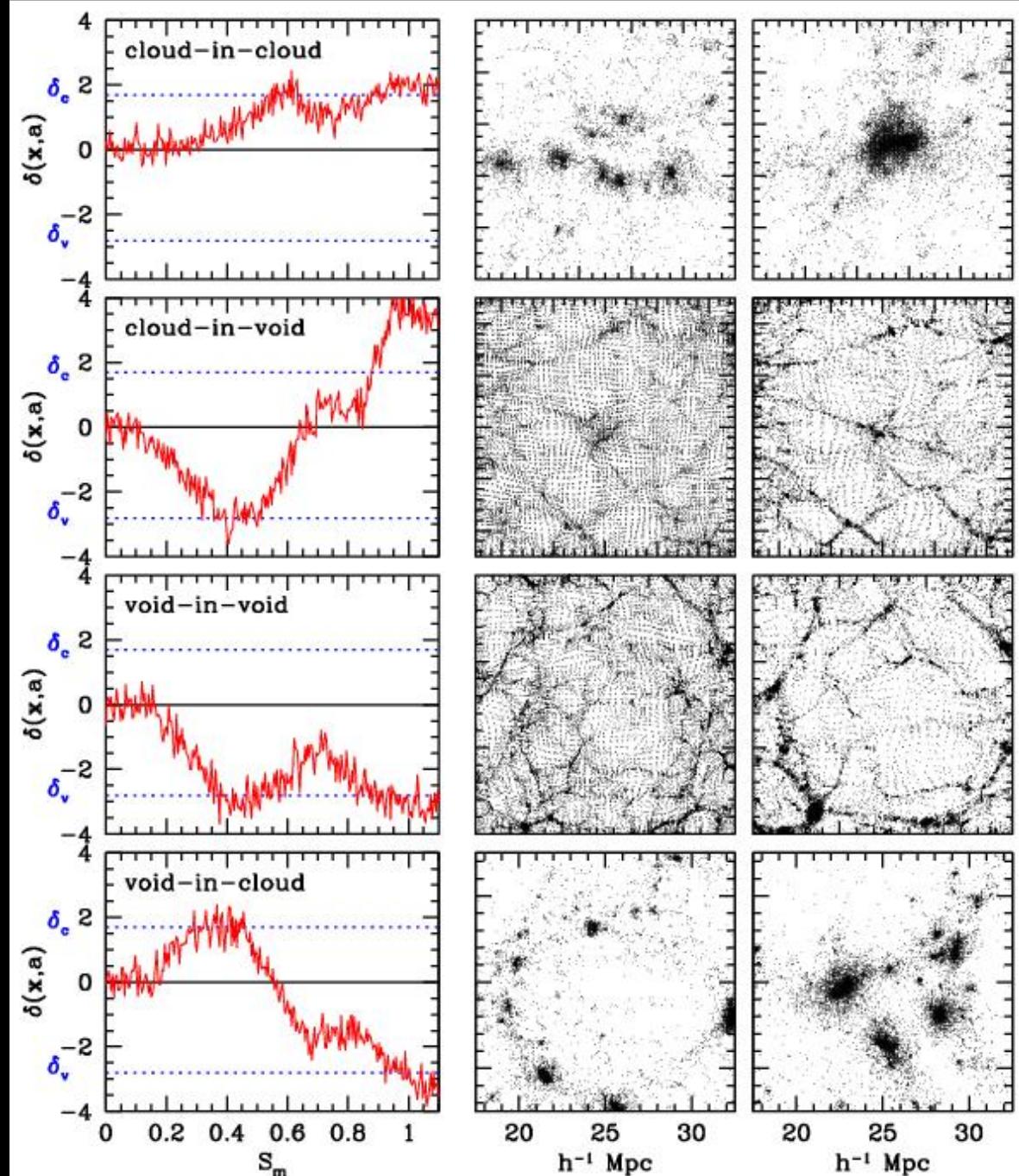
Sheth & vdW 2004

Press & Schechter 1974

Bond et al. 1991

Excursion Modes:

- Cloud-in-Cloud
- Cloud-in-Void
- Void-in-Void
- Void-in-Cloud



Extended Press-Schechter

Barrier Excursions

- Spherical linear collapse overdensity:

$$\Delta_{lin}(r, S_m, t) > \delta_c$$

- Collapse time:

$$a_{coll}(r) = \delta_c / \Delta_{lin}(r, S_m)$$

- Initial density field:

prediction object formation time:

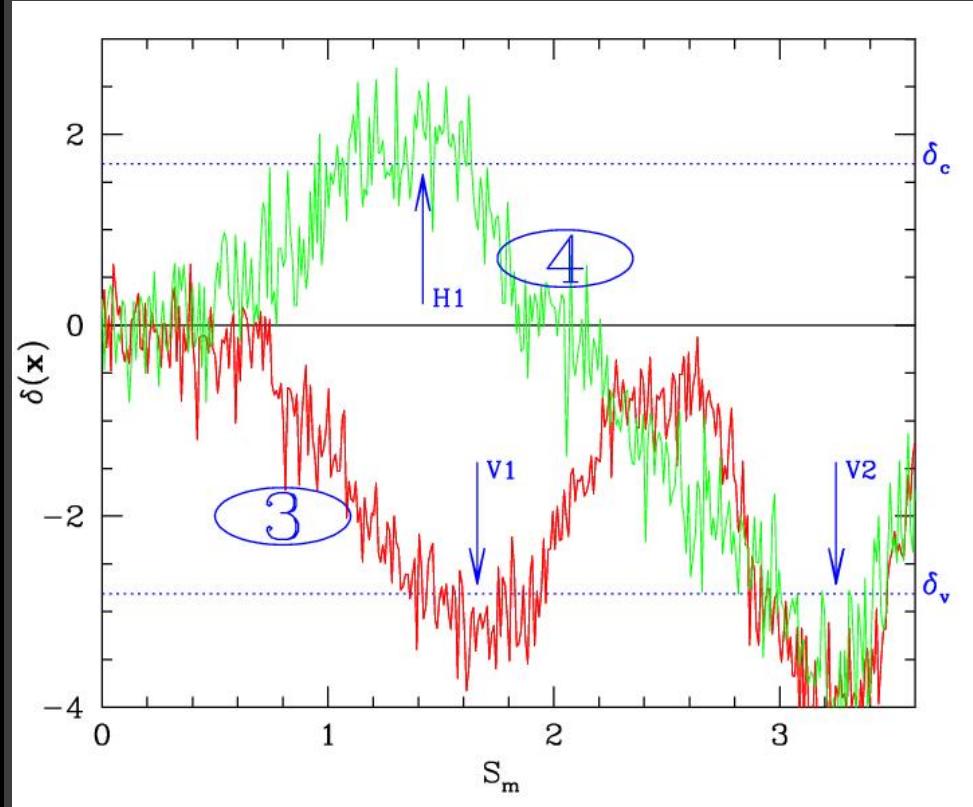
$$a_{coll}(r) \Leftrightarrow \Delta_{lin,0}(r, S_m)$$

dependent on:

Collapse Barrier δ_c

- Critical density value:

$$\text{EdS , } \Omega_0=1: \quad \delta_c \sim 1.69$$



cumulative random walk:

$$\delta_s(\vec{x}; \lambda_m) = \int_{|k| < k_m} \frac{d\vec{k}}{(2\pi)^3} \hat{\delta}(\vec{k}) e^{-i\vec{k} \cdot \vec{x}}$$

Void Volume Distribution

(Sheth & vdW 2004)

- Small Void tail suppressed
- Peaked Void Size Distribution
Characteristic Void Size
- Self-Similar Evolution:
increasing
characteristic Void Size
- Volume-filling:
at any cosmic epoch,
for power-law $P(k)$,
approximately void-filling
- Excess Void Expansion:
“Super-Hubble” expansion

$$\nu_v(M) = \frac{|\delta_v|}{\sigma(M)}$$

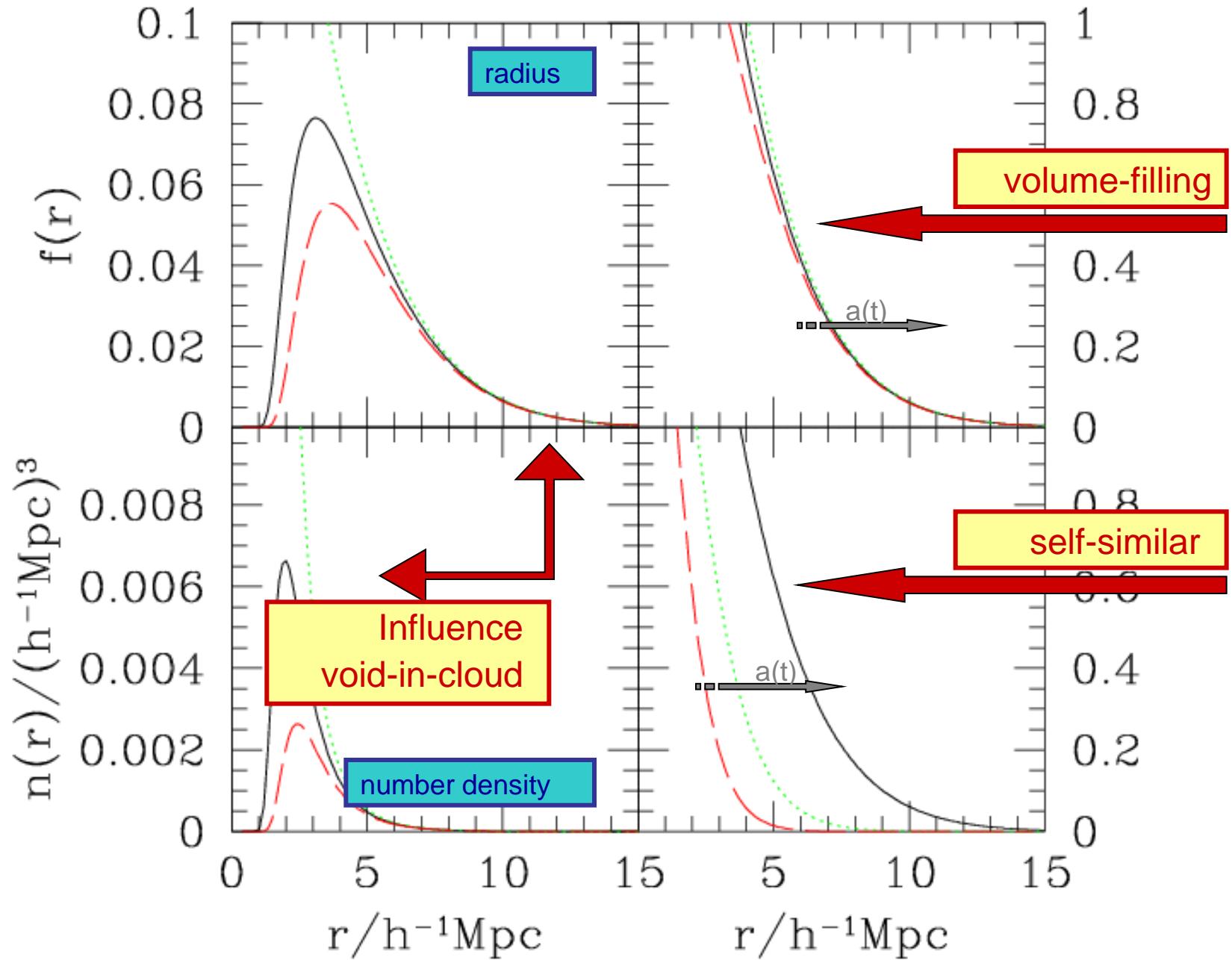
$$\mathcal{D} \equiv \frac{|\delta_v|}{(\delta_c + |\delta_v|)}$$

