

# Zel'dovich's Legacy in Discovery and Understanding the Cosmic Web

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**University of Kansas**

# OUTLINE

- \* A few words about Ya.B. Zeldovich
- \* Cosmological background
- \* Zel'dovich approximation
- \* Adhesion approximation (Skeleton of the structure)
- \* Phase space and Lagrangian Submanifold
- \* Multi-Stream Field
- \* Flip-Flop Field
- \* Lagrangian skeleton and flip-flop field

Summary

# Relativistic Astrophysics and Cosmology

Zeldovich and independently Edwin Salpeter were the first to suggest that accretion disks around massive black holes are responsible for the huge amounts of energy radiated by quasars.

Zel'dovich and Starobinski showed Hawking that, according to the quantum mechanical uncertainty principle, rotating black holes should create and emit particles.

The Sunyaev-Zel'dovich effect (Scatter of CMB photons on hot electrons in clusters of gal.)

The Zeldovich spectrum of primordial fluctuations (scale-free power spectrum)

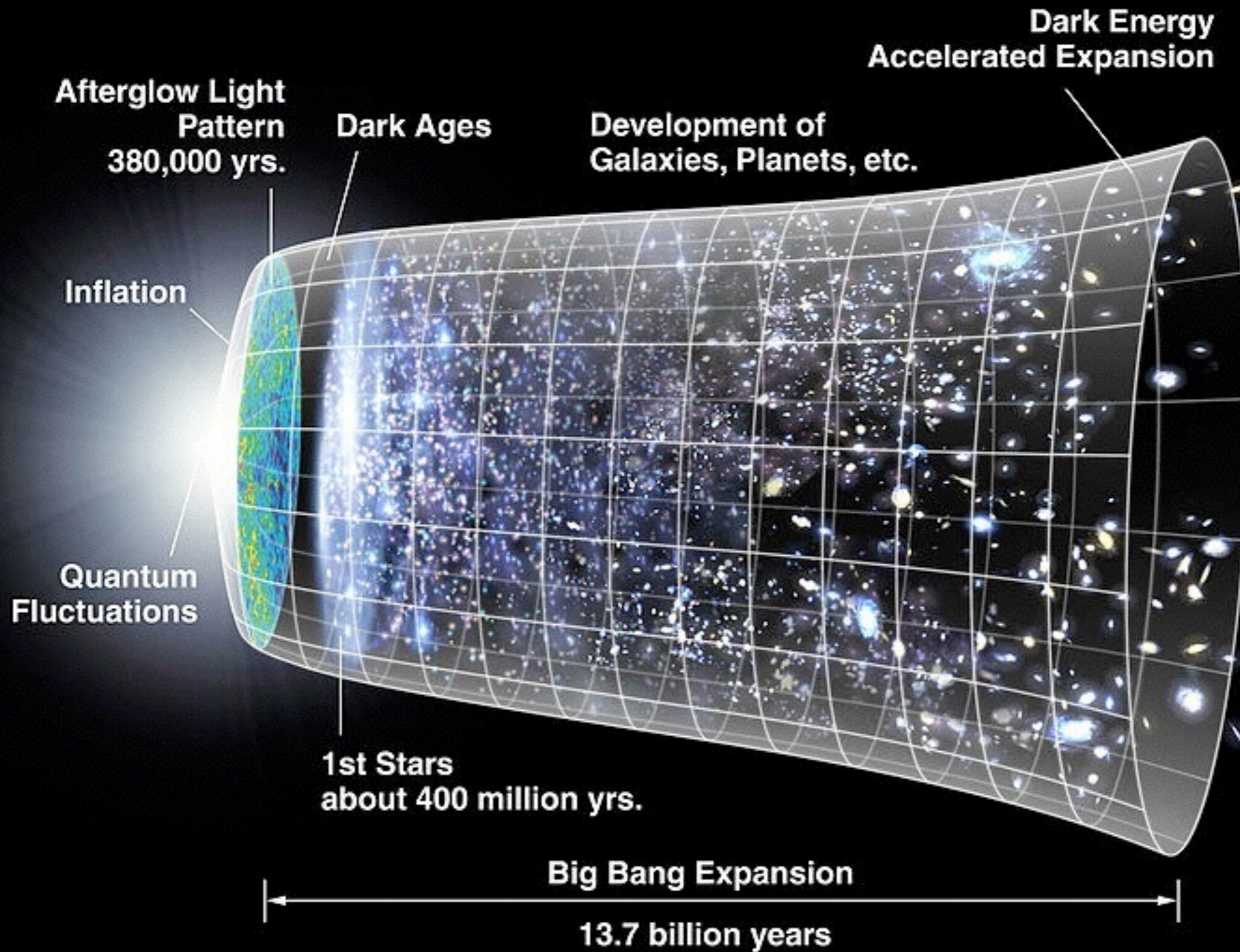
He argued that the relativistically-invariant theory of vacuum would result in non-zero minimum of the vacuum energy with the equation of state  $p_{vac} = -\epsilon_{vac}$ . Lambda term in Einstein eq. must be placed on the right hand side.

\*\*\*\*\* The Zel'dovich approximation in 1970 \*\*\*\*\*

# From recommendation for Zel'dovich to Academy of Sciences of USSR

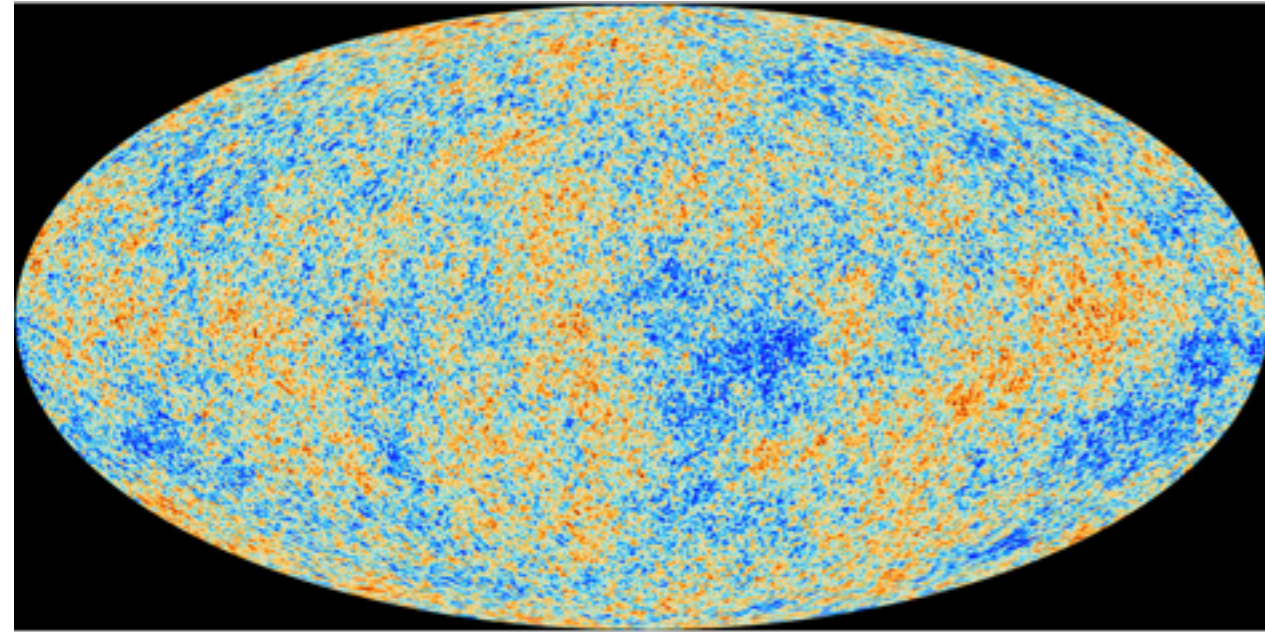
“It is characteristic for Ya.B. Zel'dovich to widely use the methods of hydrodynamics along with “conventional” methods of theoretical physics. This ability to use the both techniques **-very rare among theorists** - is a very advantageous trait of Zel'dovich, allowing him to solve problems which can be solved by neither pure hydrodynamicists nor “conventional” theorists.”

Lev Landau, 1946



# Structure in the Universe

*13.7 billion years ago*



Plank map

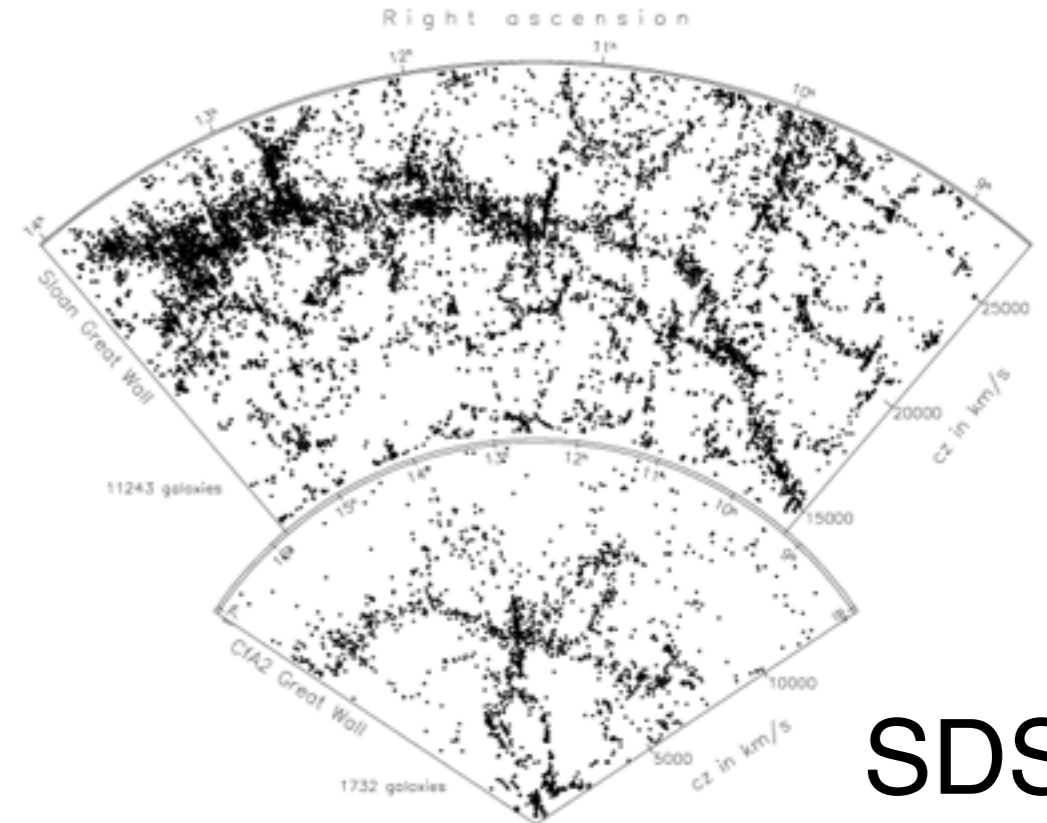
Temperature fluctuations of Cosmic Microwave Background

$T=2.73$  K

(fluctuations:  
of the order of  $1/100,000$ )

**NOW**

Below is the image in its original context on the page: [www.astro.princeton.edu/~mjuric/universe/](http://www.astro.princeton.edu/~mjuric/universe/)



**SDSS**

Galaxy distribution in  
a thin slice

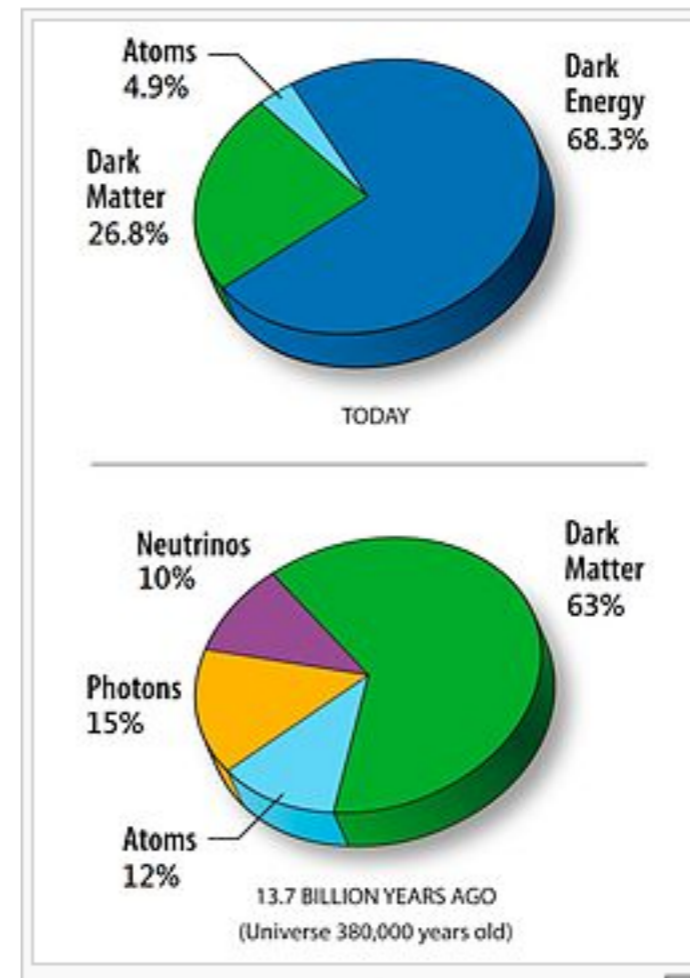
# Why did complexity begin to grow?

- Thanks to gravitational instability of Dark Matter

Why only Dark matter?

more mass – stronger gravity

NOW



*13.7 billion years ago*

Atoms **alone** could NOT develop structures because they were unable to overcome the expansion of the universe!

That's why we focus on the growth of complexity in Dark Matter.

**Gravitational Instability:  
An Approximate Theory for Large Density Perturbations**

**YA. B. ZELDOVICH**

**Institute of Applied Mathematics, Moscow**

**Received September 19, 1969**



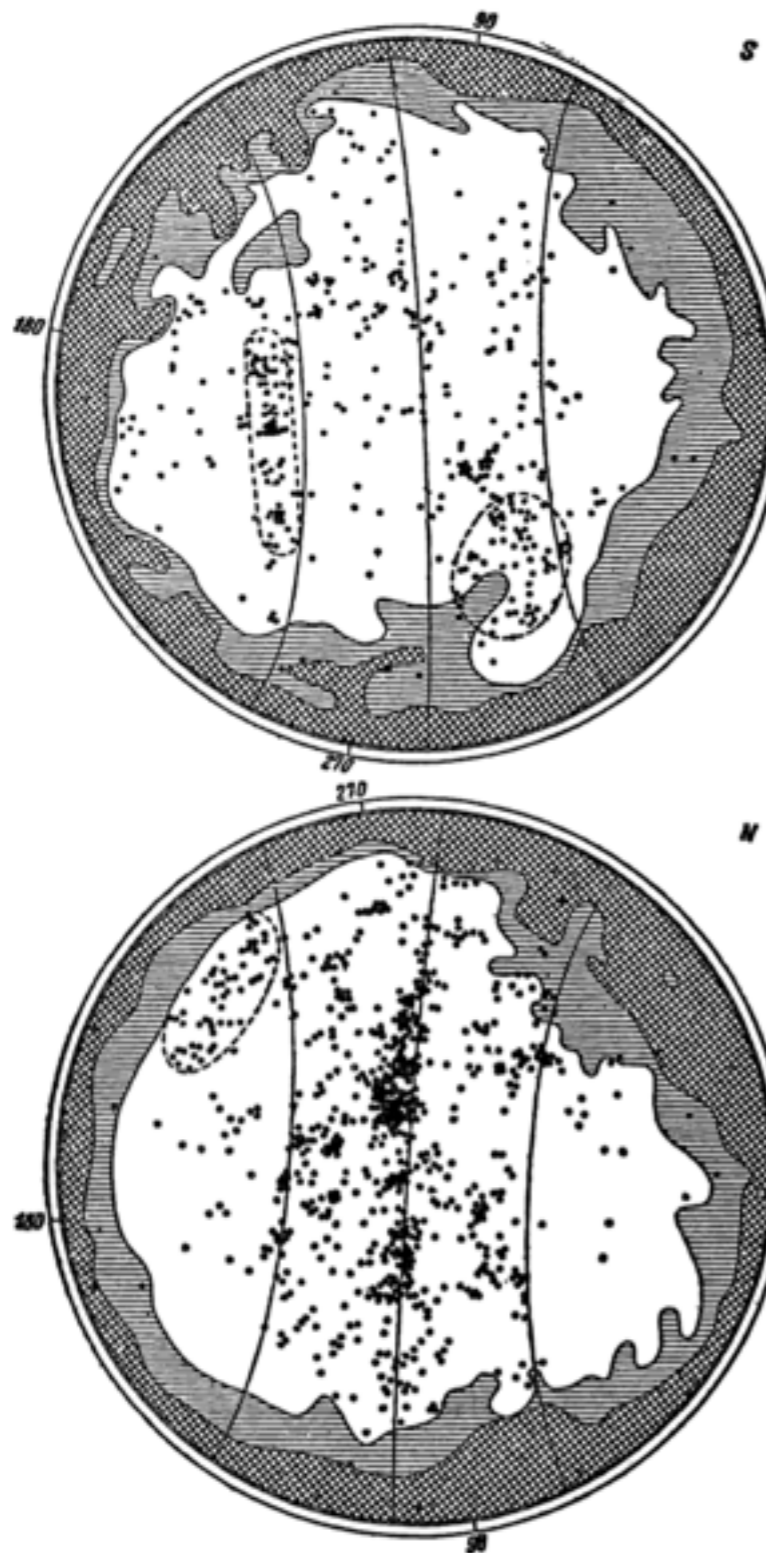
# Large-Scale Structure of the universe in 1970