void mass distribution function

$$n_v(M) dM = \sqrt{\frac{2}{\pi}} \frac{\rho_u}{M^2} \nu_v(M) \exp\left(-\frac{\nu_v(M)^2}{2}\right) \left| \frac{d \ln \sigma(M)}{d \ln M} \right| \exp\left\{-\frac{|\delta_v|}{\delta_c} \frac{\mathcal{D}^2}{4\nu_v^2} - 2 \frac{\mathcal{D}^4}{\nu_v^4}\right\}$$

for power-law power spectrum:

$$n_v(M) dM \approx \sqrt{\frac{1}{2\pi}} \left(1 + \frac{n}{3}\right) \frac{\rho_u}{M^2} \left(\frac{M}{M_{v,*}}\right)^{(n+3)/6} \exp\left(-\left(\frac{M}{M_{v,*}}\right)^{(n+3)/3}\right) \exp\left\{-\frac{\mathcal{D}^2}{2} \left(\frac{|\delta_v|}{4\delta_c} + \mathcal{D}^2 \left(\frac{M}{M_{v,*}}\right)^{-(n+3)/3}\right) \left(\frac{M}{M_{v,*}}\right)^{-(n+3)/3}\right\}$$



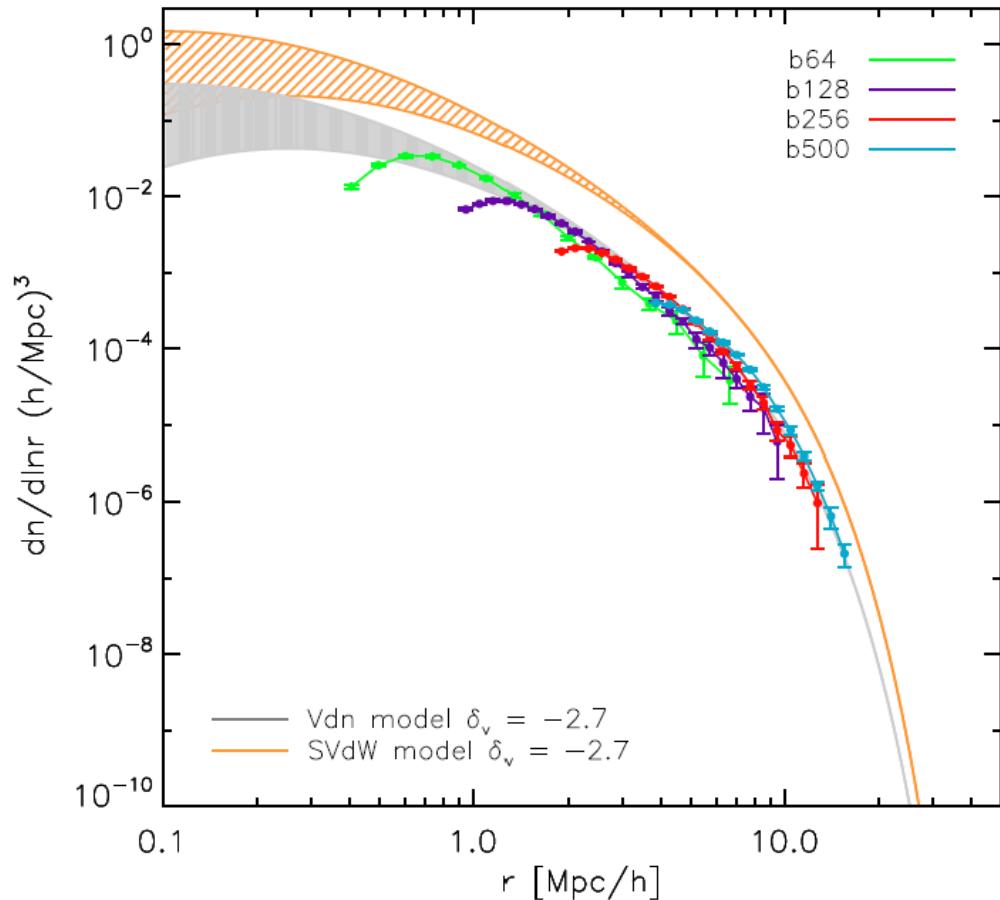
Void Population & Excursions

Jennings et al. 2013

Two-barrier SvdW
void excursion set formalism:

correcting simple assumption
 δ_v criterion:

Jennings et al. 2013:
void volume occupation constraints



Voids & Cosmology

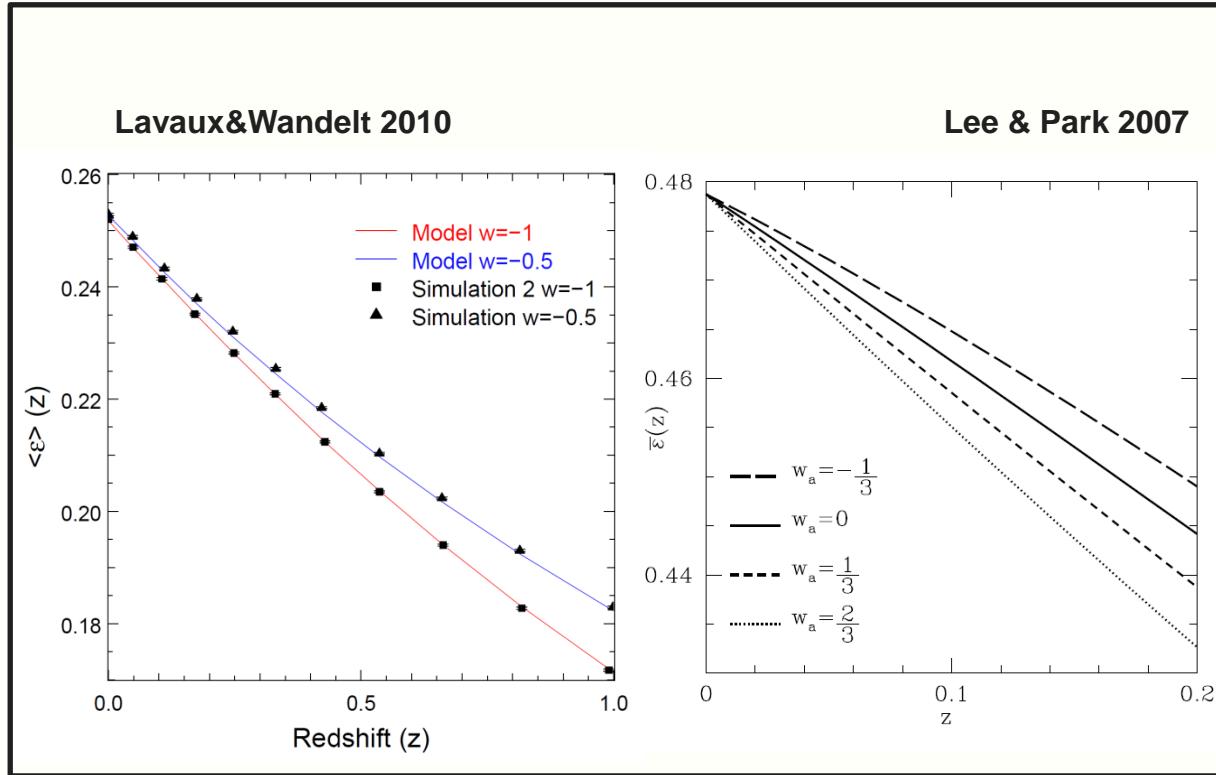
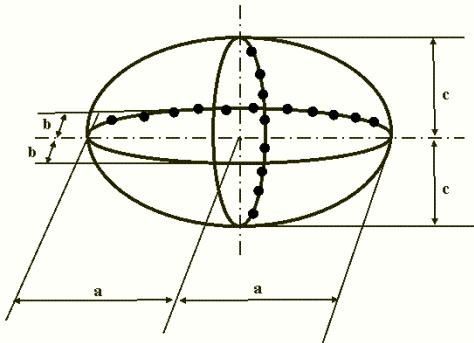
Voids & Cosmology

- **Cosmological probe:**

Voids contain significant amount of information on global cosmological parameters:

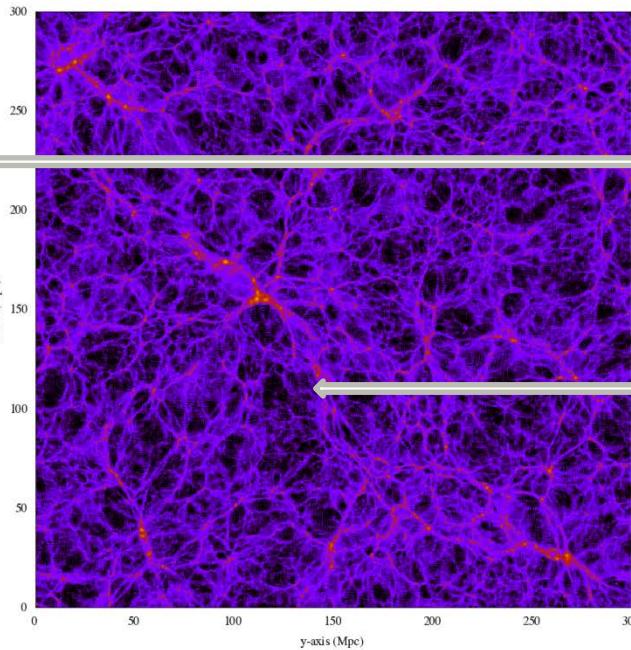
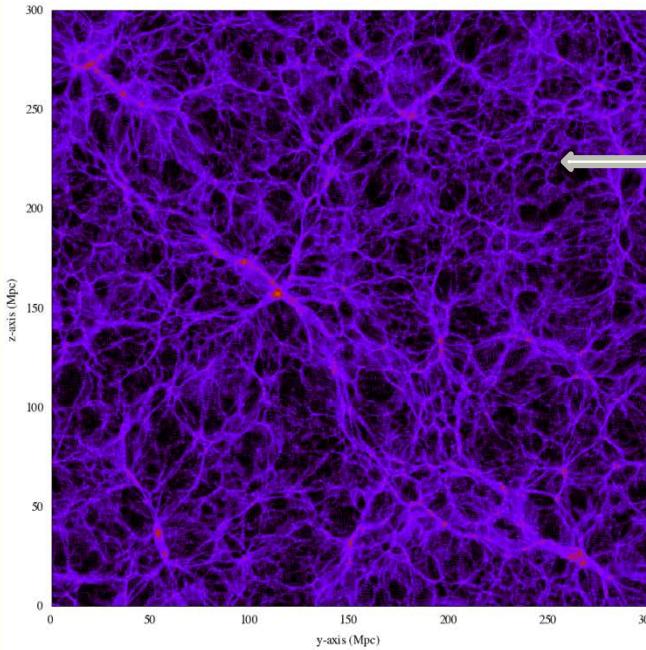
- void substructure: test GR/modified gravity
- void outflow: dark matter
- void shapes: dark energy
- supervoids: existence; CMB - ISW

Voids & Dark Energy



Evolving Void Population:

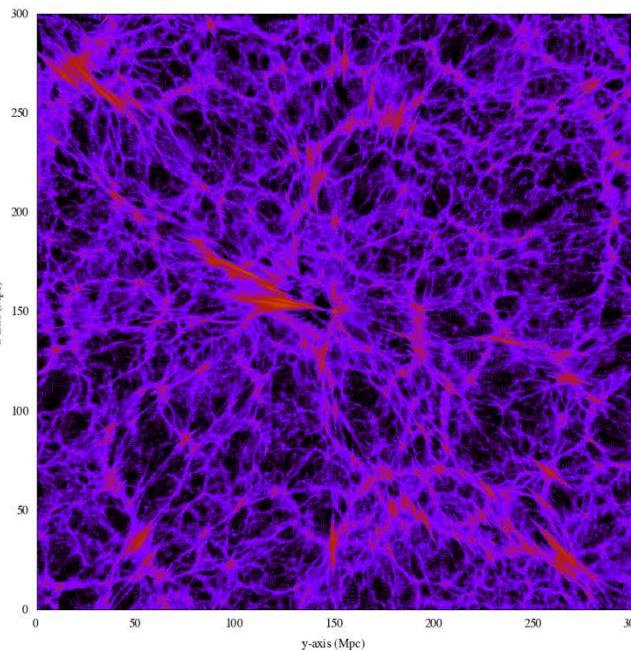
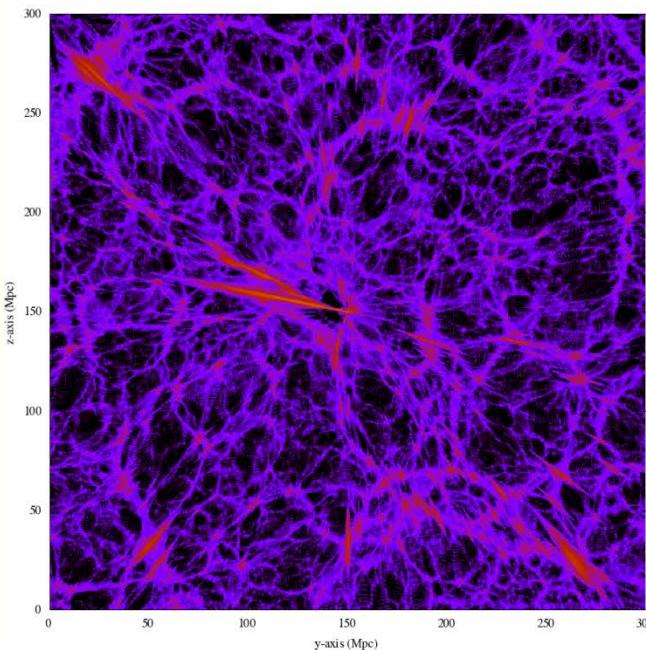
- Void shape evolution highly sensitive to:
dark energy equation of state
- Manifestation of high sensitivity void dynamics to external tidal forces