## OBSERVATIONS:

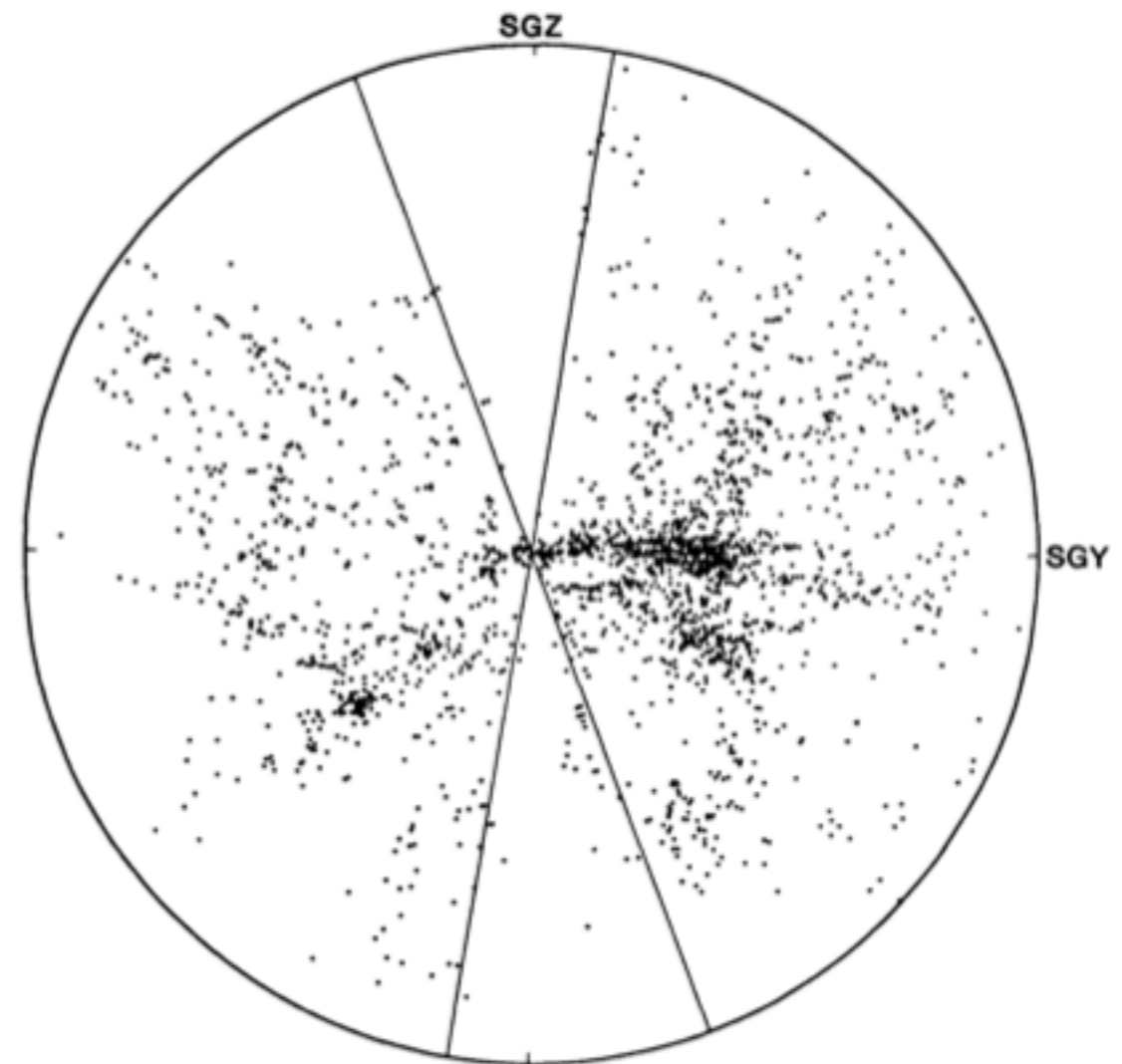
**de Vaucouleurs (1956), (1960) : The Local Supercluster**

**The first observational results that claimed the discovery of Superclusters of galaxies**

**Chincarini & Rood 1976, Gregory & Thompson 1978**



**Figure 2.** Distribution of the Shapley-Ames galaxies (1932) in (old) galactic coordinates. The zone of avoidance (dark) and of partial obscuration (grey) by the Milky Way is indicated. The supergalactic equator and parallels at  $\pm 30^\circ$  latitude are marked. Two external galaxy clouds in Hydra ( $l^I = 240^\circ$ ) and Pavo-Indus ( $l^I = 310^\circ$ ) and the elongated Dorado-Fornax-Eridanus stream or "southern supergalaxy" are outlined.



**Figure 3** All 2175 galaxies in the Nearby Galaxy Catalog (NBG) projected onto the SGY-SGZ plane. The SGY-axis is directed toward supergalactic longitude  $90^\circ$ , supergalactic latitude  $0^\circ$  ( $l^{II} = 227^\circ$ ,  $b^{II} = +83^\circ.7$ ), the SGZ-axis toward supergalactic latitude  $90^\circ$  ( $l^{II} = 47^\circ.4$ ,  $b^{II} = +6^\circ.3$ ). The radius of the outer boundary is 60 Mpc. The galactic zone of avoidance ( $b < 15^\circ$ ) is contained within the opposed wedges tilted by  $6^\circ$  with respect to the SGZ-axis. There is a zone of incompleteness ( $\delta < -45^\circ$ ), which is projected across most of the southern supergalactic hemisphere. Figures 3–6 are reproduced by courtesy of R. B. Tully (92).

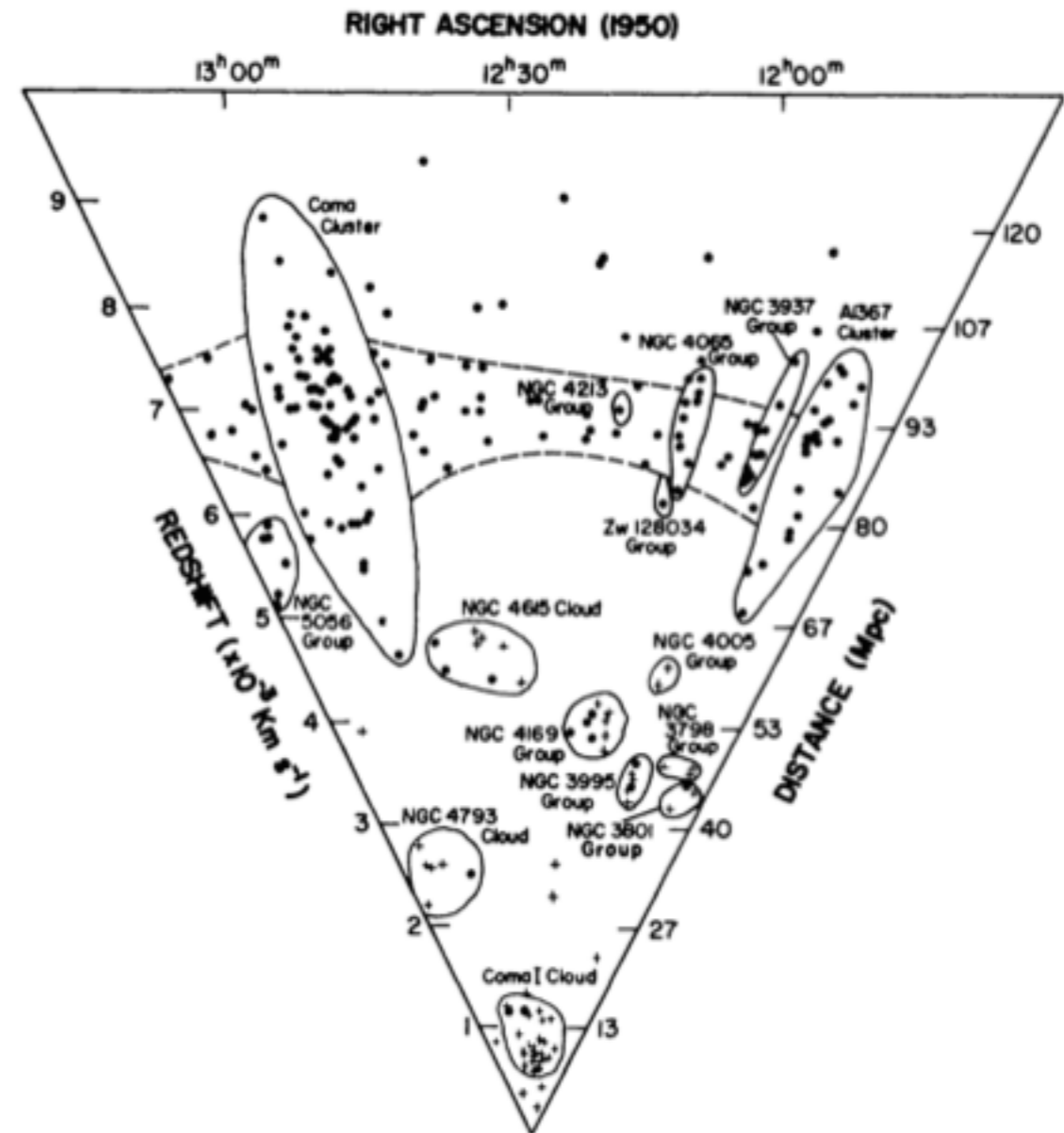
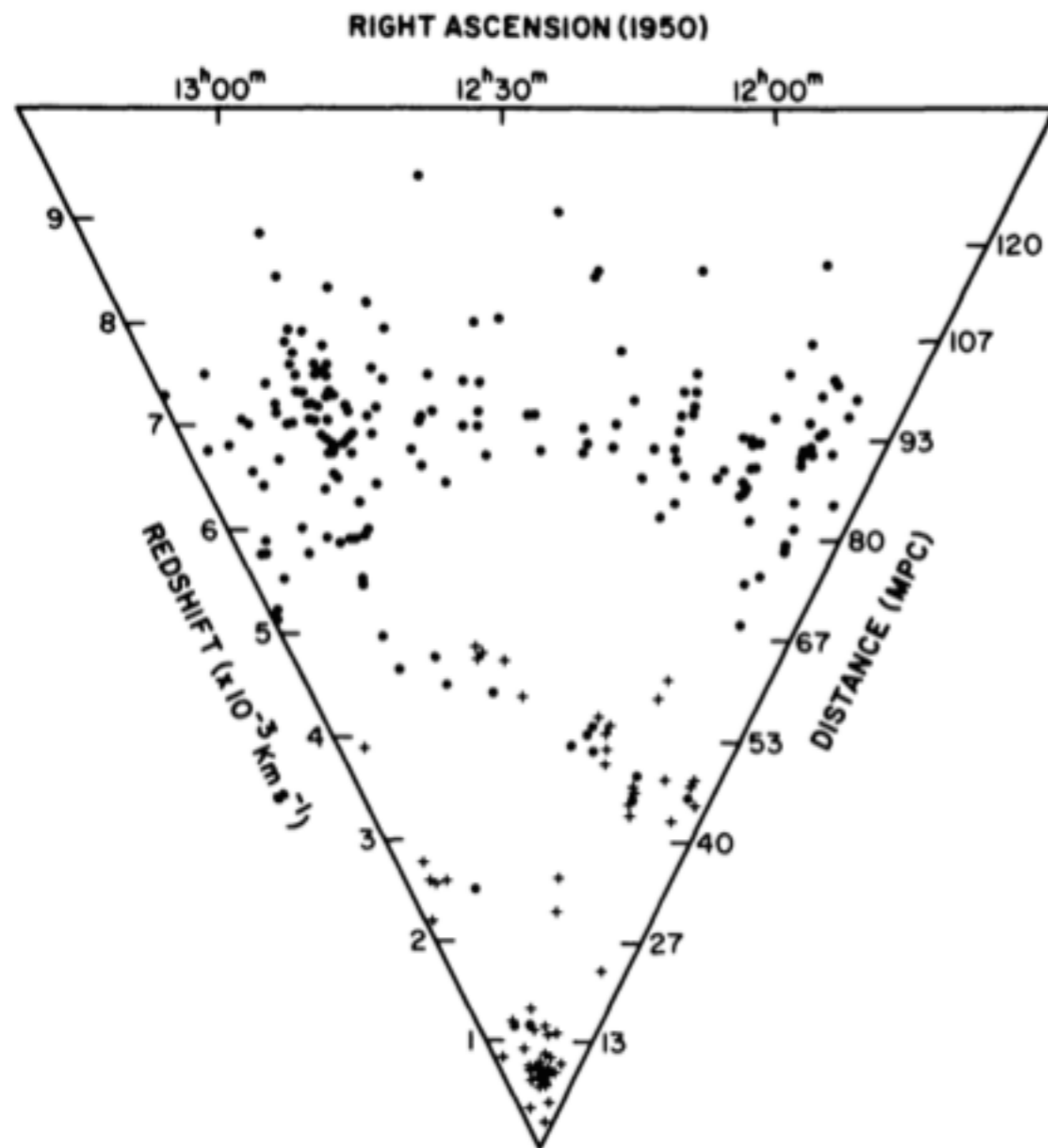


Figure 12 “Wedge diagram” of the Coma supercluster [Gregory & Thompson (40)]. As the supercluster is elongated in the east-west direction, right ascensions have been chosen as position coordinates; the galaxies lie between  $+19^\circ$  and  $+32^\circ$  declination. The angular size has been magnified about two times compared with the indicated distance scale.

Gregory & Thompson 1978  
 see also Chincarini & Rood 1976

# Large-Scale Structure of the universe in 1970

## THEORY:

“Gravitational Instability is similar to capitalism:

in the course of time  
rich become richer  
while  
poor become poorer

and the gap between the rich and poor is widening  
and becomes catastrophic.”

Zel'dovich

Z=14.2

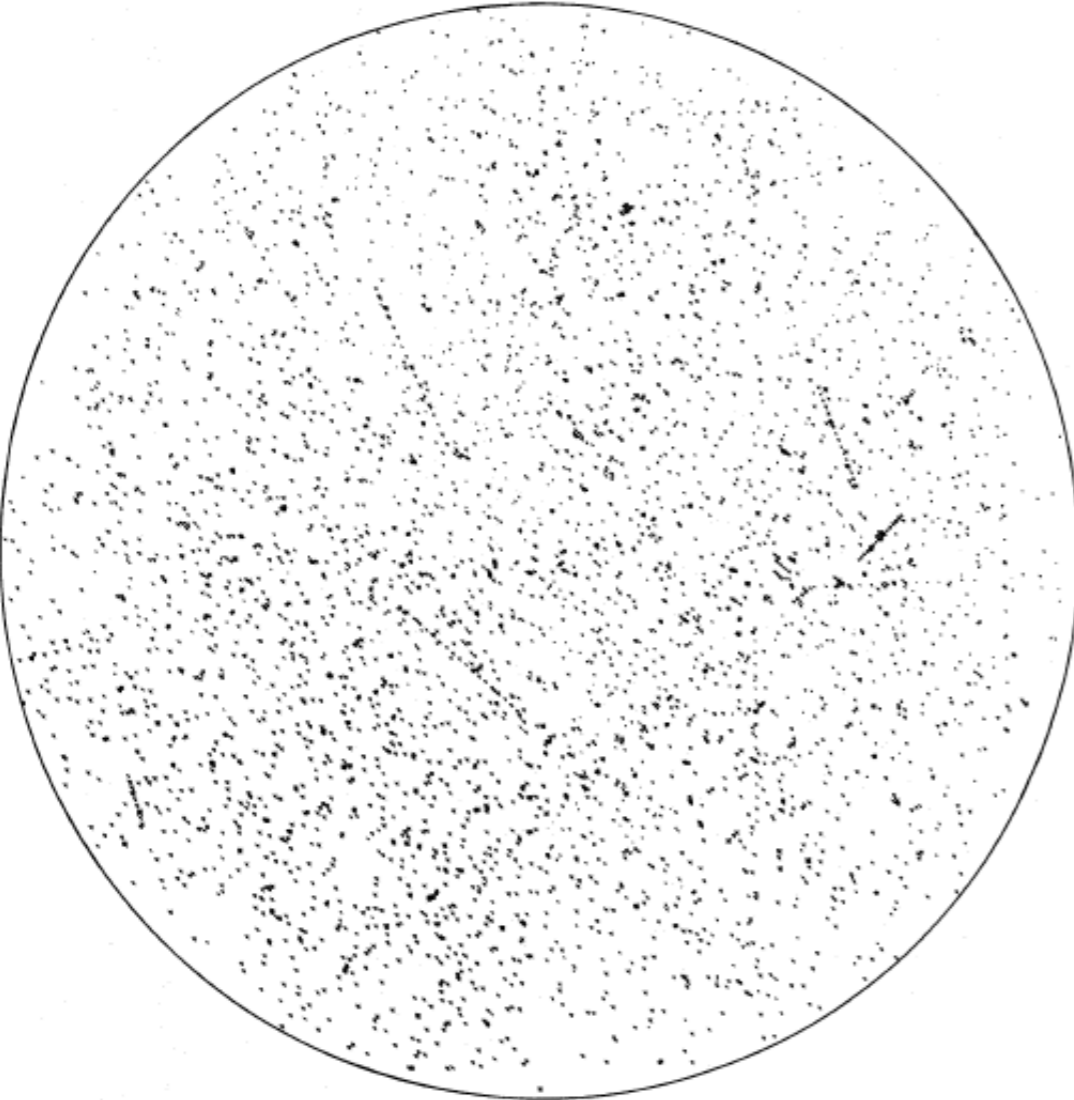


FIG. 1a

Z=0

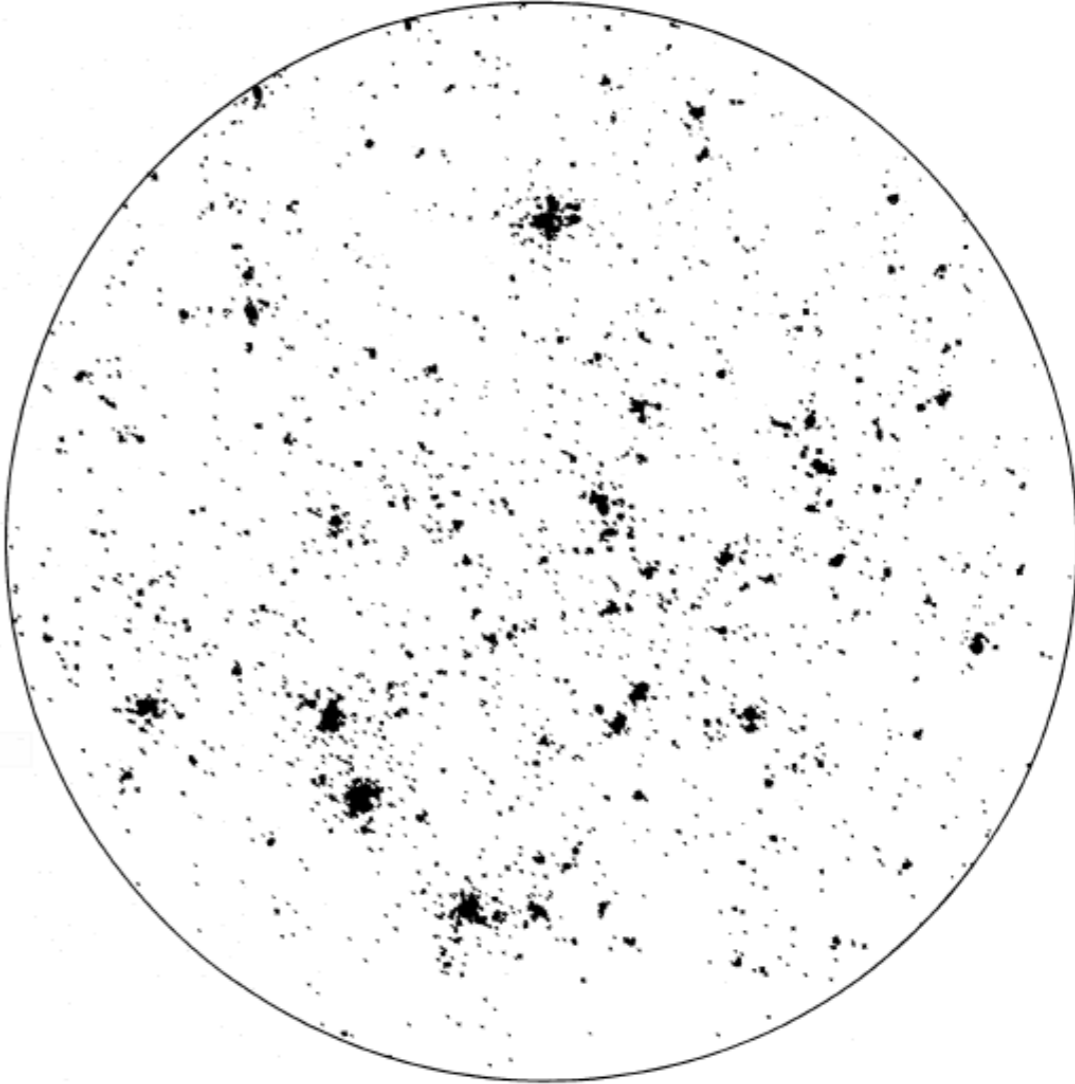


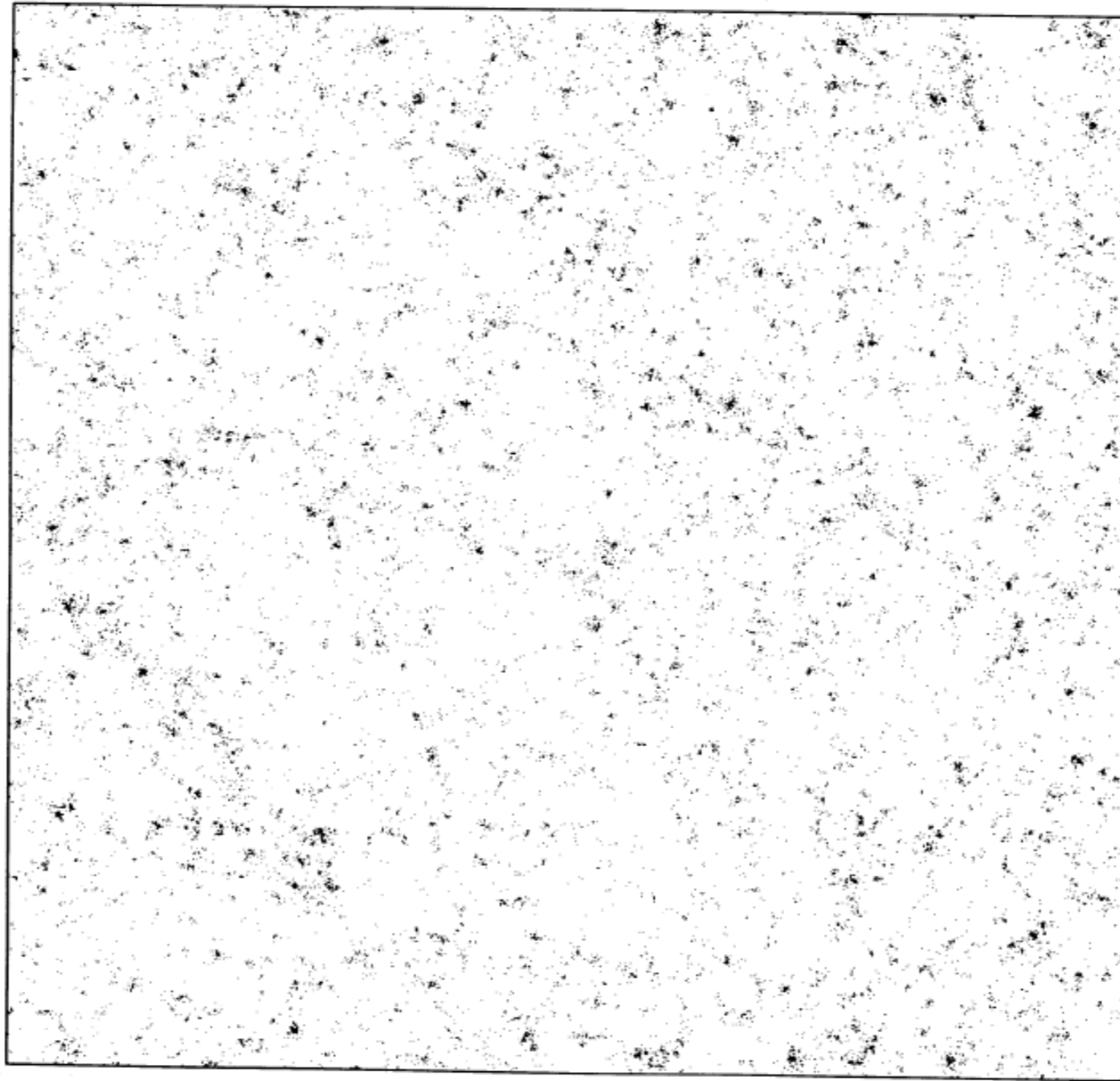
FIG. 1f

# State of art N-body simulations

1981 Efstathiou, Eastwood, *MNRAS*, 194, 503

*Clustering of particles in an expanding Universe*

511



**Figure 1.** *X–Y* projection of the particle positions for a 20 000-body numerical experiment after the system has expanded by a factor of 9.9. In this case the expansion follows that of an Einstein–de Sitter model,  $\Omega_0 = 1.0$ .

**Gravitational Instability:  
An Approximate Theory for Large Density Perturbations**

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# Zel'dovich approximation (1970)

Comoving coordinates:  $r_i$ ,

Zel'dovich approximation is a map:  $r_i(\mathbf{q}, t) = q_i + D(t)s_i(\mathbf{q})$

If  $\Phi(\mathbf{q})$  is the linear perturbation of grav. potential then  $s_i(\mathbf{q}) = -\partial\Phi/\partial q_i$

Density can be found from the conservation of mass

$$\rho(\mathbf{q}, t) = \bar{\rho}(t) \left| \frac{\partial r_i}{\partial q_k} \right|^{-1} = \bar{\rho} \left[ (1 - D(t)\alpha(\mathbf{q}))^{-1} [(1 - D(t)\beta(\mathbf{q}))^{-1} [(1 - D(t)\gamma(\mathbf{q}))^{-1} \right]$$

$\alpha(\mathbf{q}) \geq \beta(\mathbf{q})$  and  $\beta(\mathbf{q}) \geq \gamma(\mathbf{q})$  are the eigen values of the deformation tensor

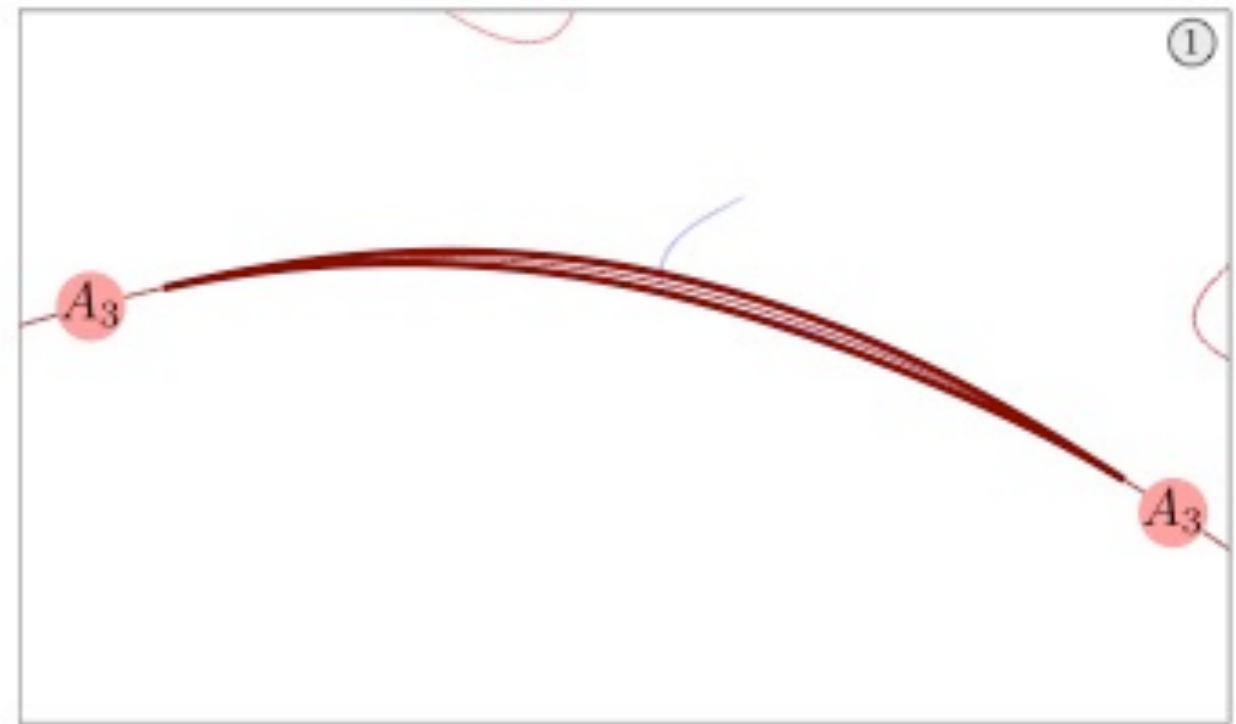
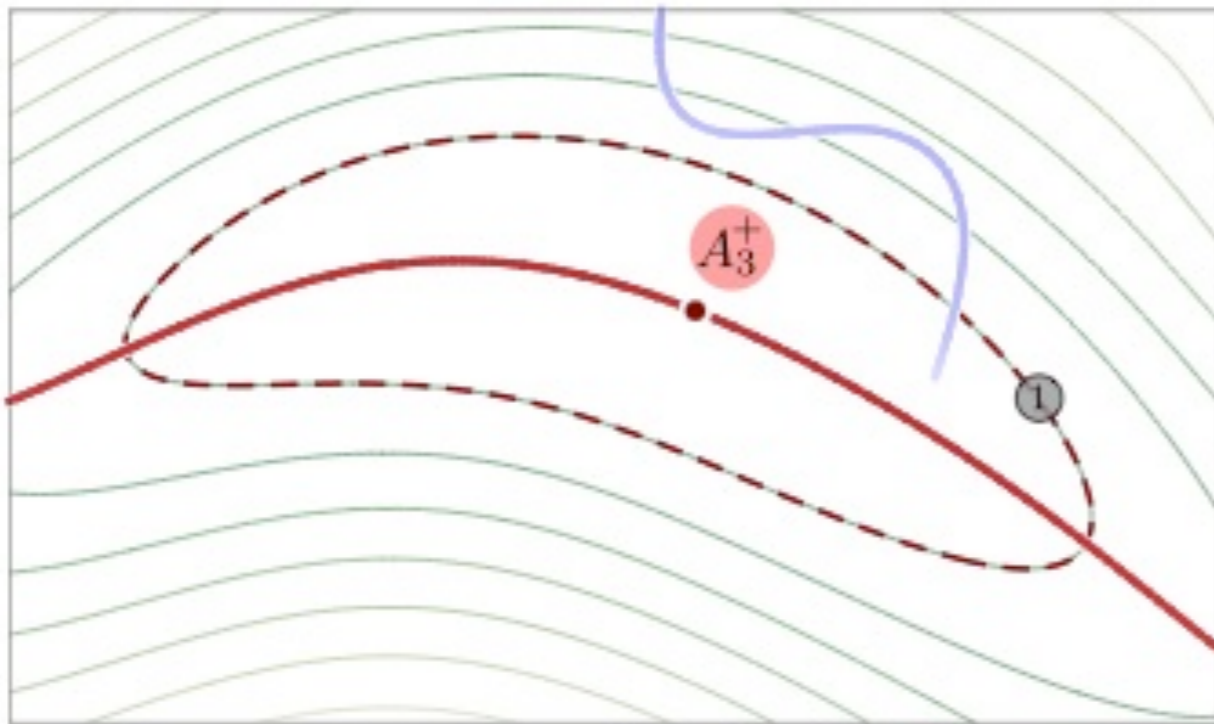
$$d_{ik}(\mathbf{q}) = \frac{\partial s_i}{\partial q_k} = -\frac{\partial^2 \Phi}{\partial q_i \partial q_k}$$

Linear density fluctuations:  $\delta\rho/\rho = D(t)(\alpha + \beta + \gamma)$ .

The Zel'dovich approximation describes anisotropic collapse and motion.





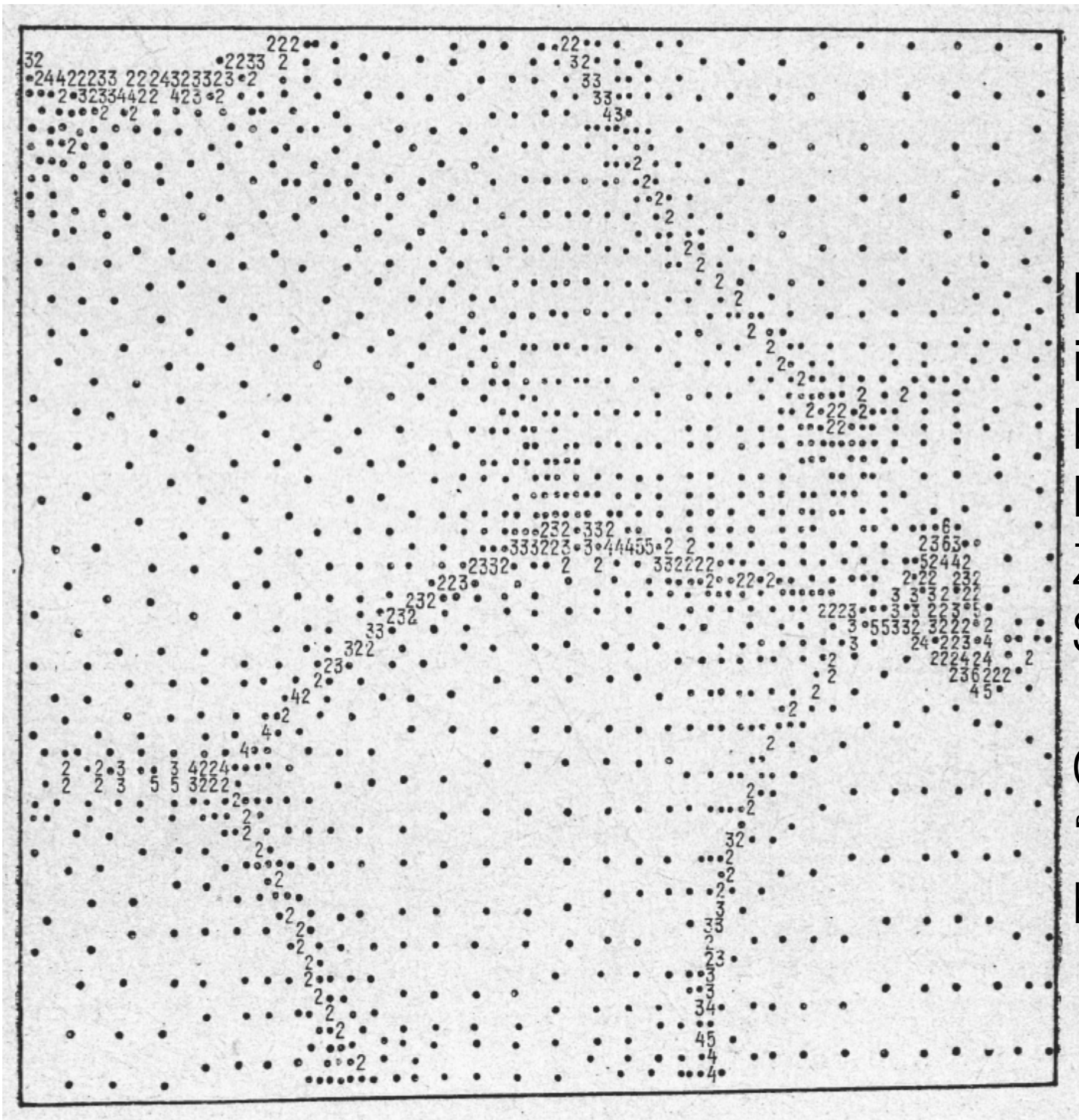


Hidding, Shandarin, van de Weygaert 2014

P.J.E. Peebles 1980 The Large-Scale Structure of the Universe

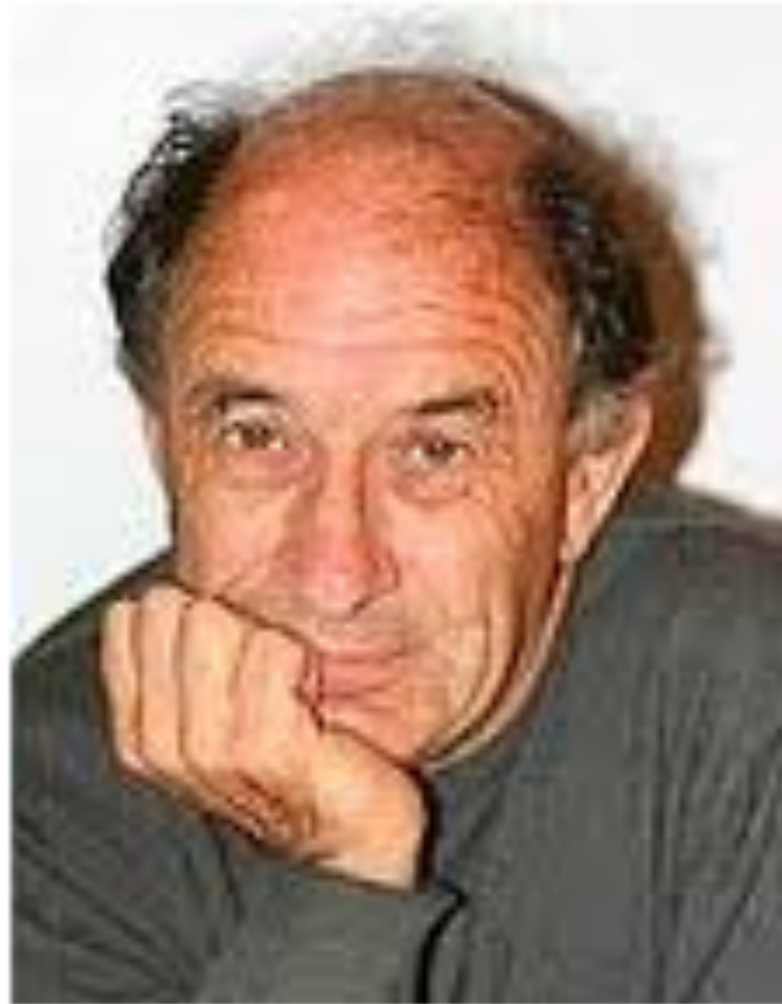
8 pages on Caustics and Pancakes with the verdict: death

Assumptions are not realistic, approximation is kinematic,  
pancakes are unstable therefore no observational traces remain



Published  
in a review paper  
by  
Doroshkevich,  
Zeldovich,  
Sunyaev 1976.