Comparison

WMAP
concordance
cosmology

SUGRA cosmology

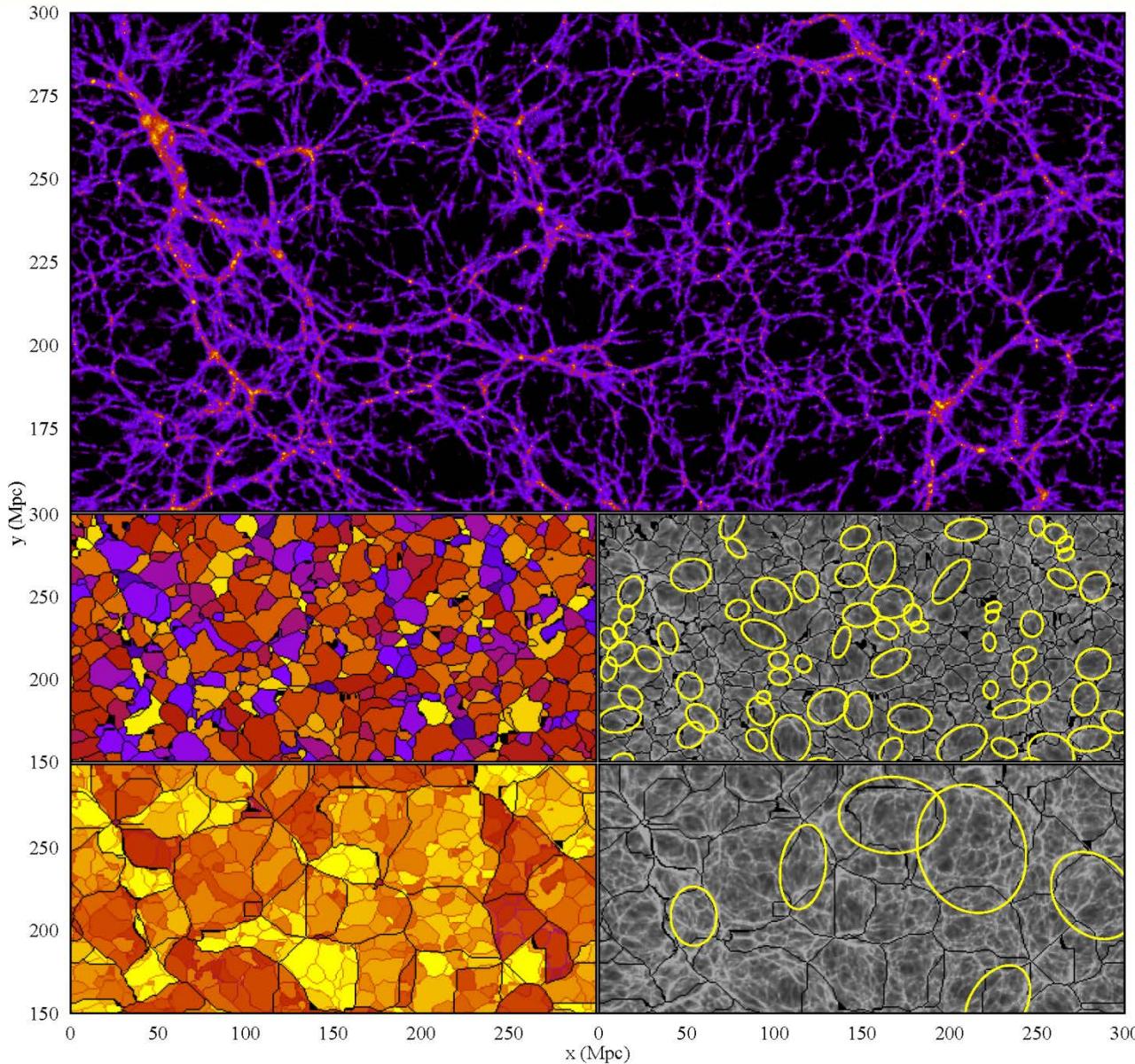


DM density $a=1$

Redshift Space

Patrick Bos 2011

Voids & Dark Energy

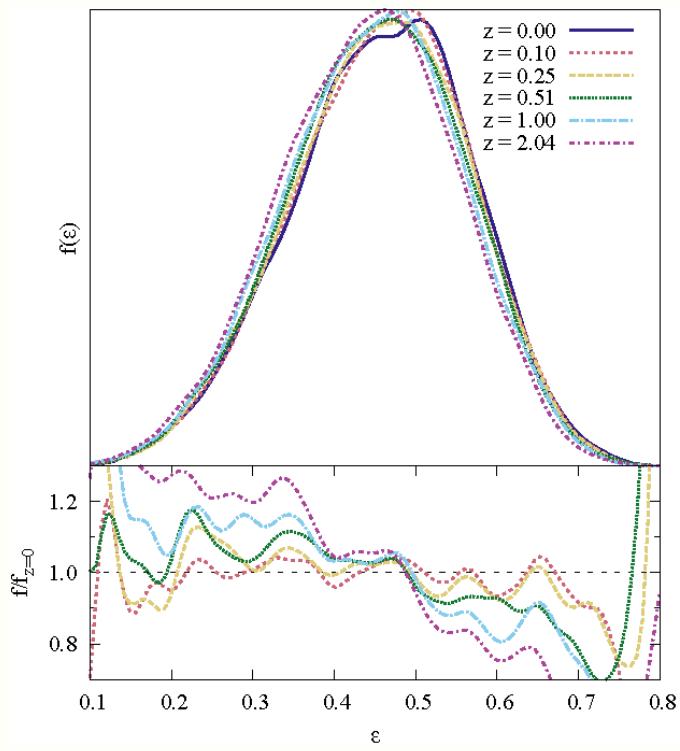


**Evolving
Void Population:**

Void Shape evolution:
sensitive probe Dark Energy

Bos, vdW, Dolag & Pettorino 2012

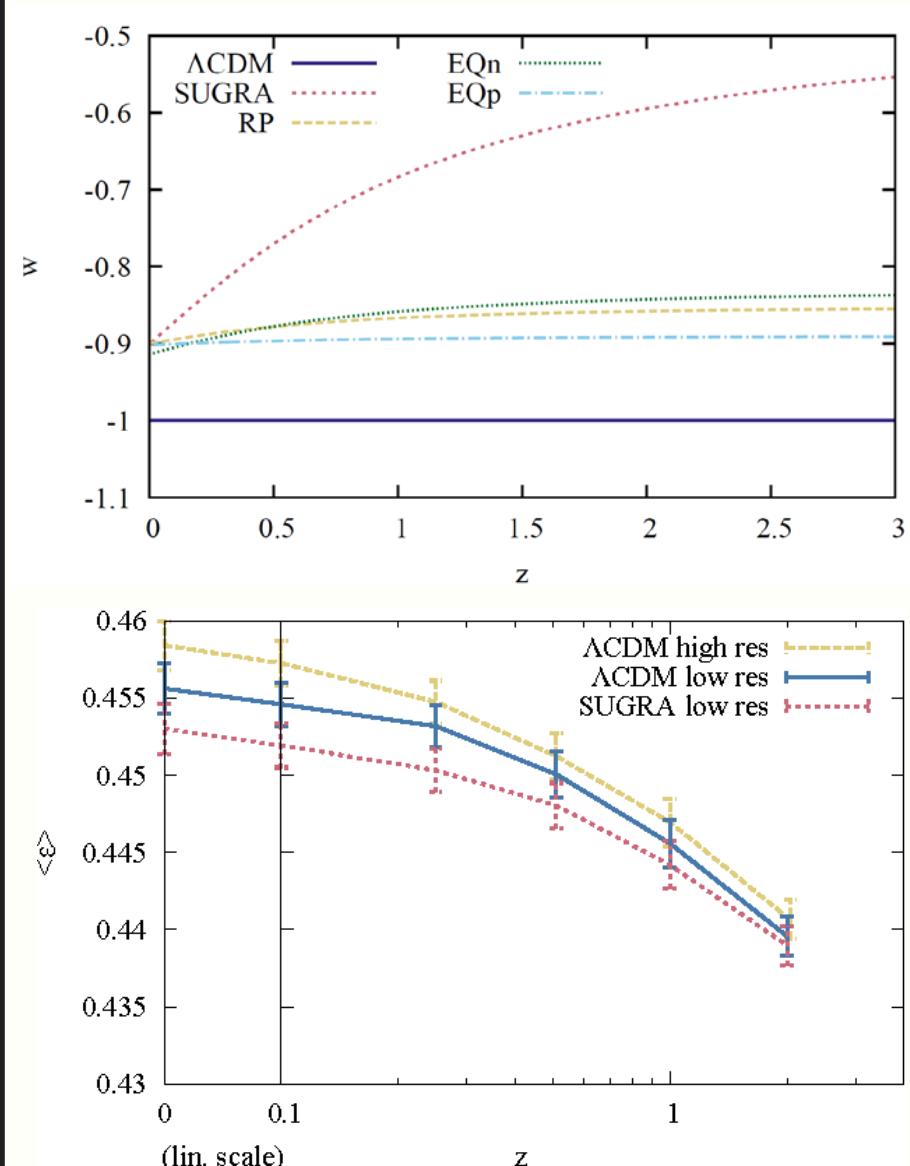
Voids and Dark Energy



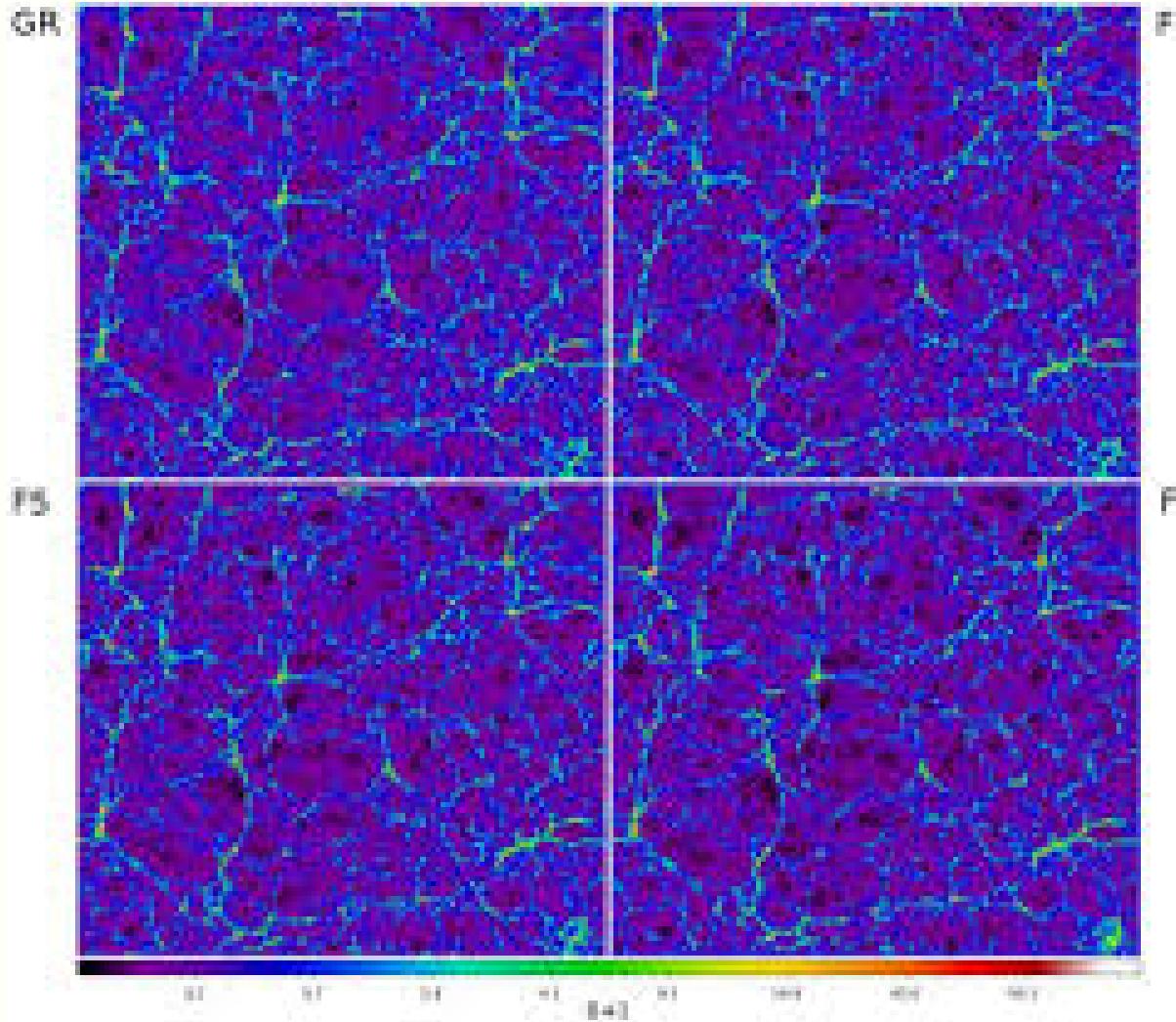
Void shape evolution LCDM:

Bos, vdW, Dolag & Pettorino 2011:

- Test of void shape evolution in Nbody simulations of 5 different DE cosmologies.
- Confirms that voids get less spherical as they evolve in time ! They become more elongated !
- Distinctly different void shape evolution of DM distribution in different DE models.



Modified Gravity & Voids



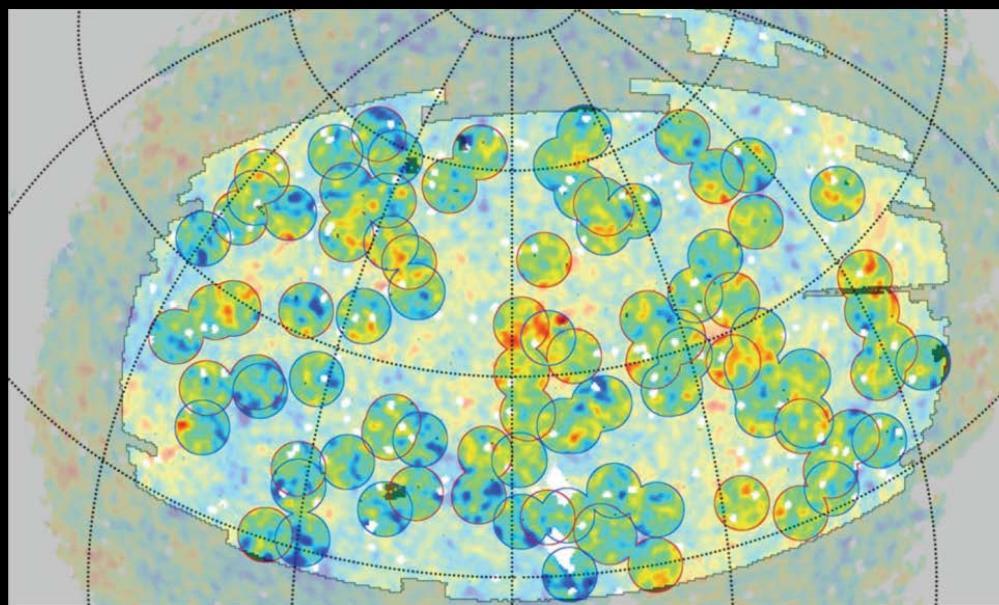
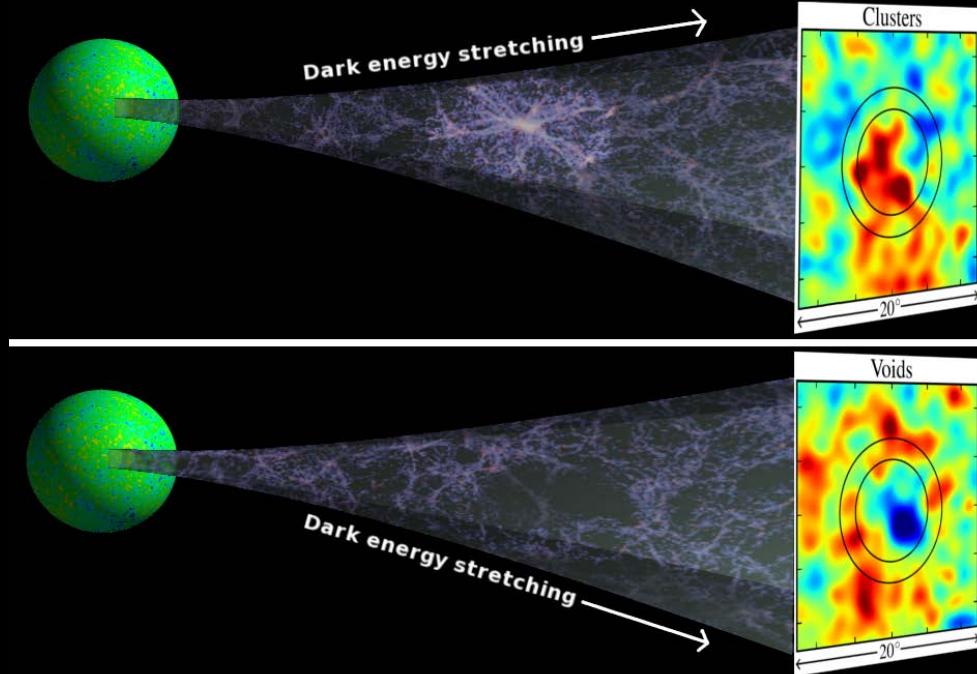
Void Population