Caption says:  
“the figure is made  
by S. Shandarin”



Vladimir Igorevich  
Arnold  
1937 - 2010

*“Mathematics is a part of physics.*

*Physics is an experimental science, a part of natural science.*

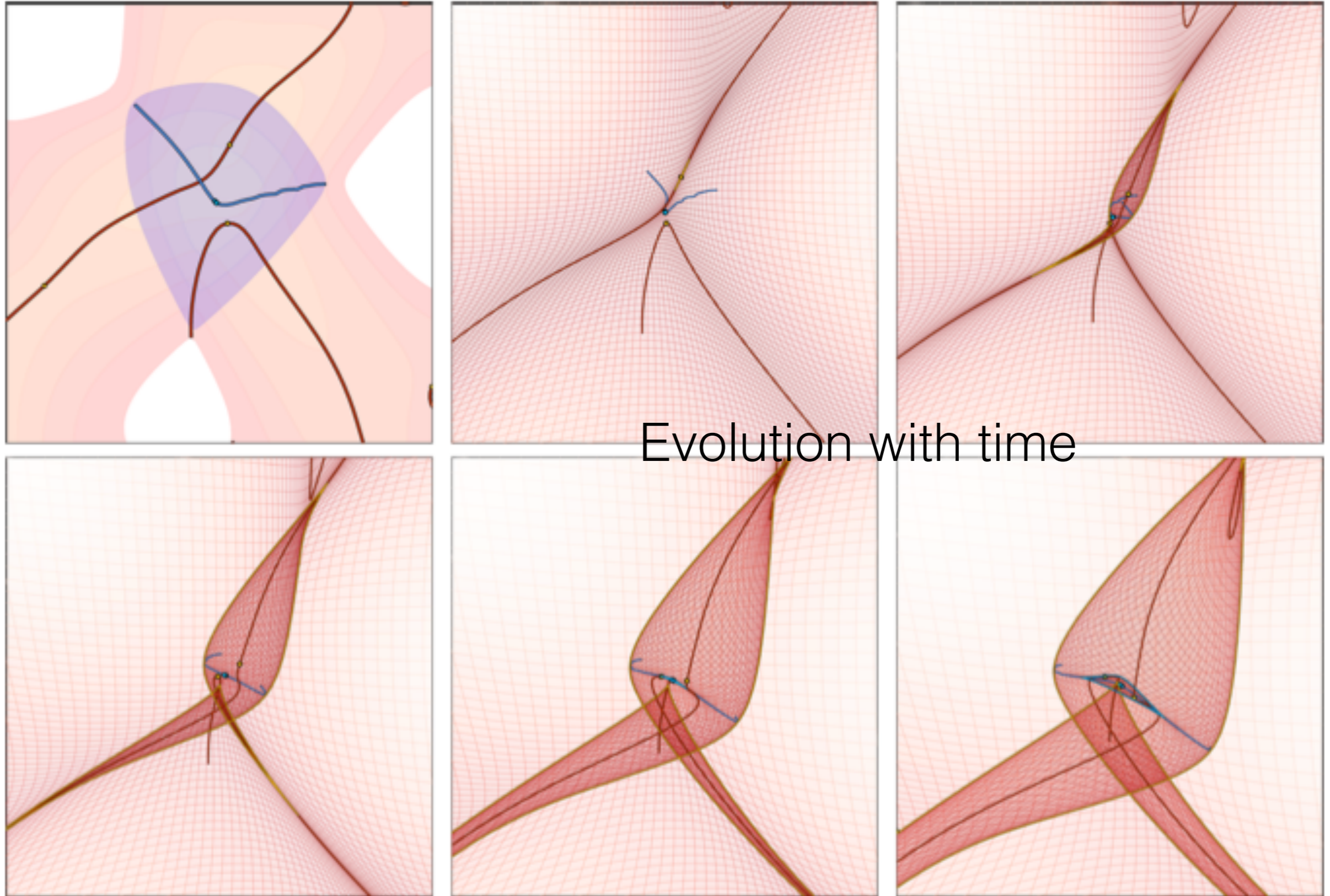
*Mathematics is the part of physics where experiments are cheap.”*

ADE classification of caustics in 3D (Arnold 1981)

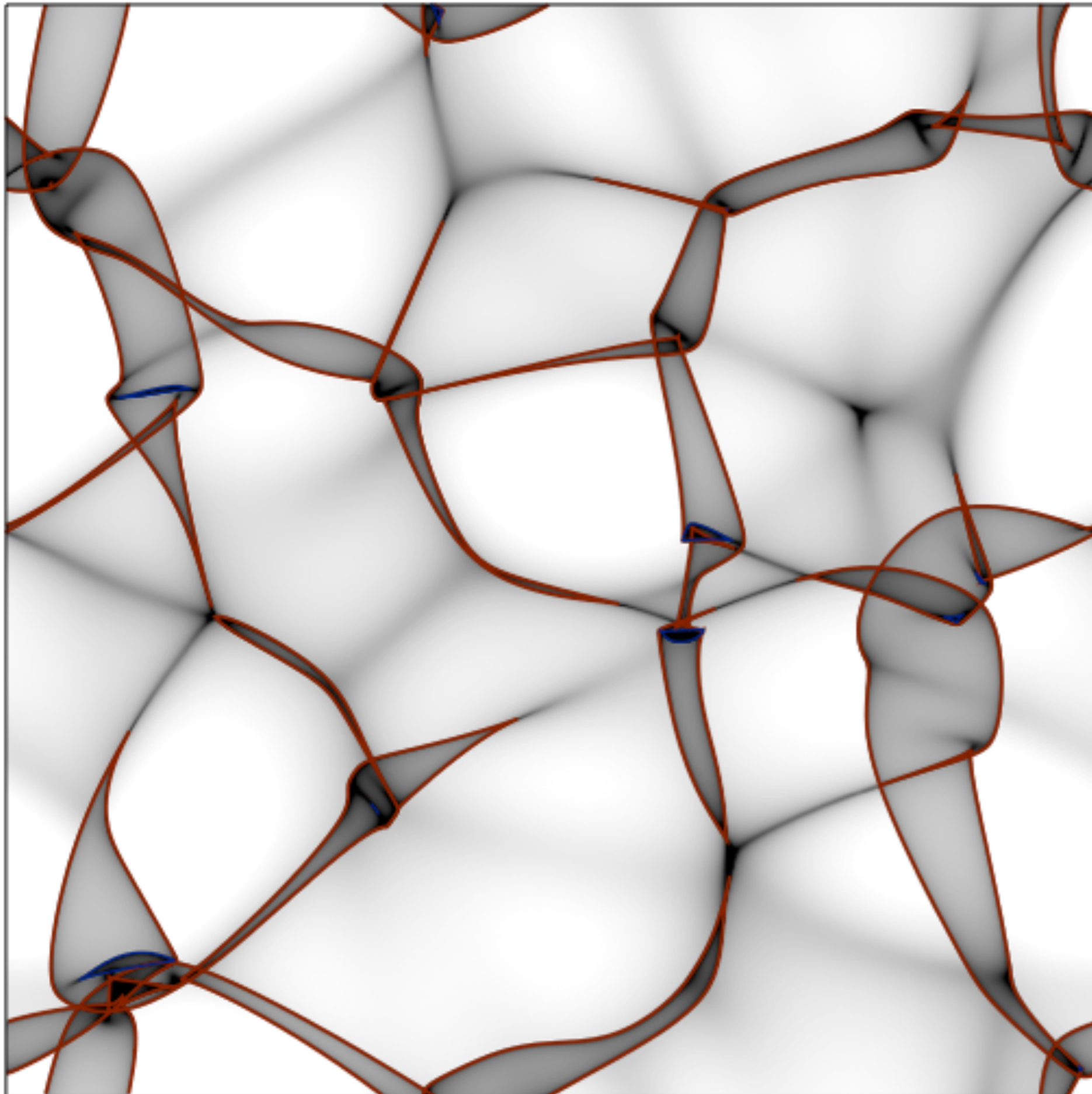
Arnold  
Shandarin  
Zeldovich  
1981

$n=3$ , Euler space, series A				$n=3$ , Euler space, series D					
type	instantaneous caustics			bicaustic	type	instantaneous caustics			bicaustic
	$t < 0$	$t = 0$	$t > 0$			$t < 0$	$t = 0$	$t > 0$	
$A_3$					$D_4^-$				
$A_3(+)$	$\emptyset$	$\circ$			$D_4^-(+)$				
$A_3(-)$					$D_4^+(+)$				
$A_4$					$D_4^+(-)$				
$A_4(+)$					$D_5$				
$A_4(-)$									
$A_5$									

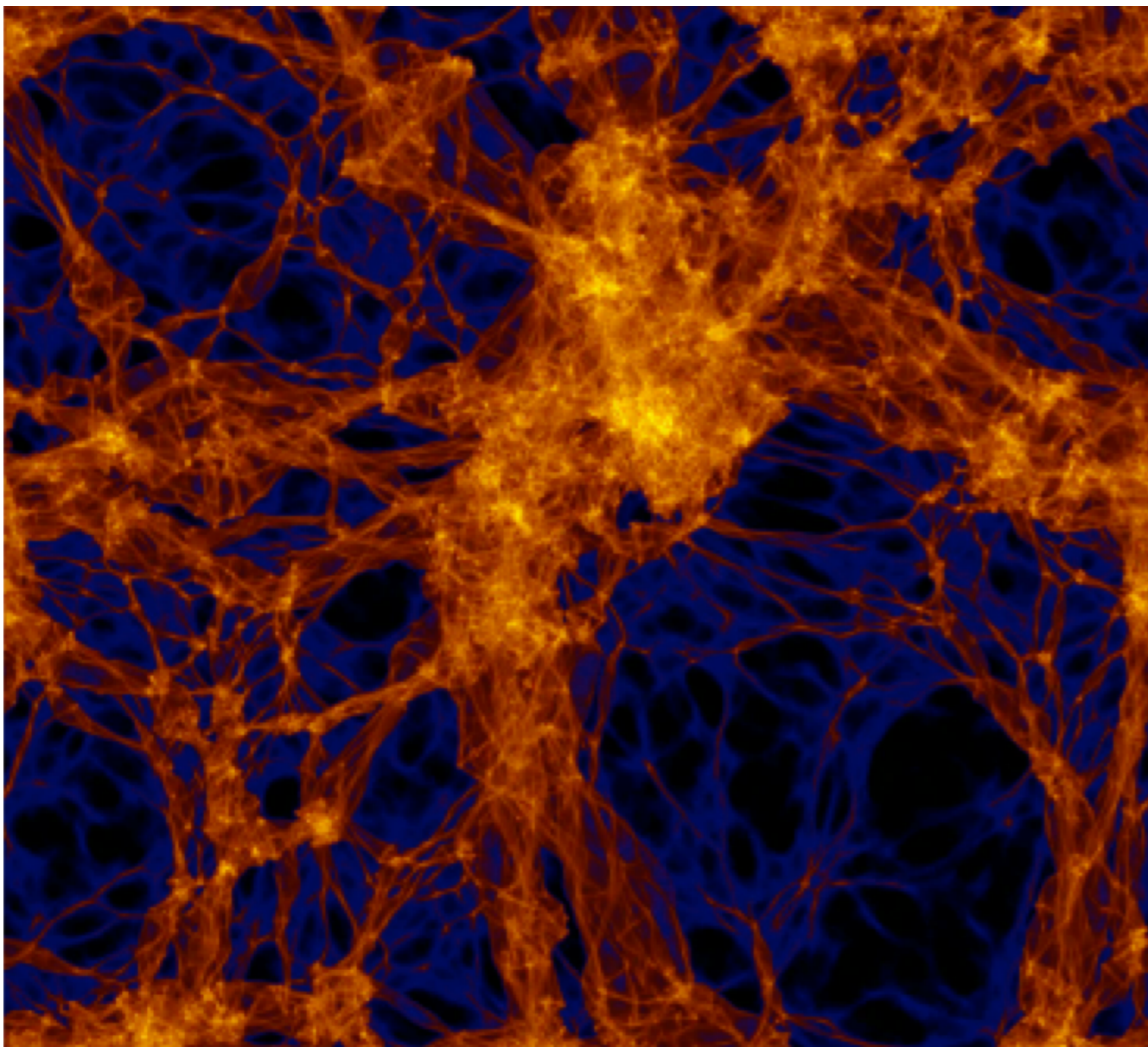
# Pancake connectivity in 2D



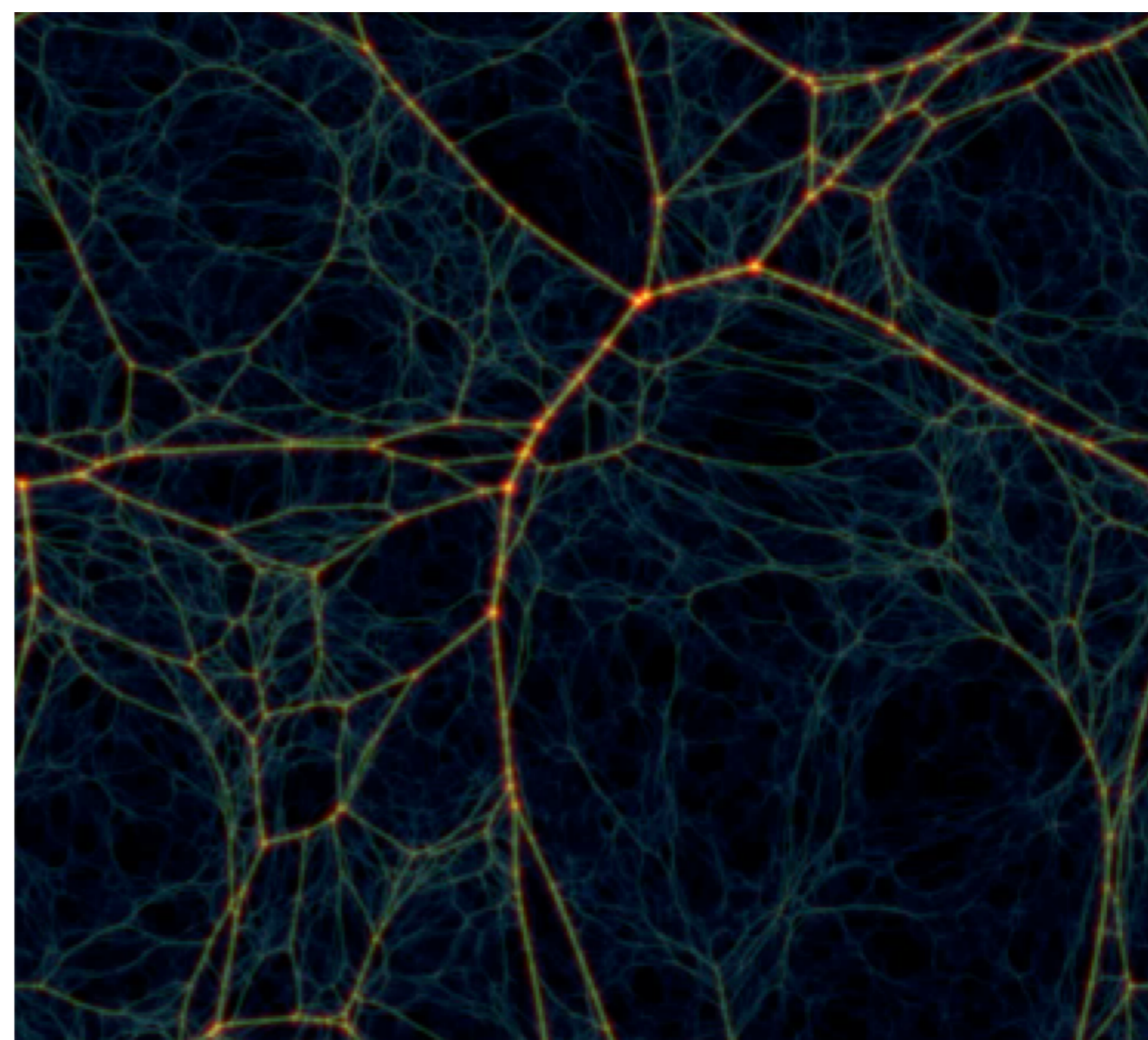
Hidding, Shandarin,  
van de Weygaert 2014



Hidding,  
Shandarin,  
van de Weygaert  
2014



Zel'dovich Approximation



Adhesion Approximation  
Skeleton of structure

Hidding 2010



W E B

or

IRREGULAR HONEYCOMB

Klypin & Shandarin 1983; Shandarin & Klypin 1984  
First demonstration of filaments in 3D N-body  
by plotting density contours

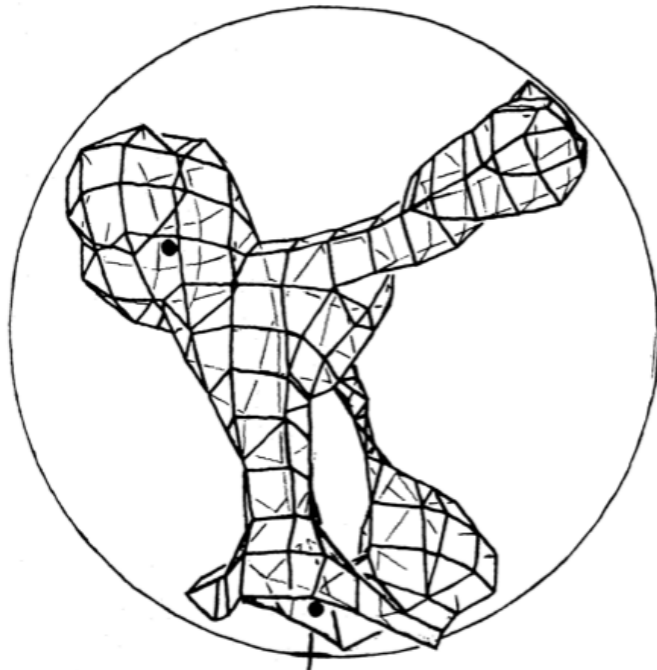


Figure 4. A surface of constant density level is plotted for the same region as that in Fig. 3.

'Cosmic chicken'  
C. Frenk

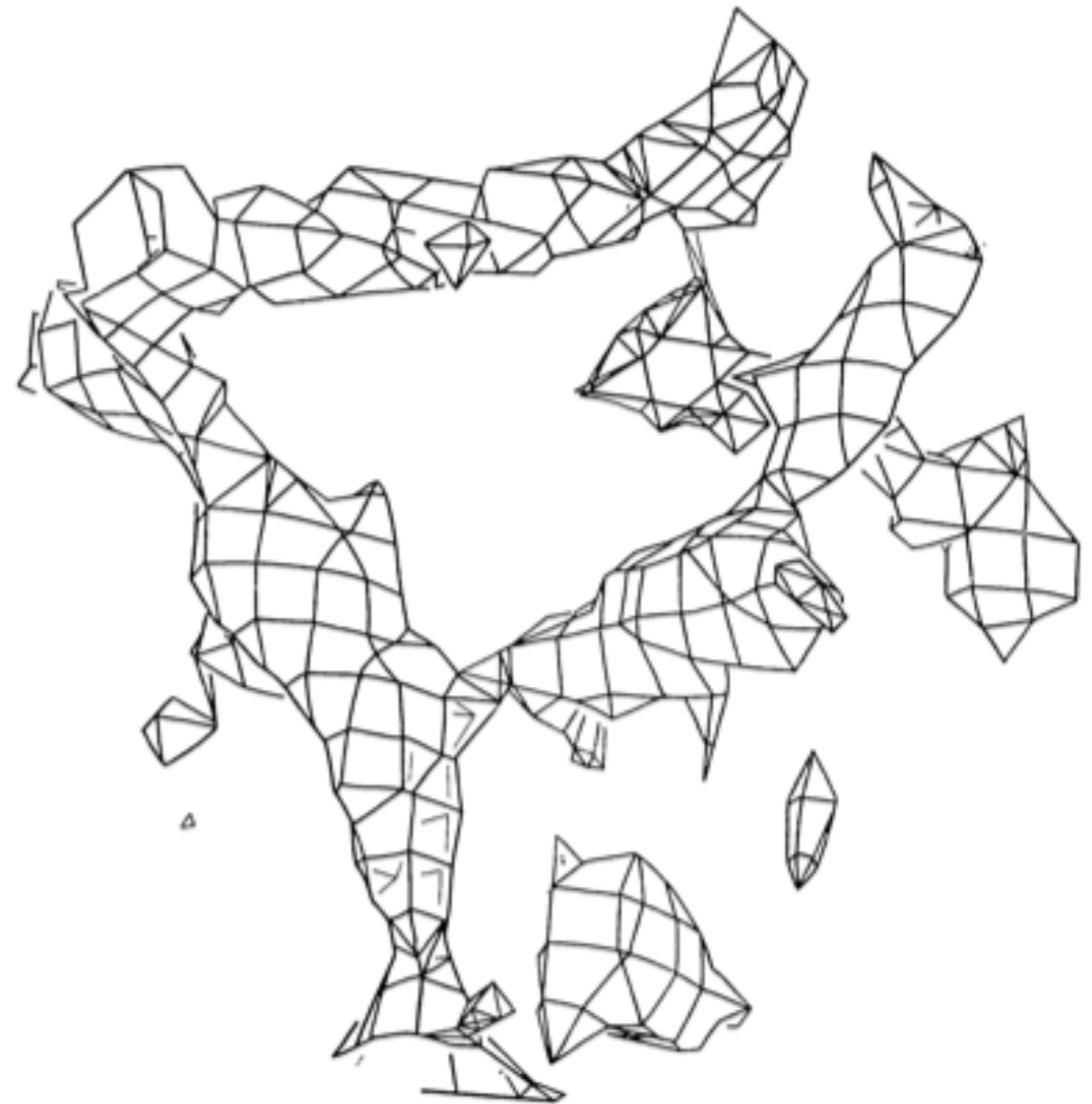


FIG. 1. A typical model isodensity surface,  $\rho = 2.5 \bar{\rho}$ , within a randomly selected sphere of radius  $45 h^{-1}$  Mpc.

“... one may conventionally divide the cluster formation process into three steps. In the first stage, contraction along the axis with the maximum deformation rate would rise to a pancake, but its development would not halt any contraction that may take place along the other two directions.

In the second step some of the pancake material would contract along a second direction, forming curvilinear (not straight, in general) structures or “filaments”, having a finite thickness substantially smaller than their length, as depicted in Fig. 1.

Finally, in the third stage flows along the filaments would produce compact clumps. Thus clusters would be born at points where even during the filament stage contraction had been occurring along all three directions.”

Shandarin and Klypin 1984, Sov. Astron. 28, 491

**This describes the evolution of structure with time.**

« The order in which the physically significant structures arise is basically the inverse of that in the classical pancake picture:

first, high-density peaks,

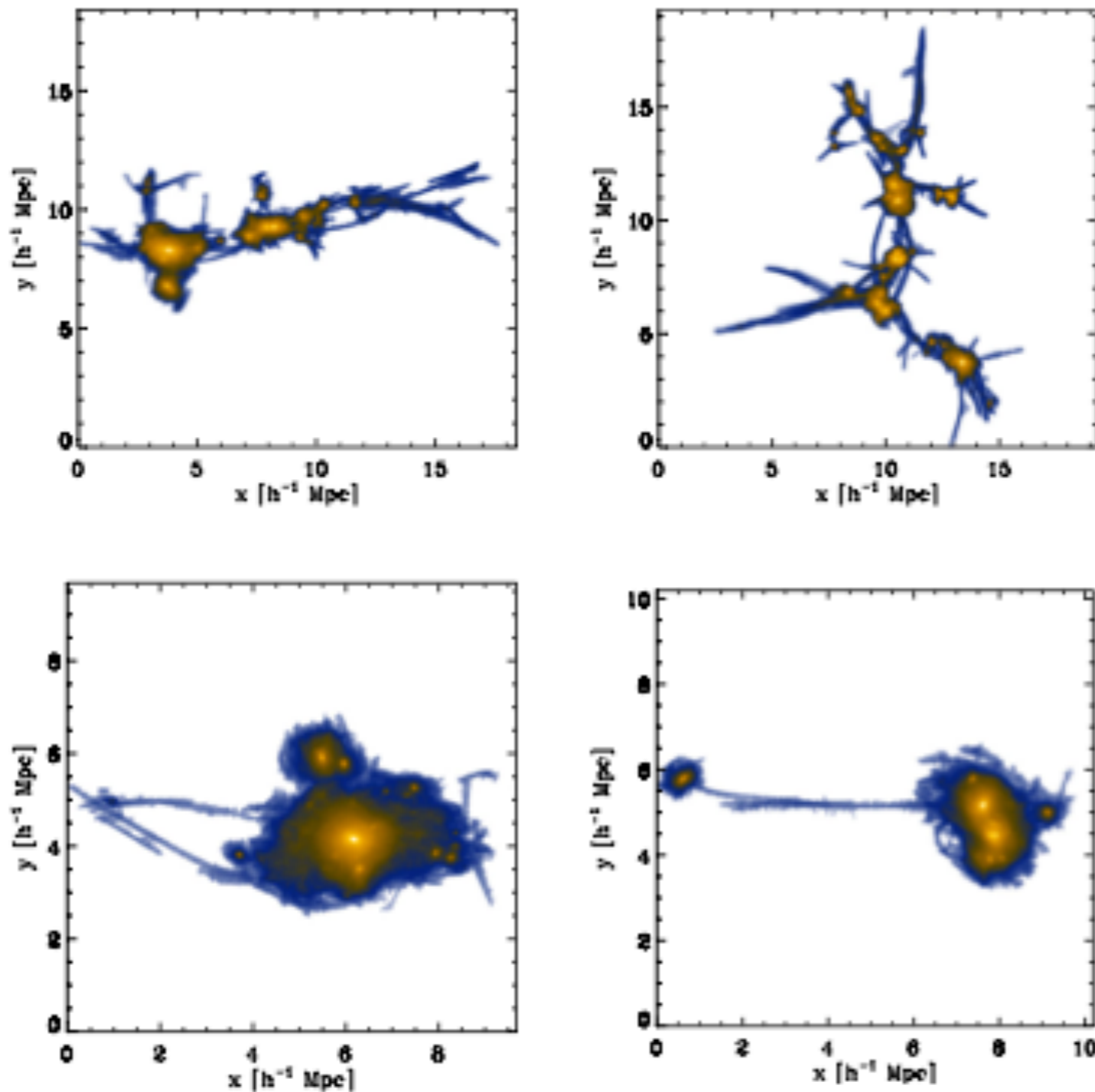
then filaments between them,

and possibly afterwards the walls,  
defined as the rest of the mass between voids.”

Bond, Kofman, Pogosyan 1996, Nature, 380, 603  
(top of page 605)

**This is not a description of the evolution in time, instead this describes unfolding of four generic structures in the excursion set with decreasing threshold.**

**It is universal for all generic fields, no exceptions!**



$z=1$

$z=0$

**Figure 2.** The ‘standard’ linking length  $b = 0.2$  selects large parts of a WDM simulation once the forces are captured accurately enough that filaments do not artificially fragment. Density projections of the particles belonging to the two most massive FoF ‘haloes’ in our WDM T4PM simulation of a 250 eV DM model are shown. Objects at  $z = 1$  (top row) and  $z = 0$  (bottom row) are shown. These haloes have a mass of  $2.7 \times 10^{14}$  and  $1.6 \times 10^{14} h^{-1} M_{\odot}$  ( $z = 1$ ), and of  $6.4 \times 10^{14}$  and  $2.8 \times 10^{14} h^{-1} M_{\odot}$  ( $z = 0$ ).

Angulo et al 2013

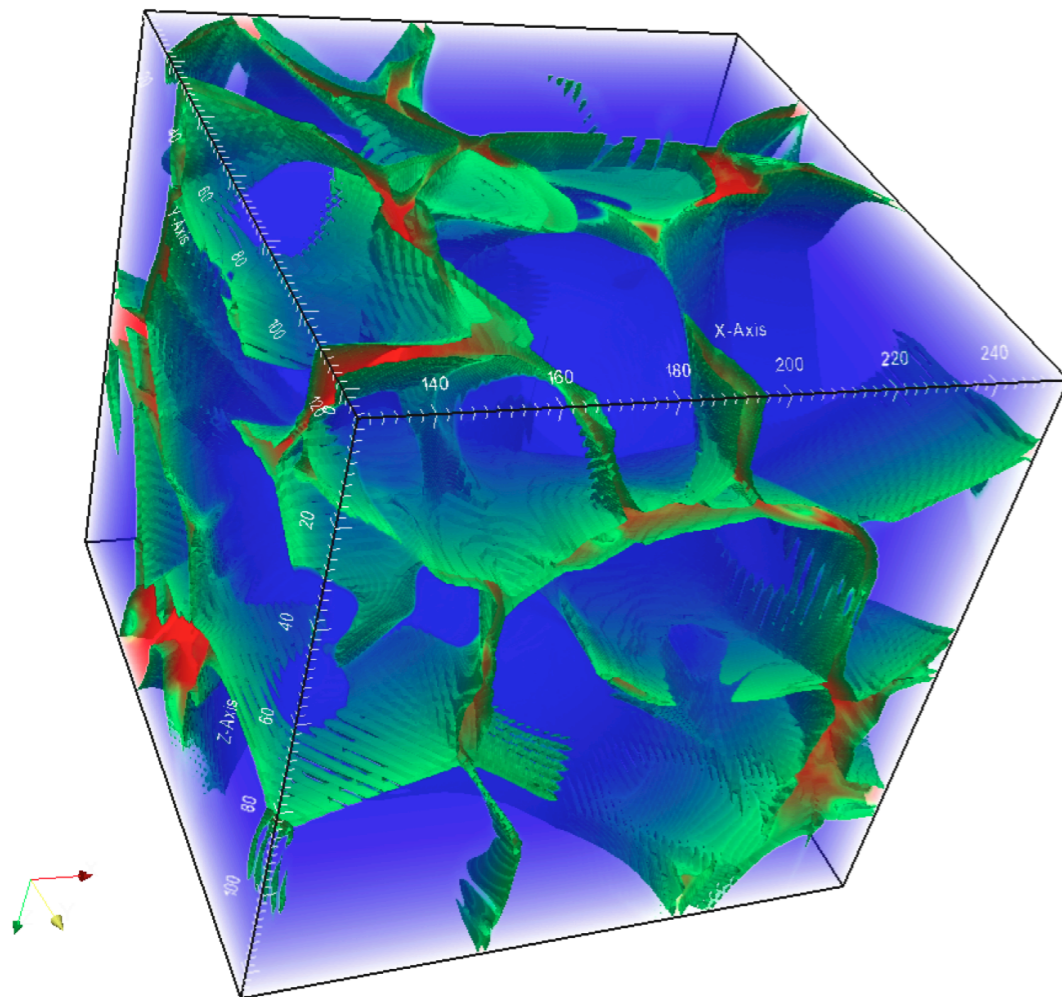
# Three-dimensional numerical model of the formation of large-scale structure in the Universe

A. A. Klypin and S. F. Shandarin *The Keldysh Institute of Applied Mathematics, Academy of Sciences of USSR, Miusskaja Sq. 4, Moscow 125047, USSR*

Received 1982 November 15; in original form 1982 April 28

## Summary:

(5) The regions of high density seem to form a single three-dimensional web structure. However, it is not clear from our simulations whether honeycomb structure arises or not.



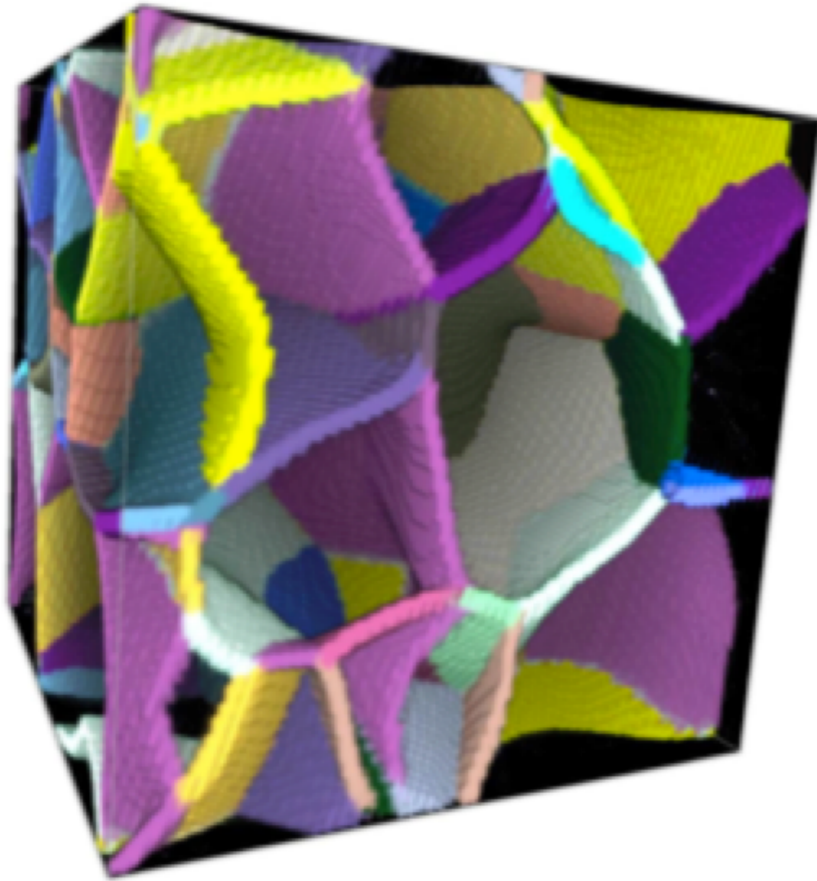
A new method of the analysis of cosmological N-body simulations based on computing the Lagrangian submanifold suggested by Shandarin, et al (2012) Abel et al (2012) allowed to demonstrate that the irregular honeycomb structure is formed in Dark Matter distribution.