**excellent probe for
modified gravity theories:**

e.g. Li et al. 2010, ...

Hellwing 2014

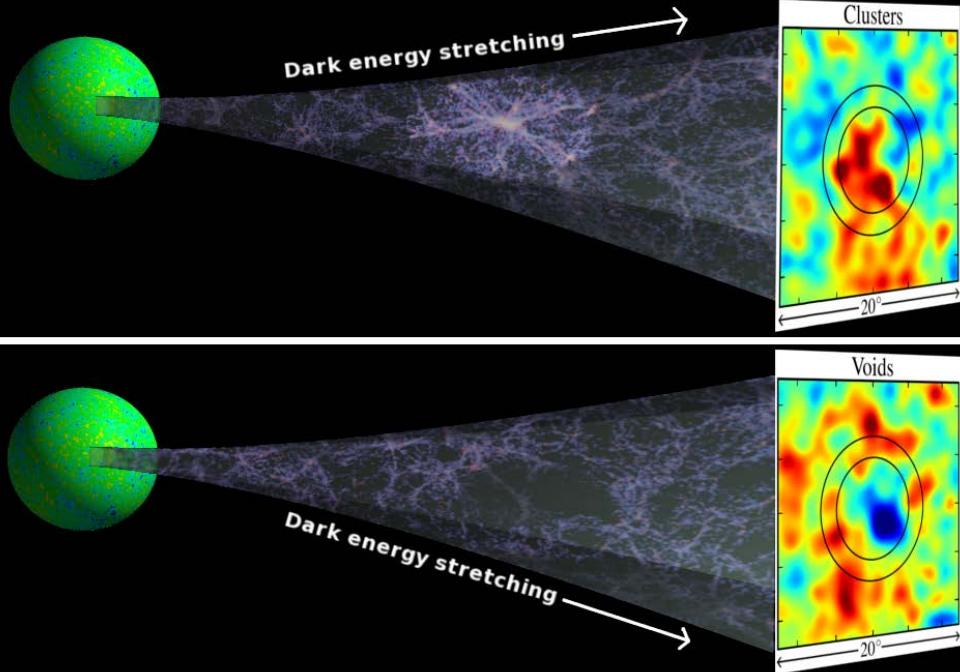
(Super)void & ISW



Granett et al. 2009

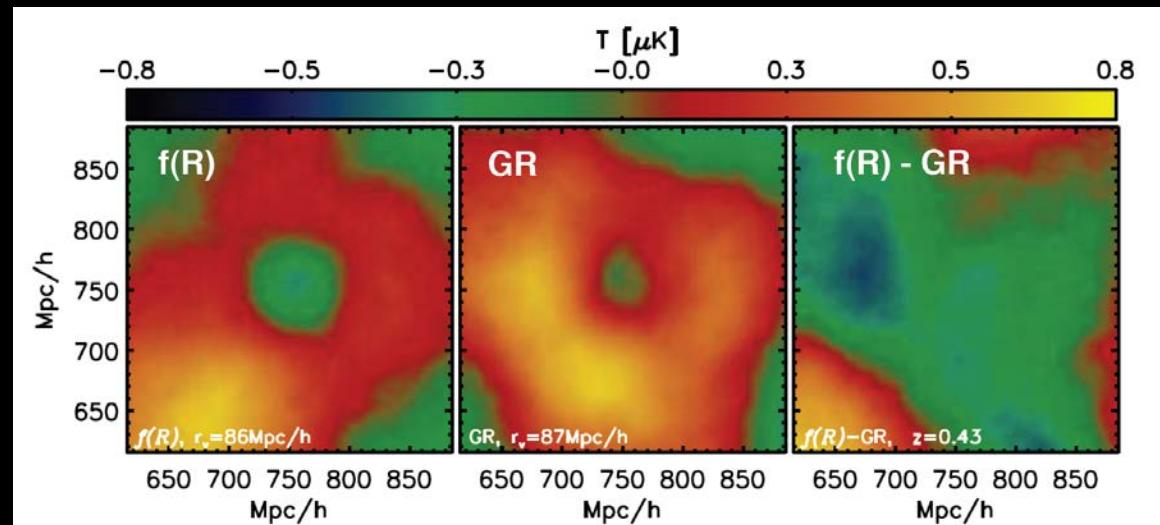
identification of
(LRG) supercluster & supervoid
Imprint on CMB

(Super)void & ISW



Cai et al. 2013:

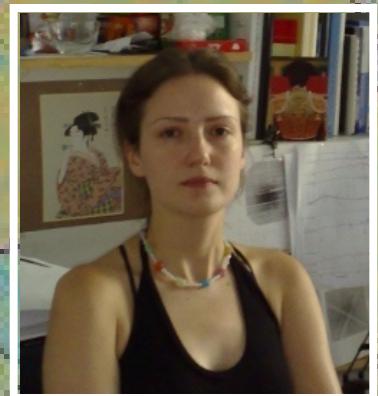
stacked ISW signal
may be used to discriminate
between gravity scenarios.



Void Galaxies

Void Galaxy Survey

- Kathryn Kreckel-Stanonik MPIA Heidelberg
- Burcu Beygu Kapteyn Inst.
- Rien van de Weygaert Kapteyn Inst.
- Jacqueline van Gorkom Columbia Univ.
- Thijs van der Hulst Kapteyn Inst.
- Jim Peebles Princeton Univ.
- Erwin Platen Kapteyn Inst.
- Miguel Aragon-Calvo JHU, Baltimore
- Katarina Kovac MPA, Garching



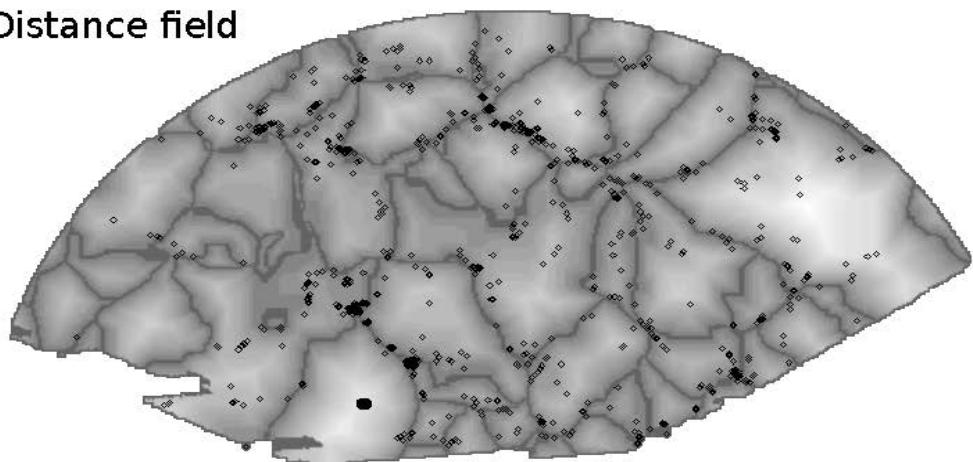
Void Galaxy Survey (VGS)

Multiwavelength Study

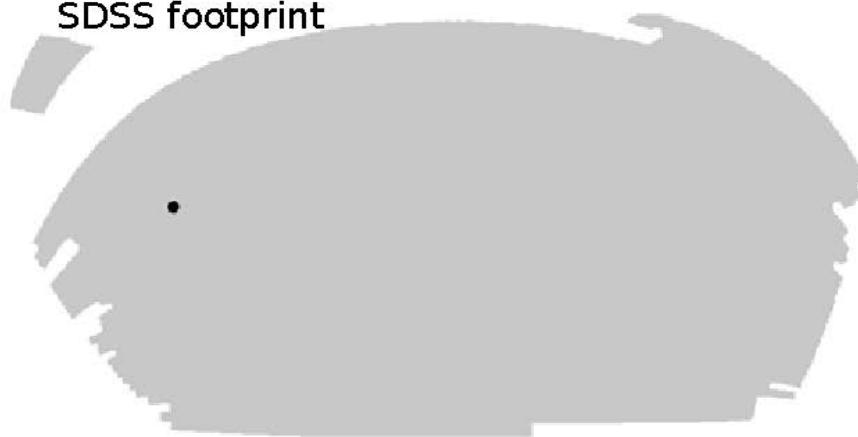
- ~ 60 galaxies located within/near the deep interior of voids in SDSS DR7
- gas content, star formation history, stellar content, kinematics & dynamics of void galaxies and their companions in a broad sample of void environments
- environmental influences on galaxy formation & evolution in isolated, supposedly pristine and undisturbed, void regions.
- Observational test theories cosmological structure formation

Geometric Void Galaxy selection

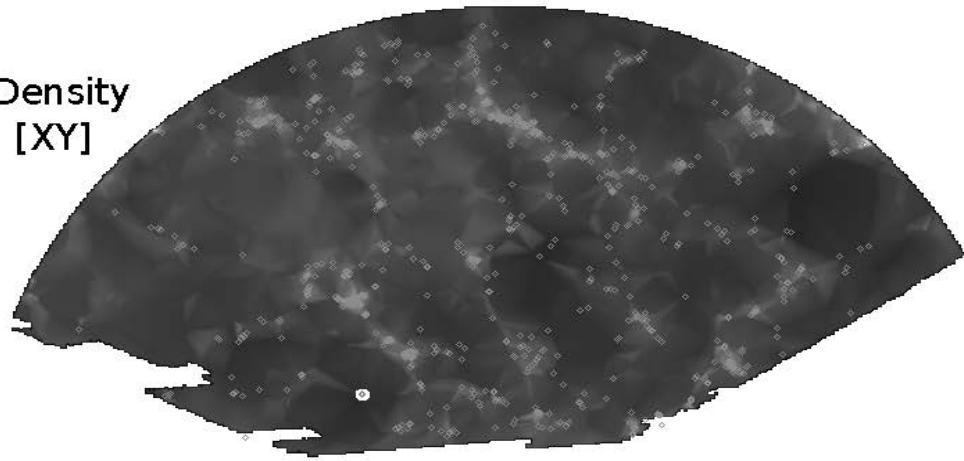
Distance field



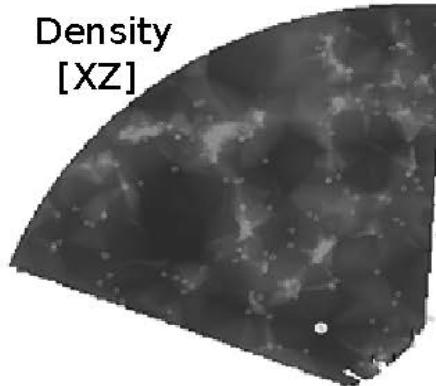
SDSS footprint



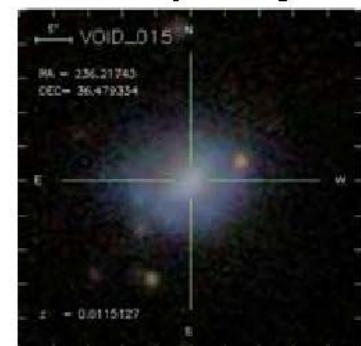
Density [XY]



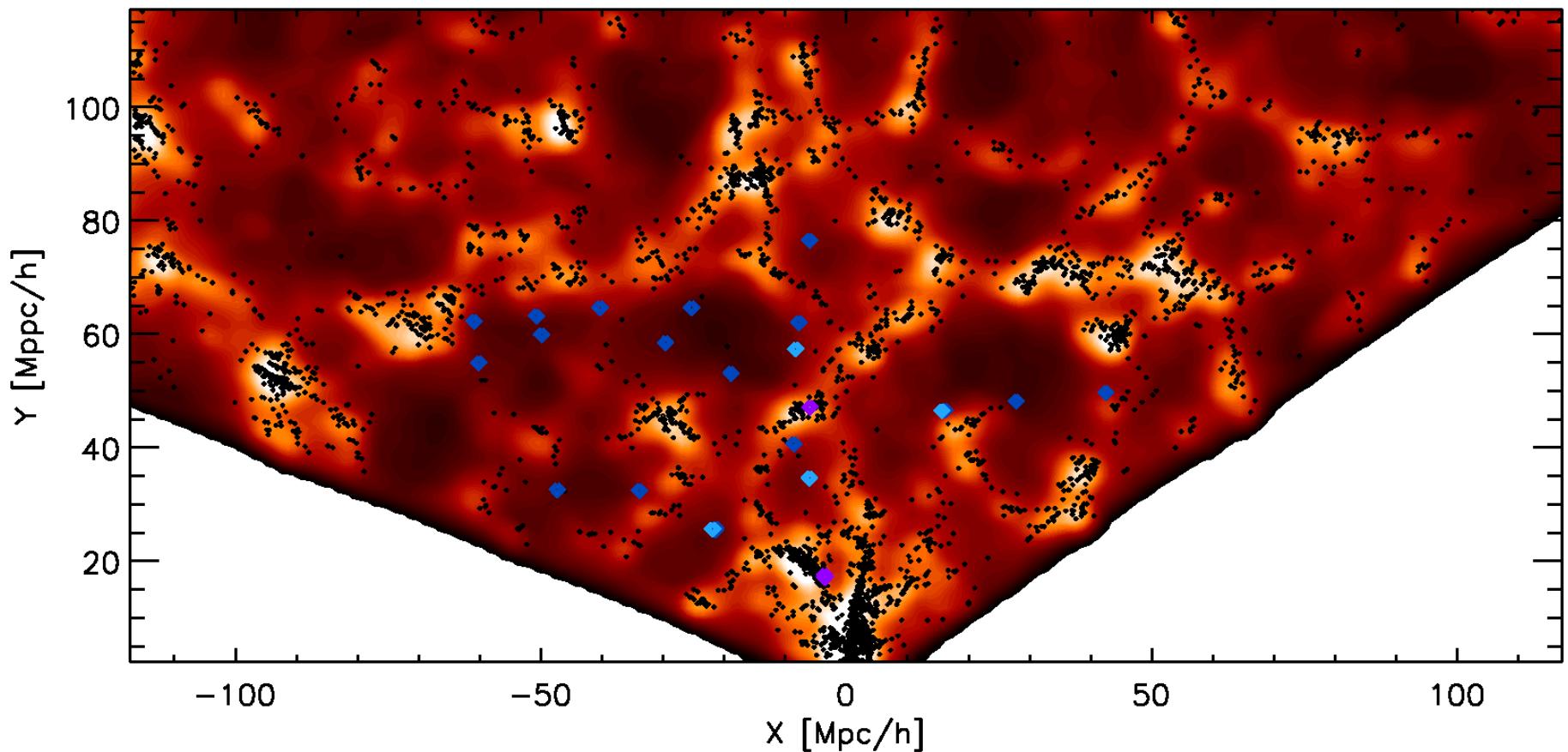
Density [XZ]