# Zeldovich's pancakes in 3D

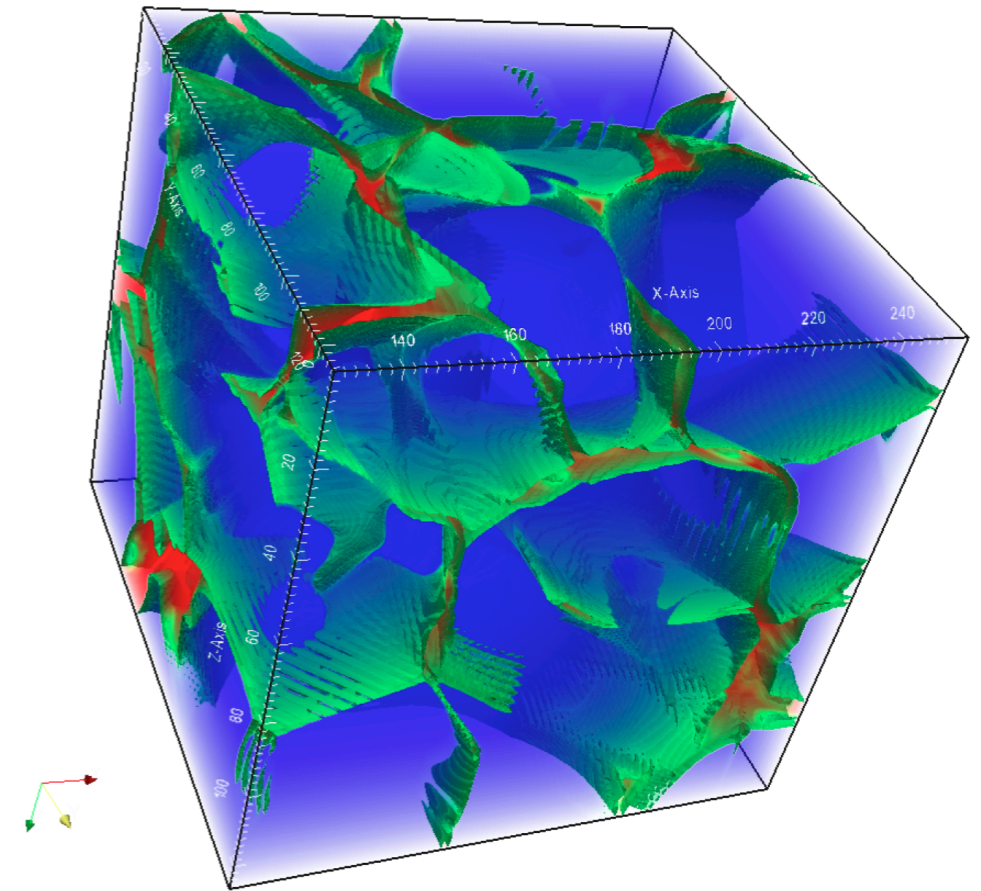
N-body Simulations

Zeldovich Approximation

(note: initial random number are different)



Sausbie 2010  
based on computing  
Morse-Smale  
complex



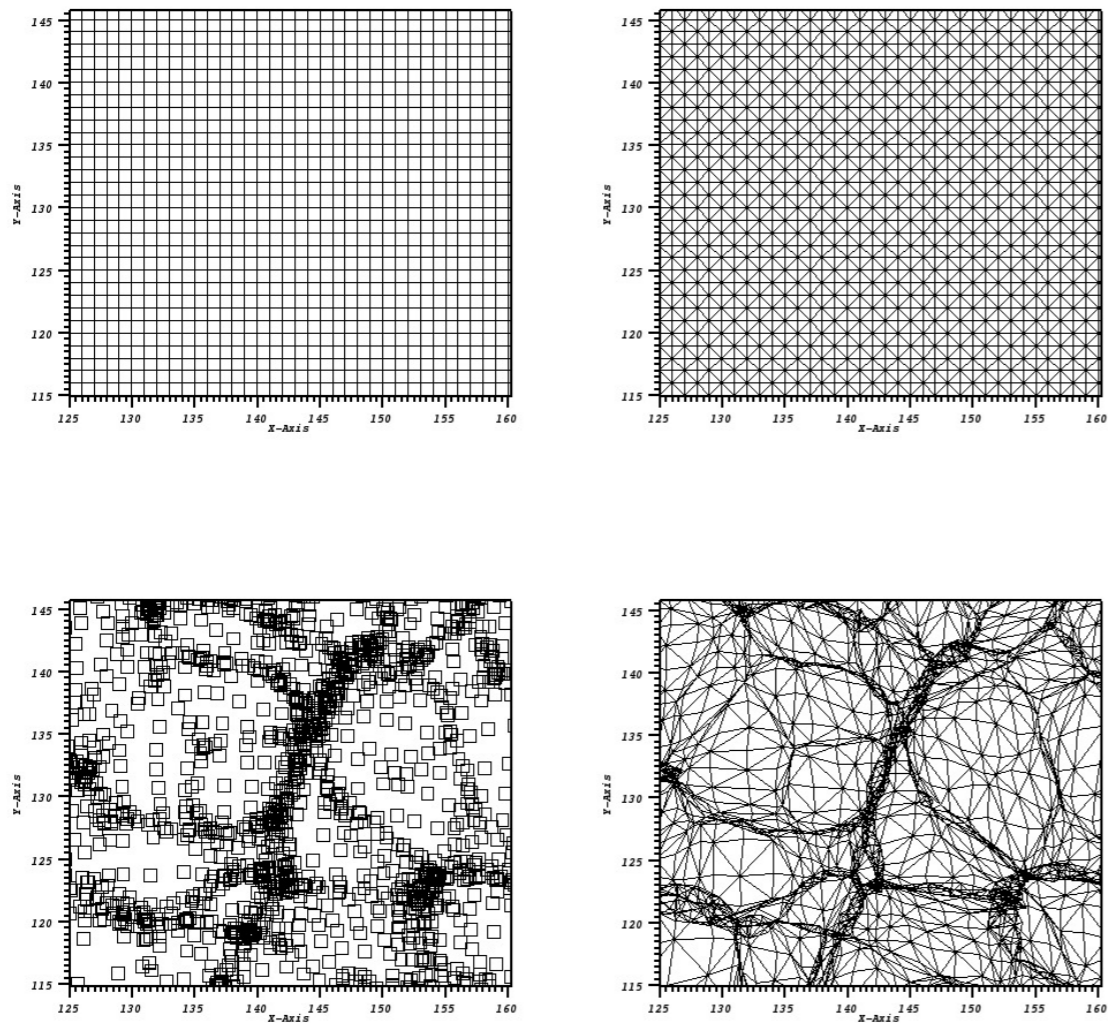
Hidding, Shandarin, van de Weygaert 2014

Multistream flows



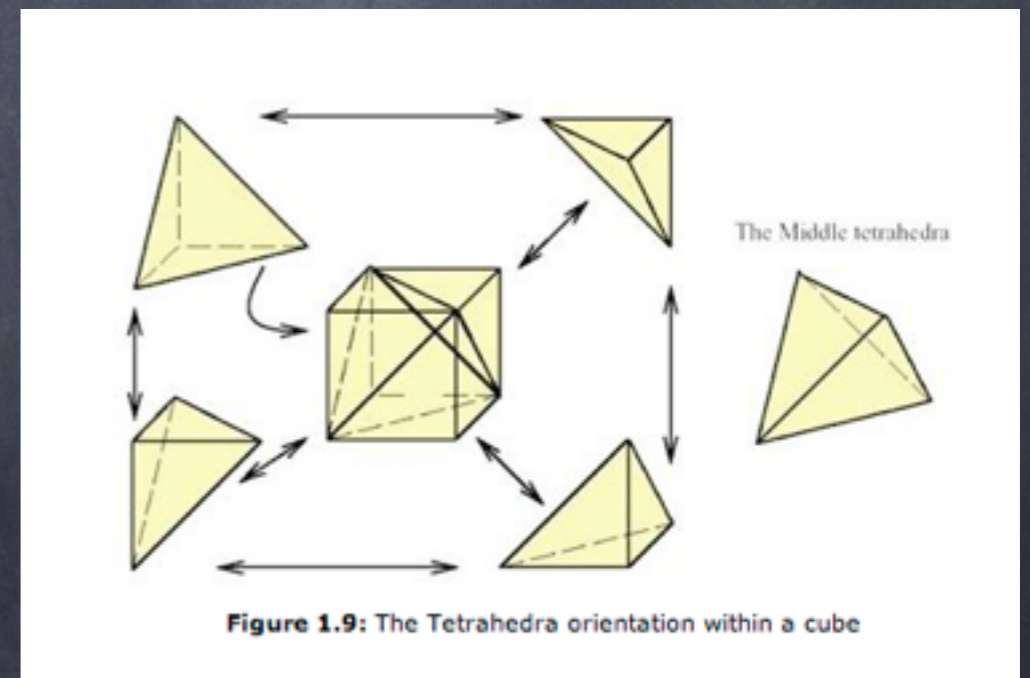
# 2D example of triangulation of Lagrangian submanifold

triangulation of initial (Lagrangian) plane



projection of Lagrangian submanifold on Eulerian space

# Decomposition of a cube in tetrahedra in 3D



# Why detecting Zeldovich's pancake in 3D simulations took so long?

Shandarin, Habib, Heitman 2012

see also

Abel, Hahn, Keuhler 2012

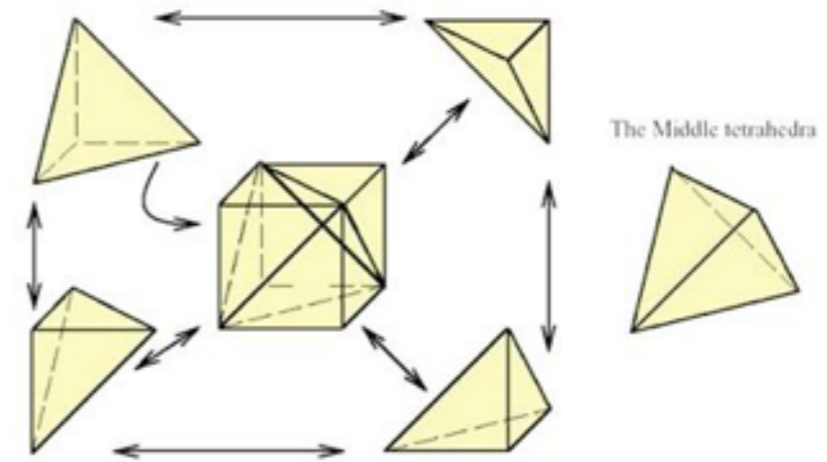
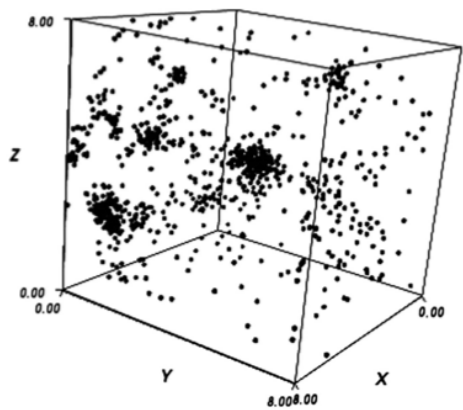
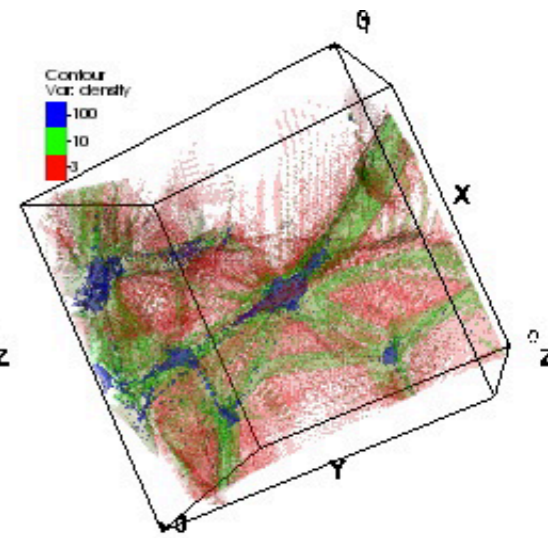
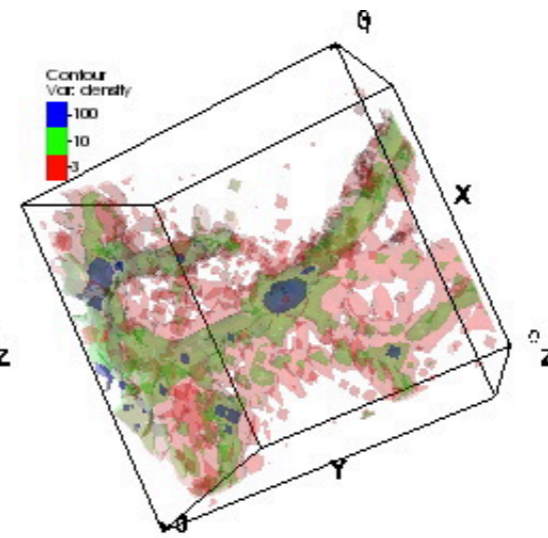
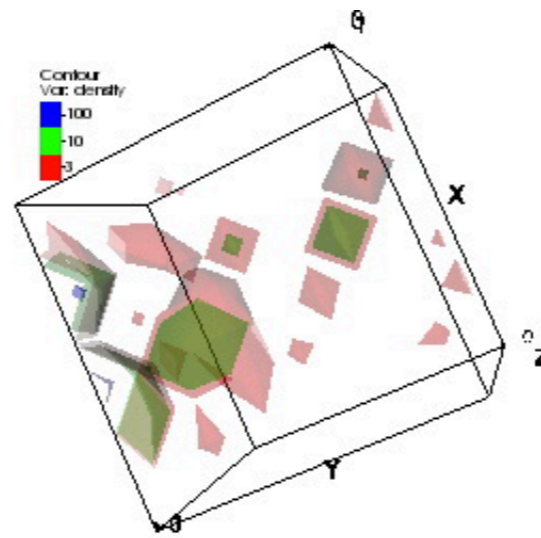


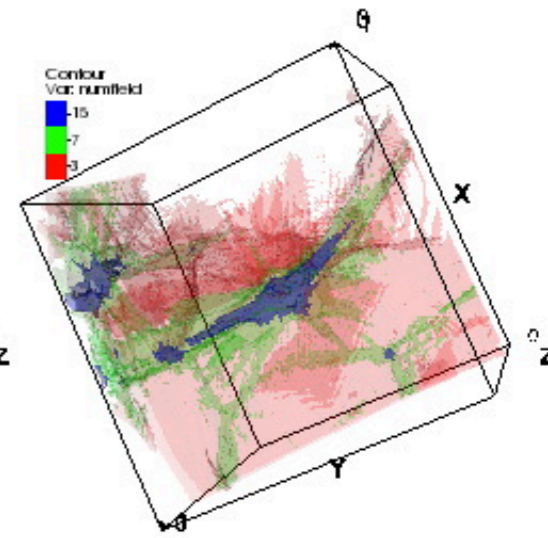
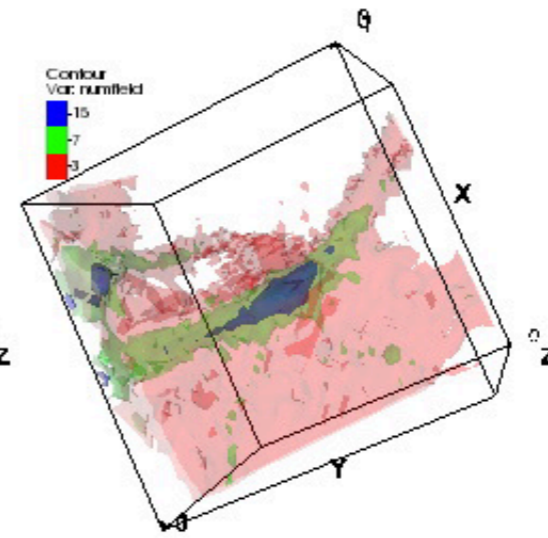
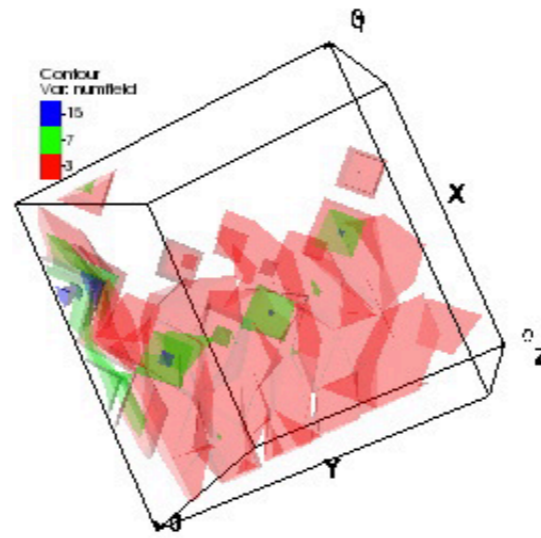
Figure 1.9: The Tetrahedra orientation within a cube