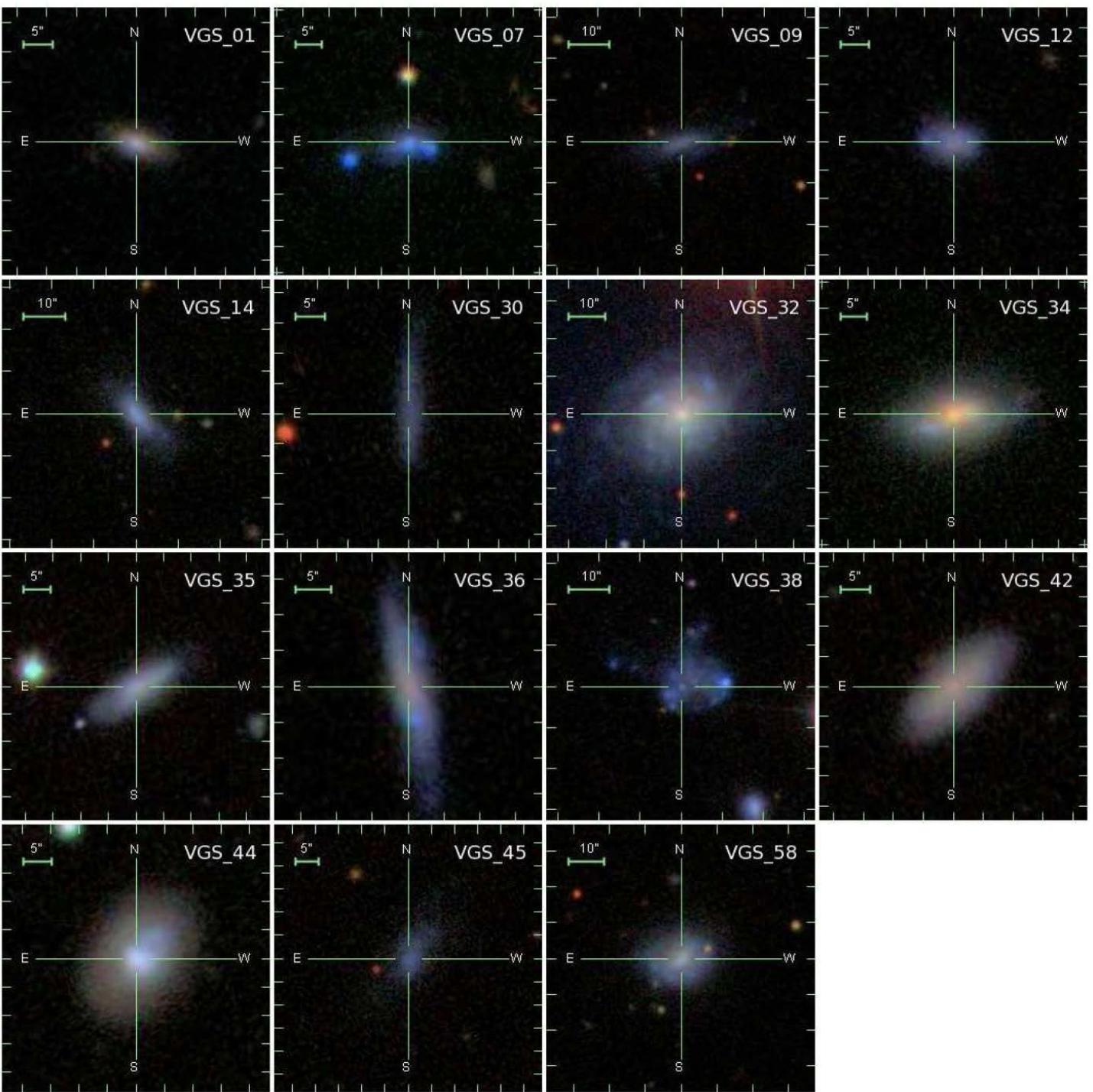
Galaxy image



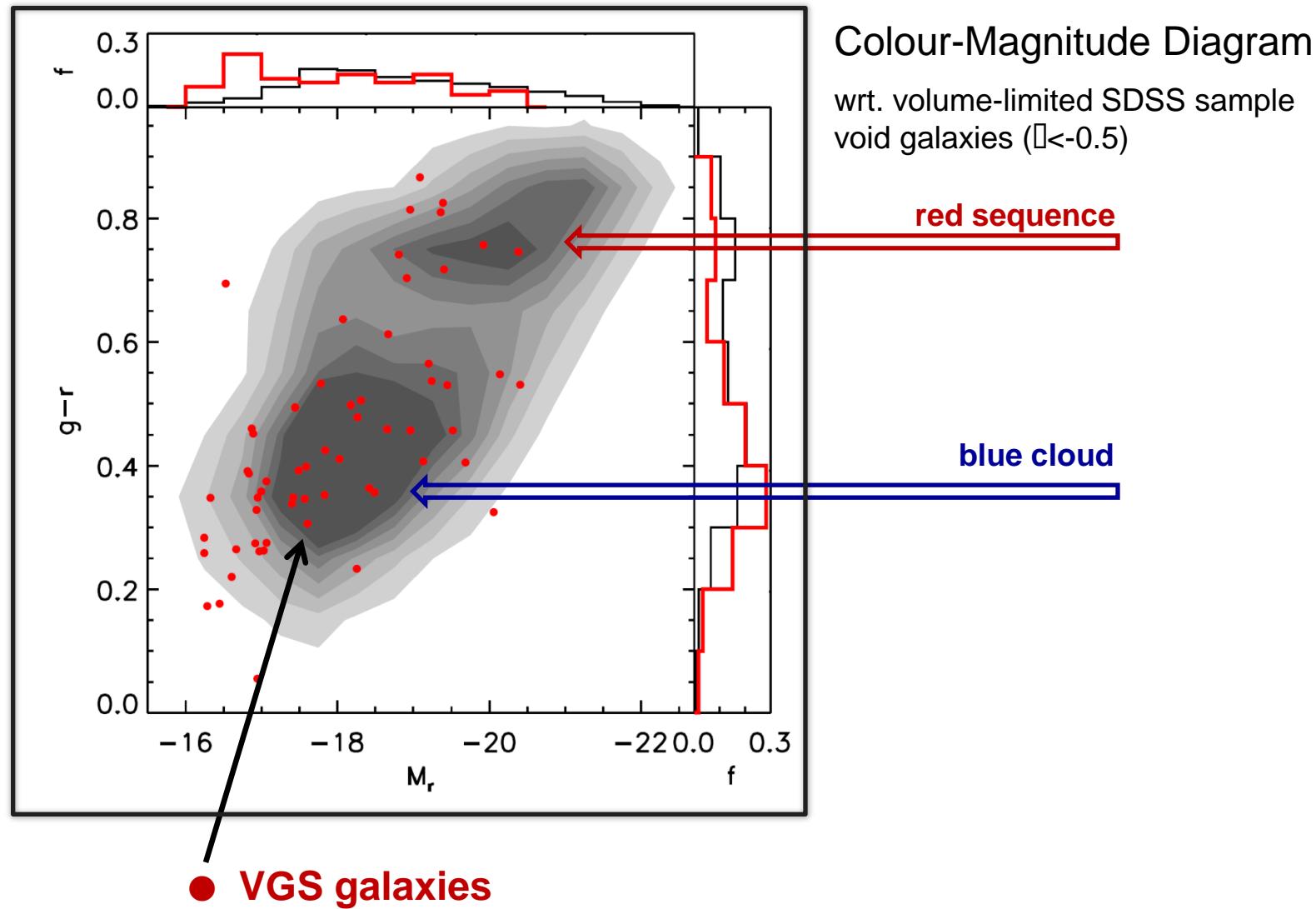
the Void Galaxy Pilot Sample



VGS Void Galaxies Sample (size scaled)

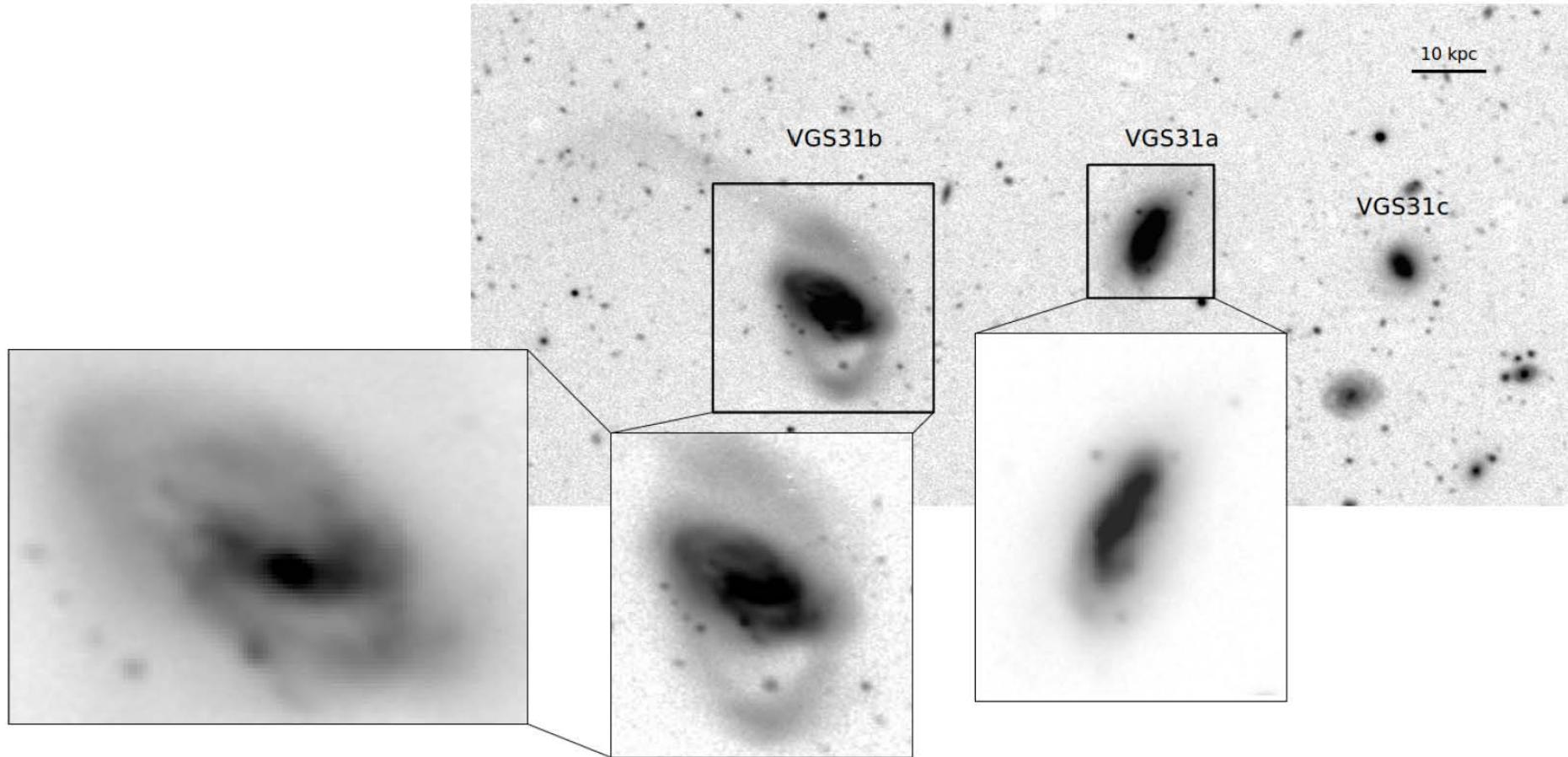


Optical Properties Selection

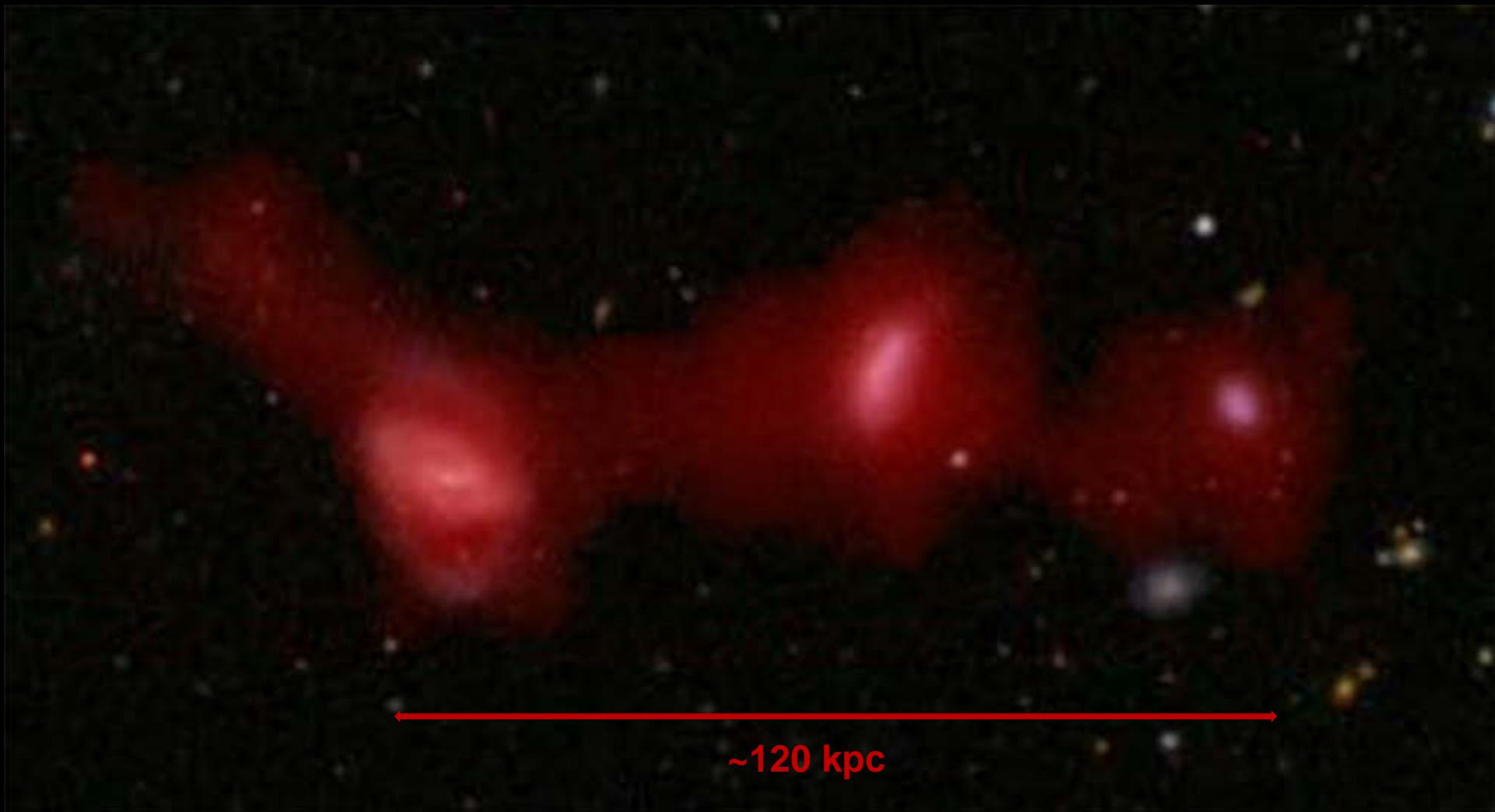


VGS_31

VGS_31: void galaxies along filament ?