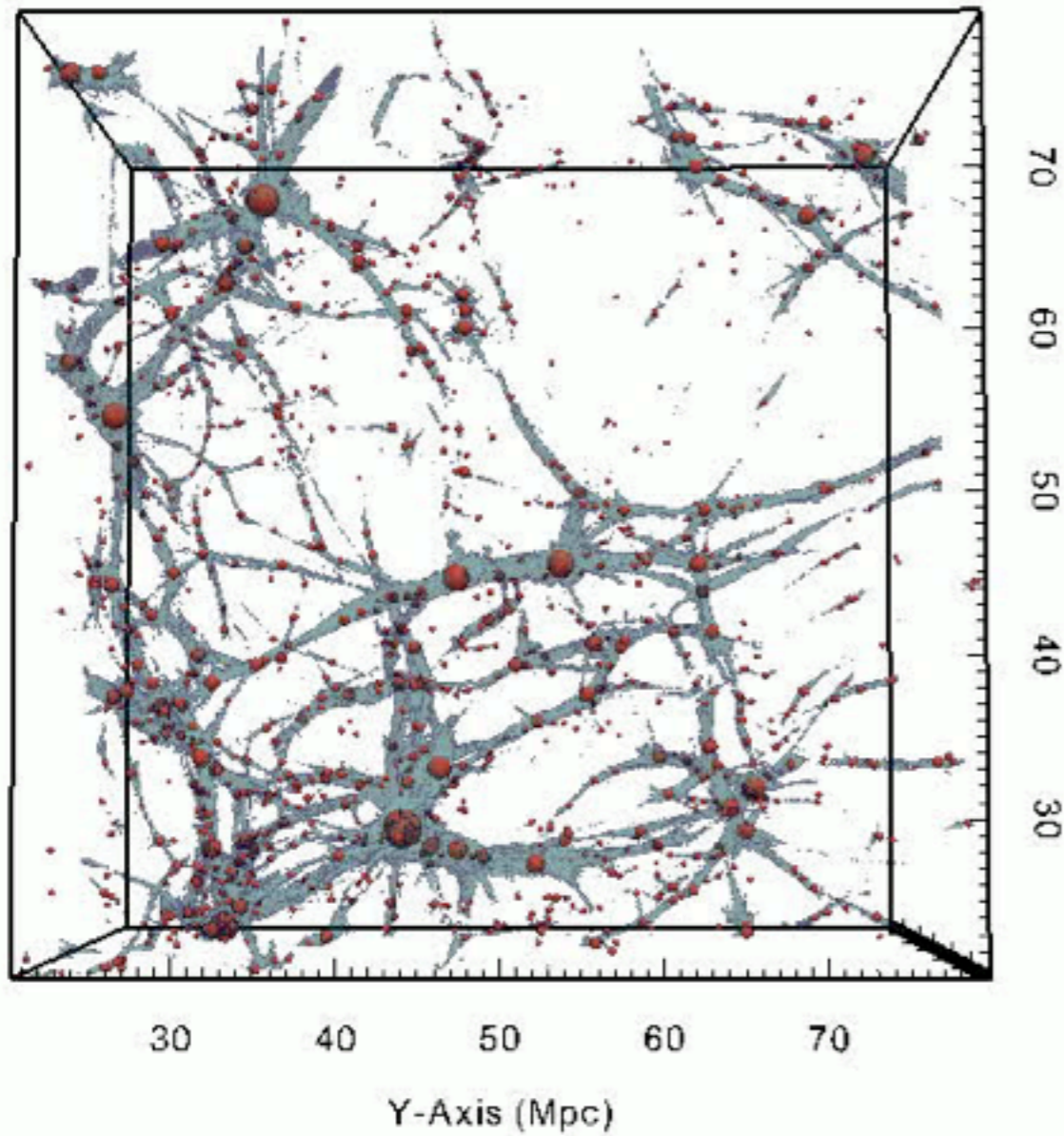
DENSITY  
FIELD



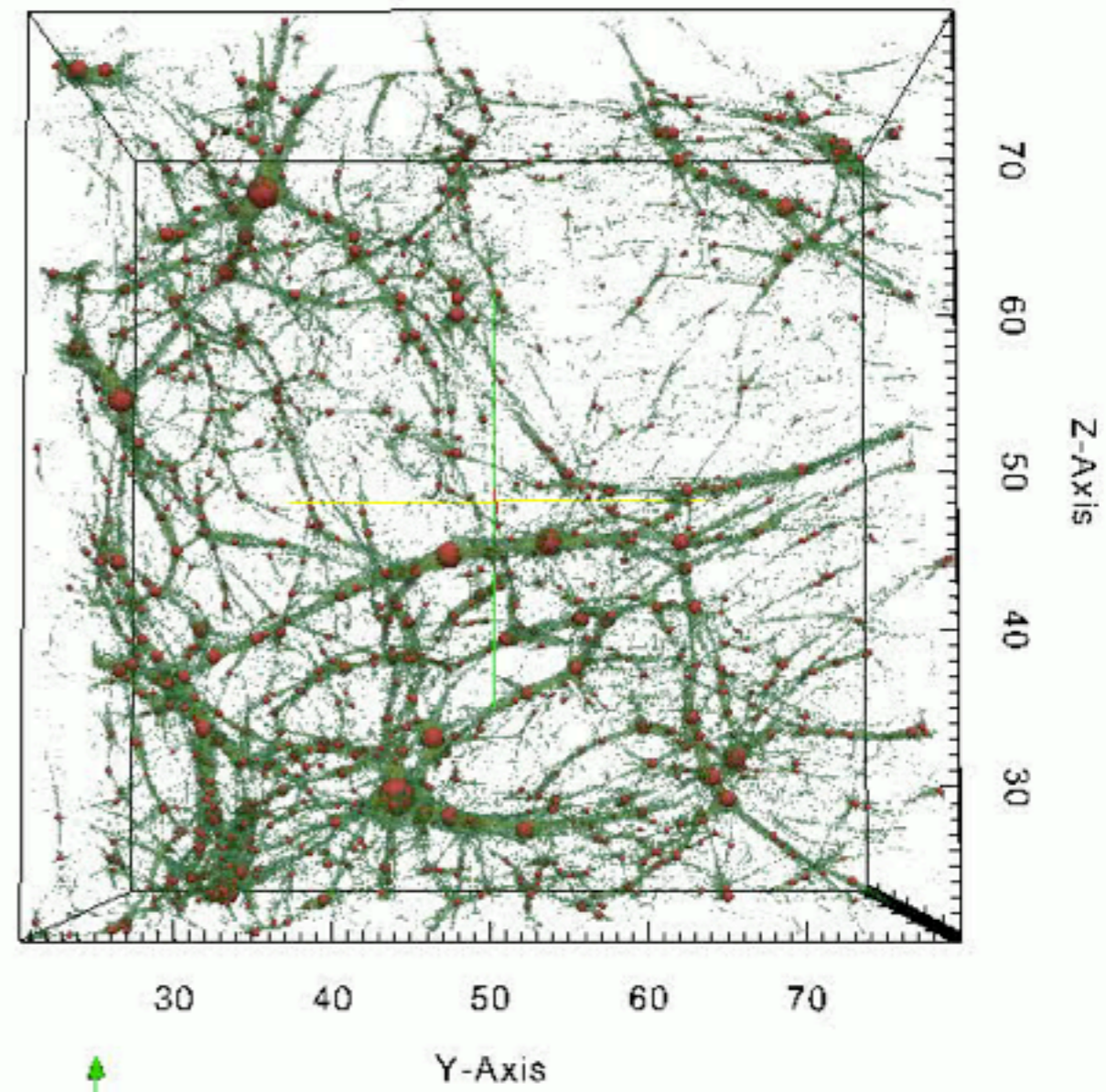
MULTI-STREAM  
FIELD



All halos are embedded in filaments



streams

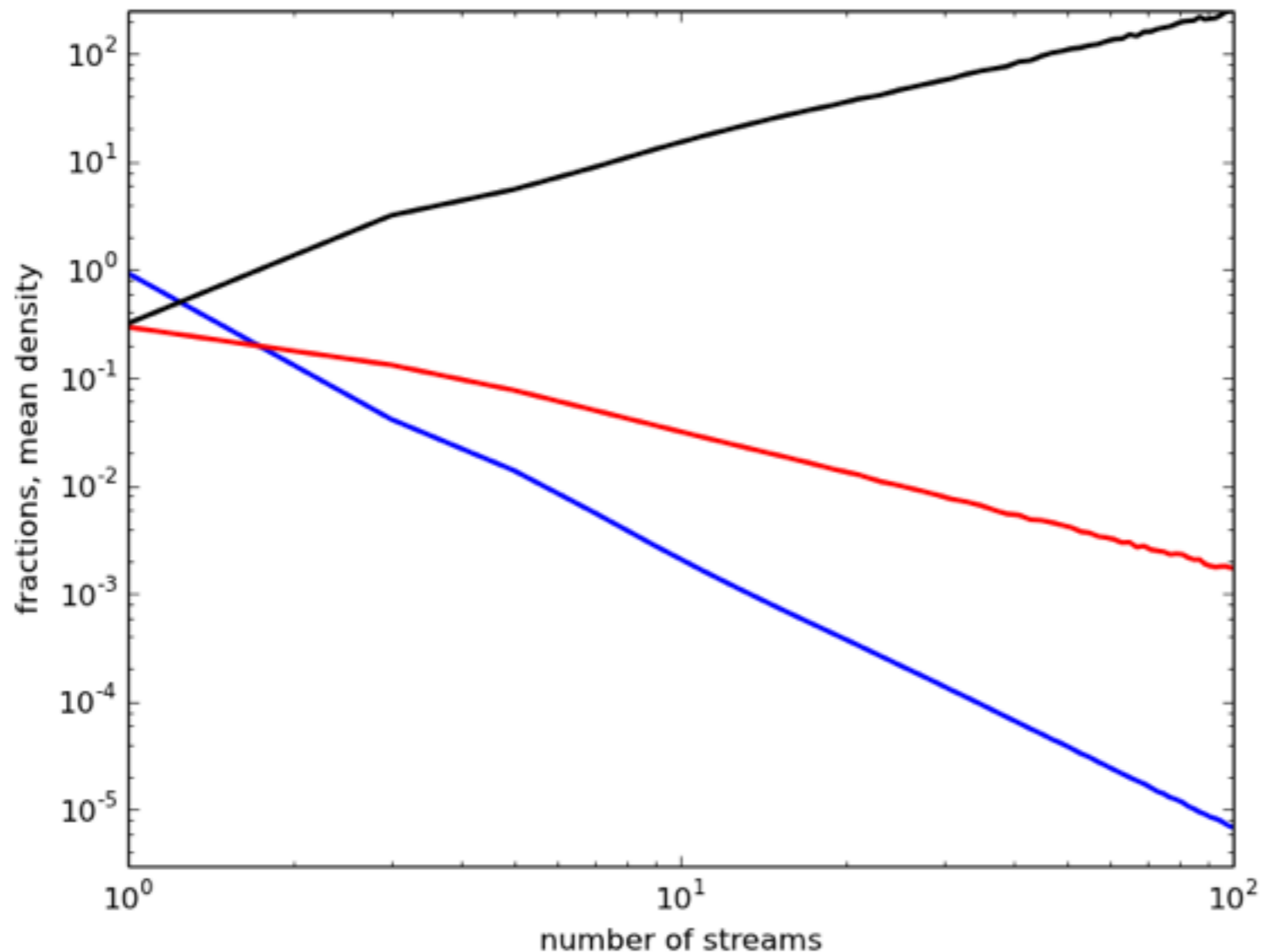


density

**Fraction of volume with given number of streams - blue**

**Fraction of mass with given number of streams - red**

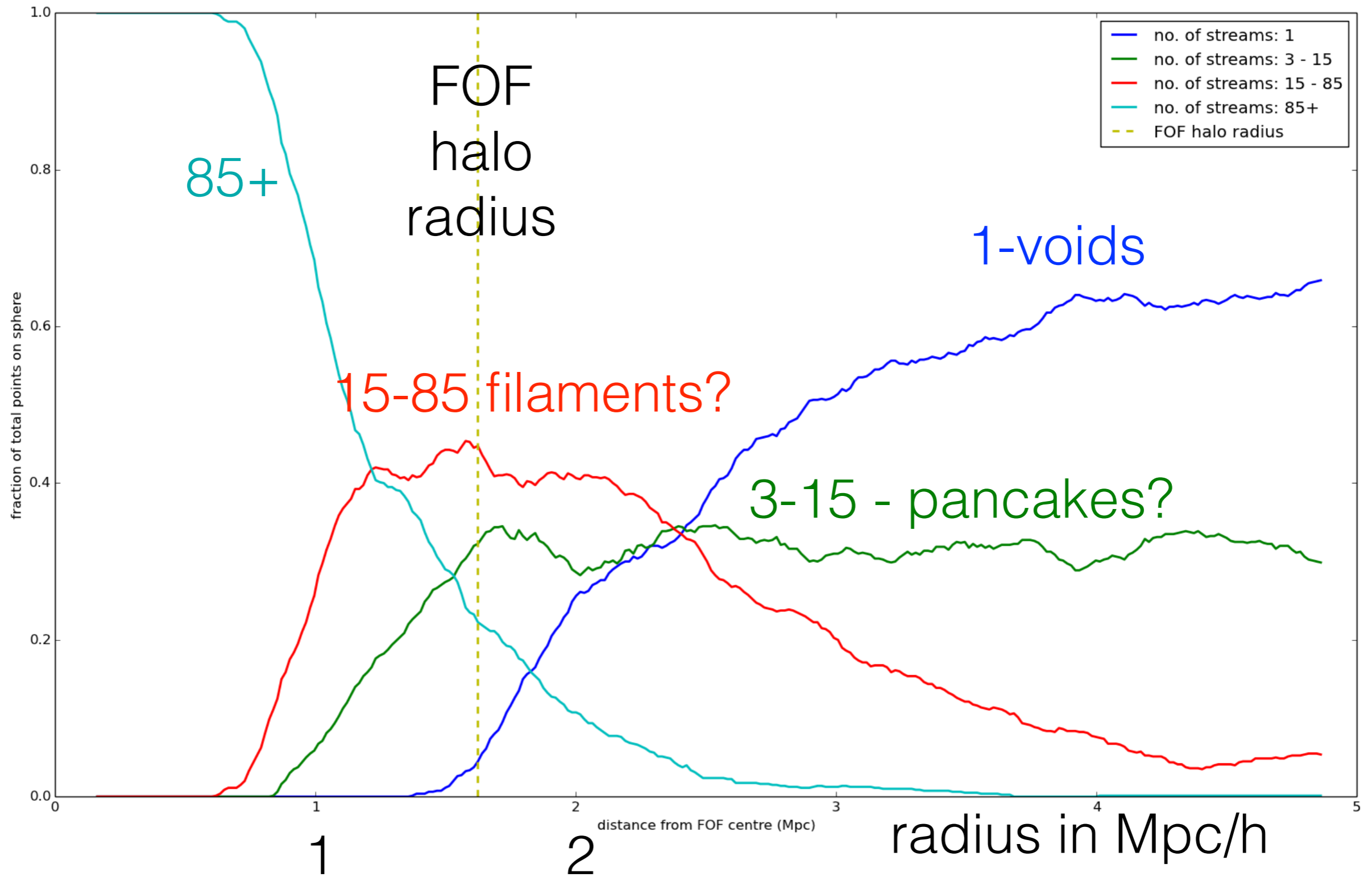
**Mean density in cells with given number of streams - black**



$$\begin{aligned} \text{den} &\geq 200 * (\text{mean den}) \\ v\_frac &= 0.03\% \\ m\_frac &= 16\% \\ \langle \text{den} \rangle &= m\_frac / v\_frac \\ &= 492 \end{aligned}$$

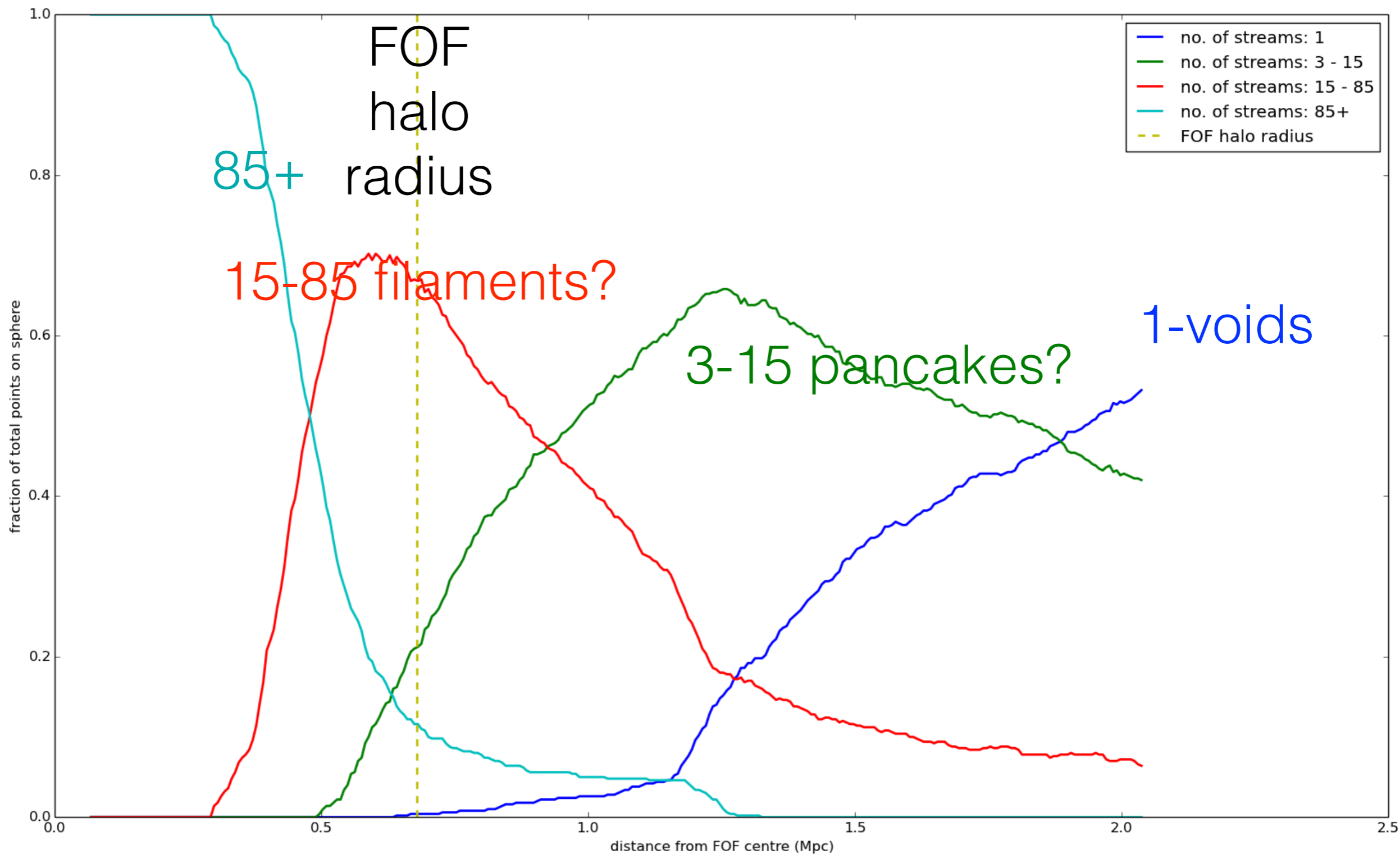
# Fraction of points with a given number of streams as a function of radius