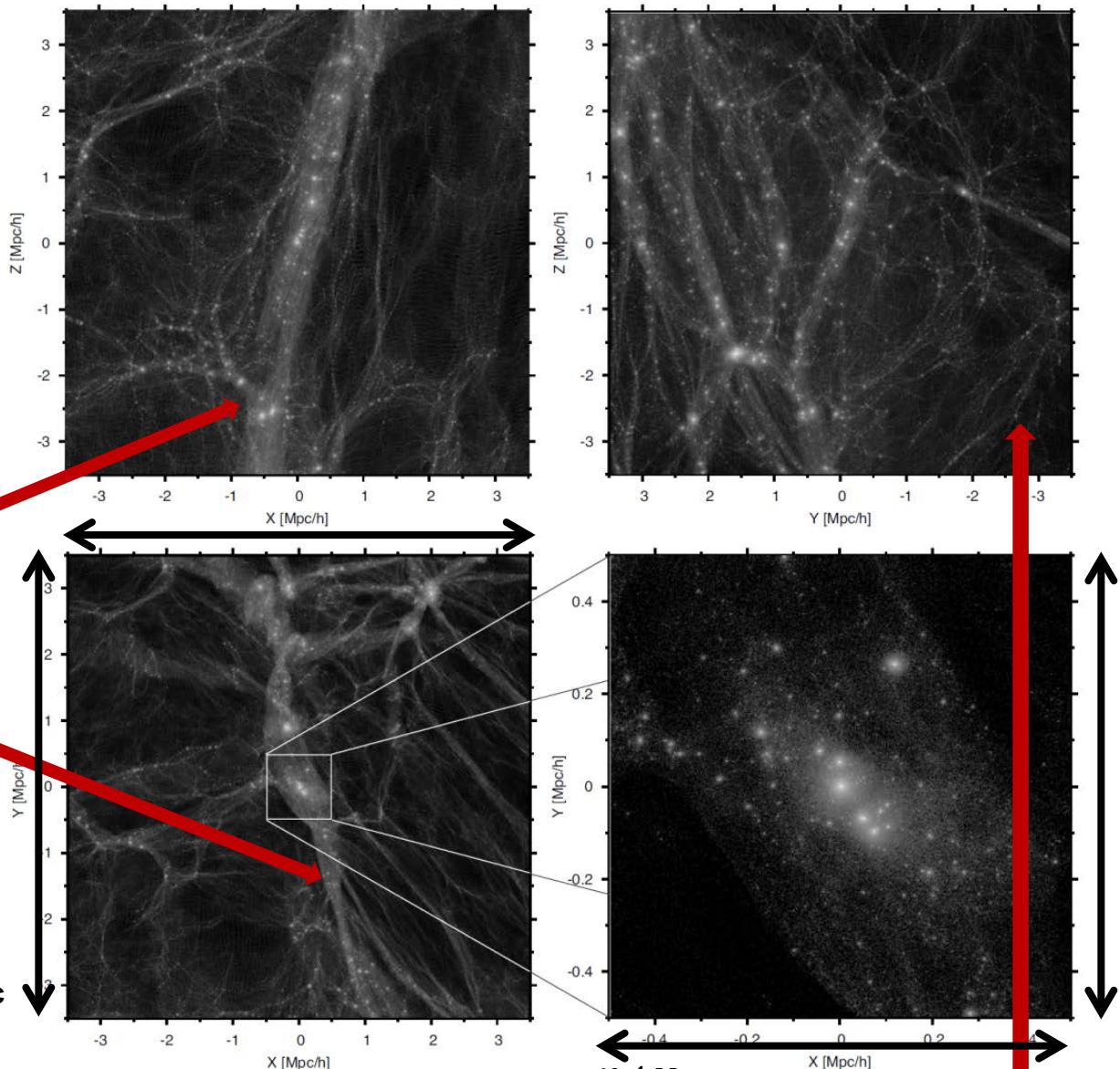


VGS_31: HI filament ?



CGV_G: on walls & filaments

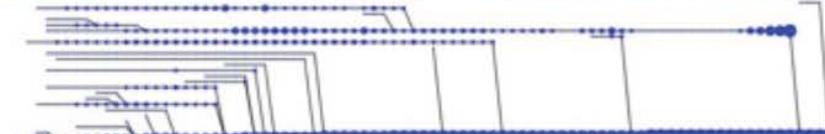
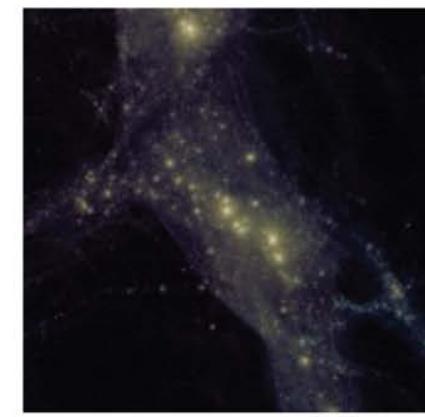
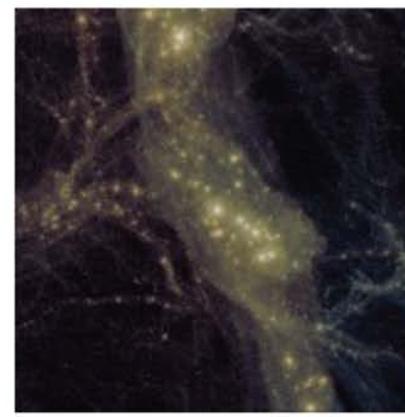
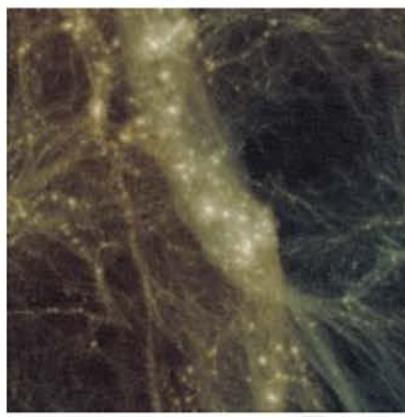
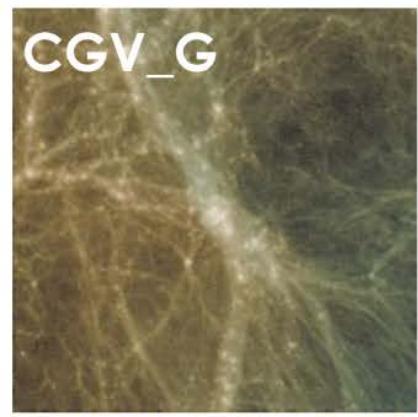
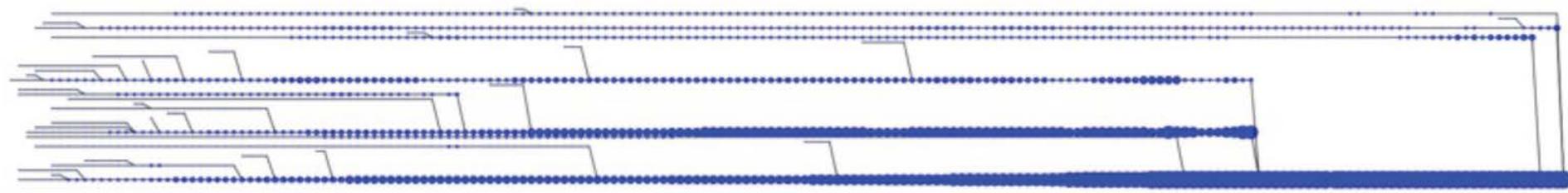
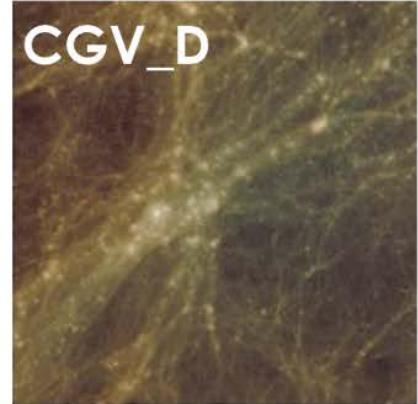
wall:
edge-on view



wall substructure:

- intricate filigree of tenous filaments
- filaments are/get largely aligned
- dynamical action of large-scale tidal force field

wall: face-on view



time [Gyr]

0

13.5

Conclusions

- **Voids are a highly interesting component of the Megaparsec matter and galaxy distribution**
- **Voids are a dynamically dominant component of the Cosmic Web**
- **Voids evolve hierarchically, and still reflect this through their multiscale structure**
- **Voids are very sensitive probes of dark energy**
- **Voids' isolated environment provides a very nice testing ground for the study of environmental influences on galaxy formation**
- **Voids may be sensitive probes of modified gravity**
- **Existence of supervoids (large underdense regions) may be challenging for cosmological standard scenario**