$M=6e14 M_{\text{sun}}$



# Fraction of points with a given number of streams as a function of radius

$M=4e13 M_{\text{sun}}$



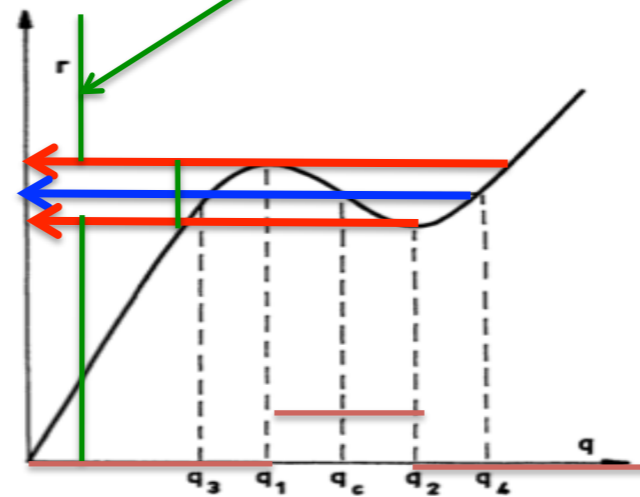
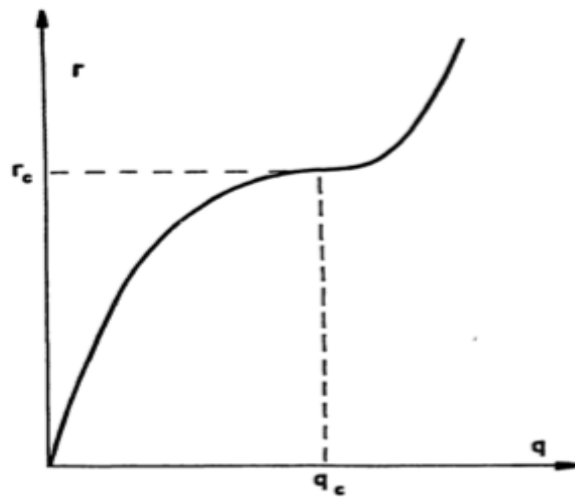
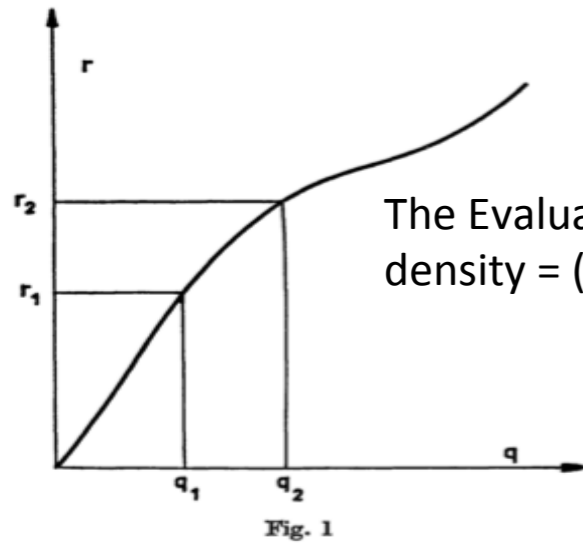
Lagrangian submanifold

Number of streams field

Flip-Flop field

# Lagrangian Submanifold (LS) is N-dim surface in 2N-dim space

Zel'dovich 1970

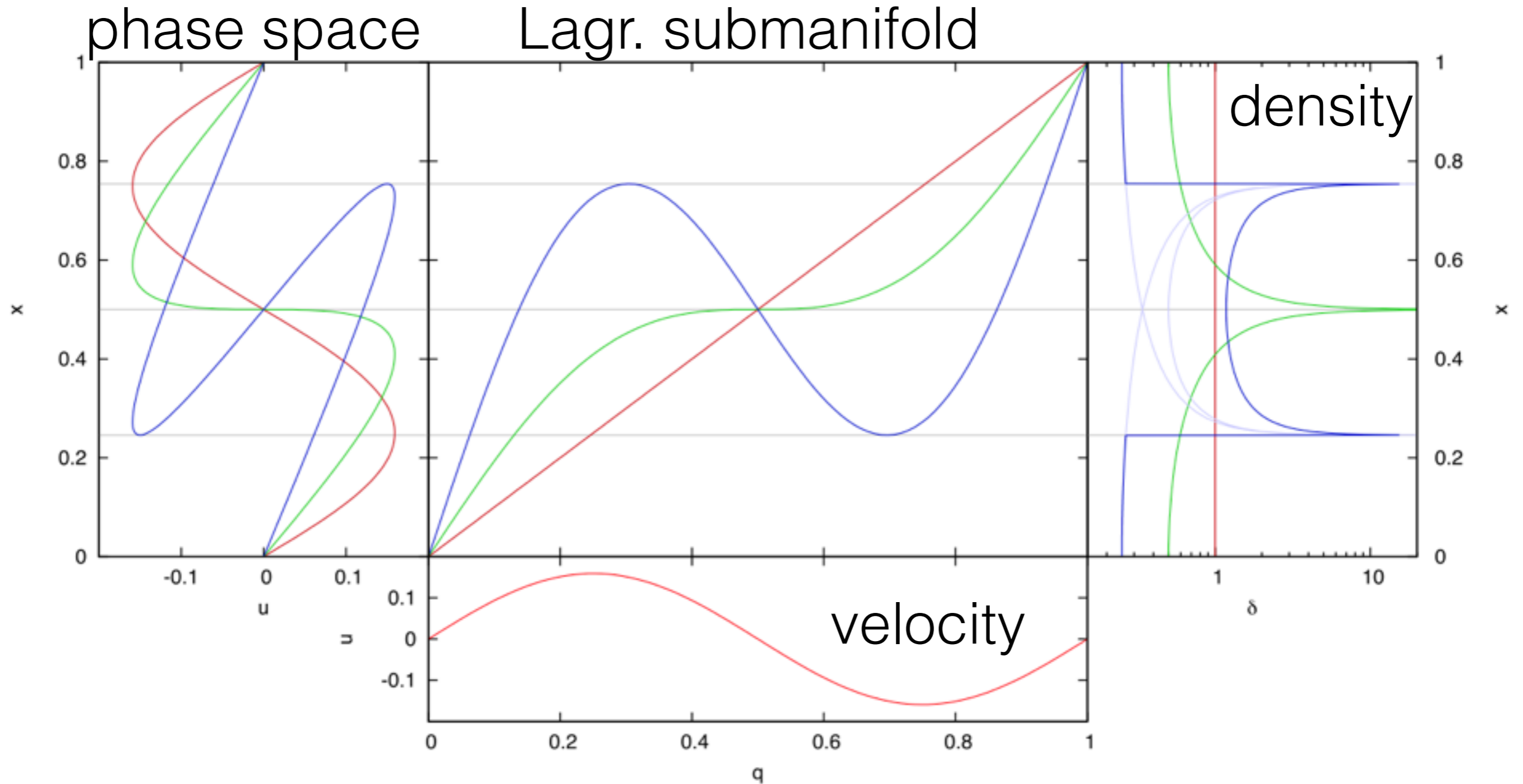


Multi-Stream Field

Flip-Flop Field



# Phase space and Lagrangian submanifold in 1D



Three stages are shown: **initial - red**  
**critical - green**  
**final - blue**

# Halo in 1D N-body simulation

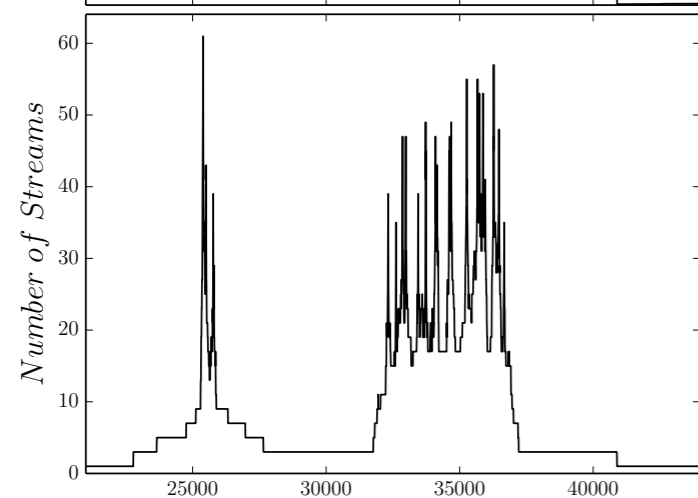
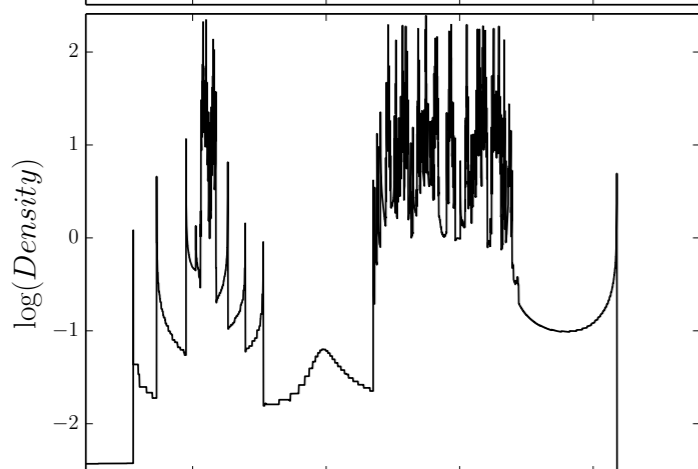
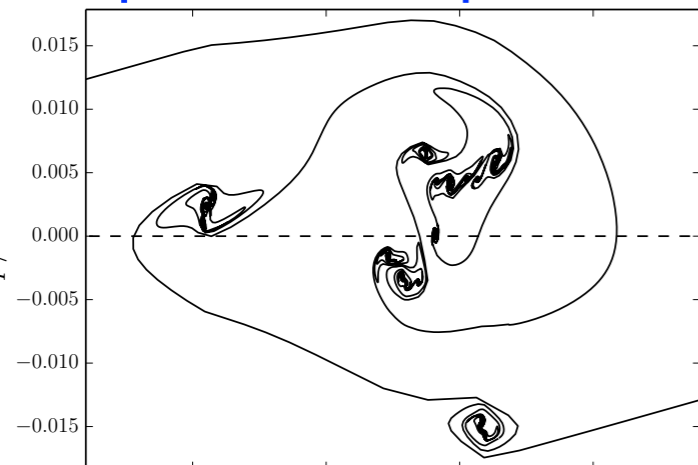
Shandarin,  
Medvedev,  
Hidding 2014  
in preparation

phase space

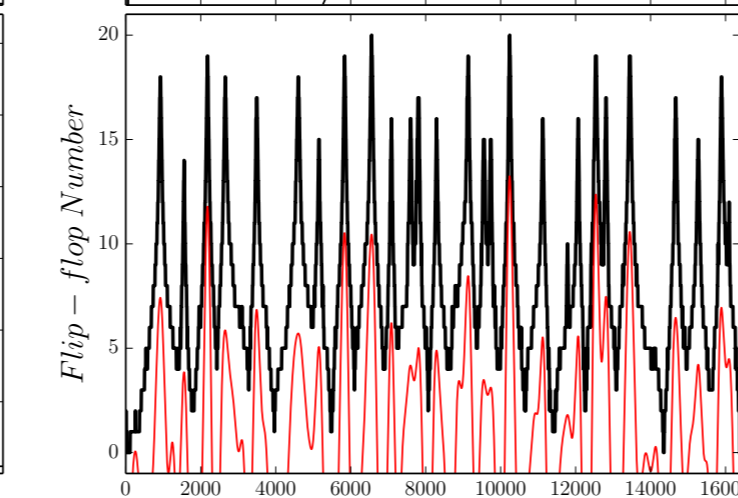
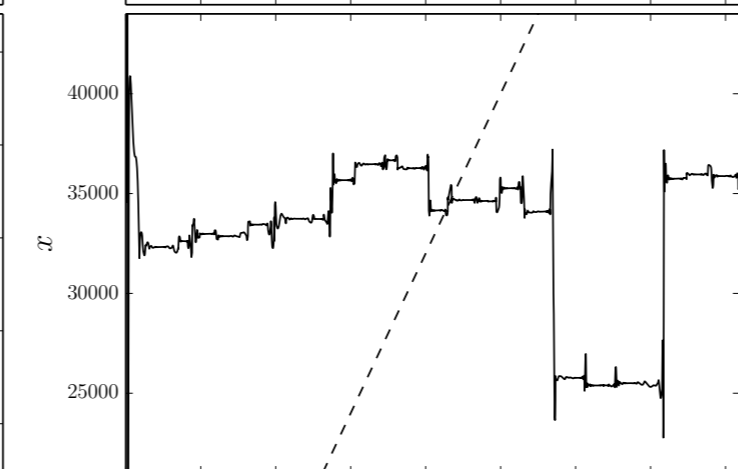
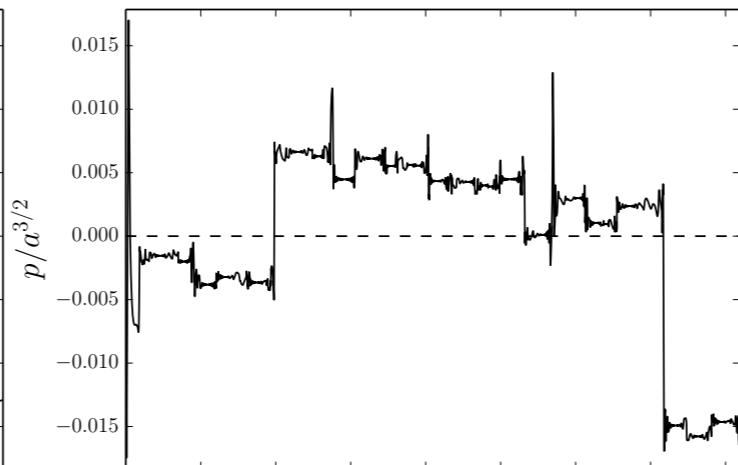
velocity

density

# streams



$x$



$q$

Phase space sheet :  
multivalued, non metric

Lagr. submanifold:  
single valued, metric

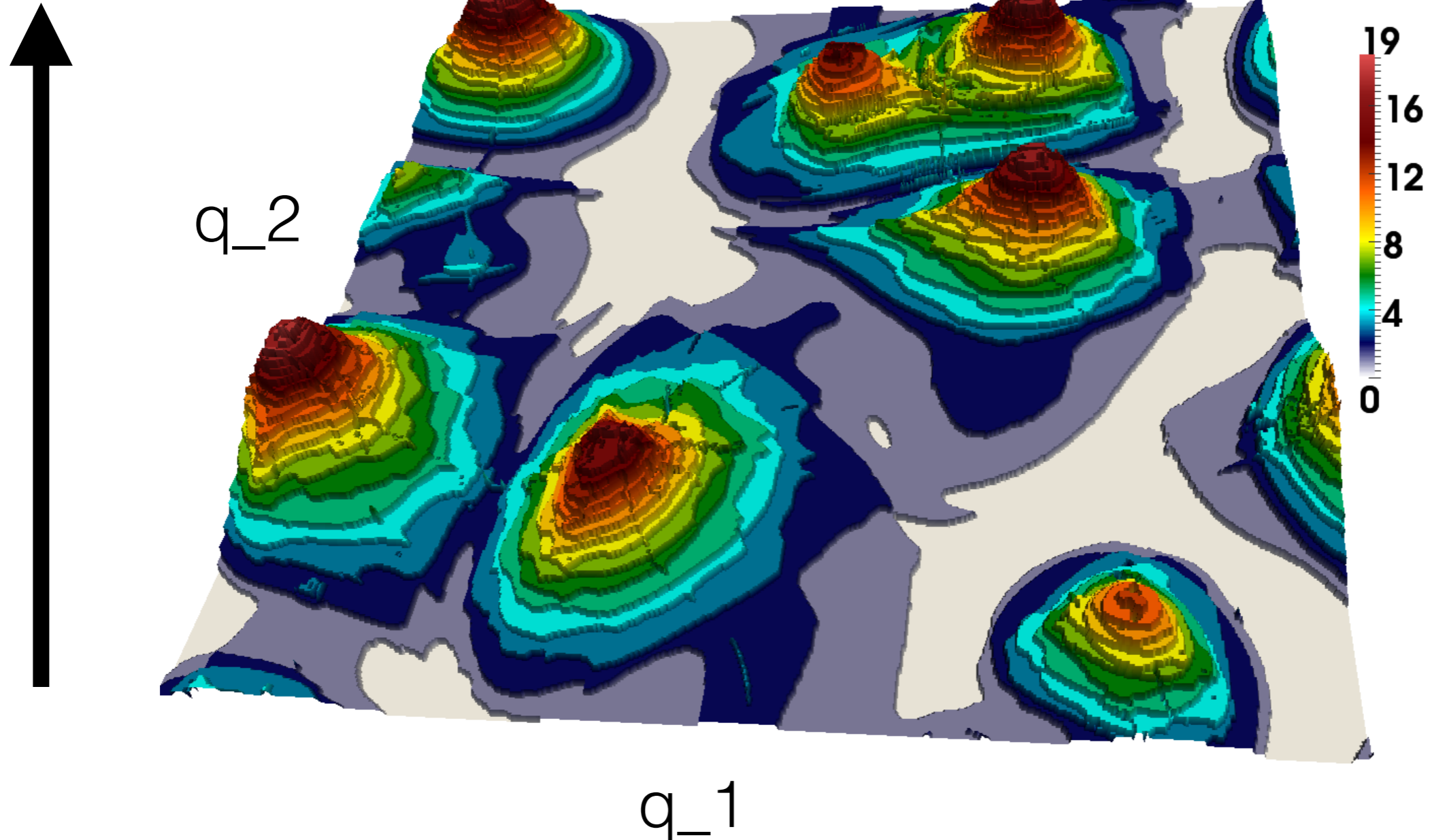
Lagrangian  
submanifold:  
 $x = x(q)$

Flip-flop  
number

2D

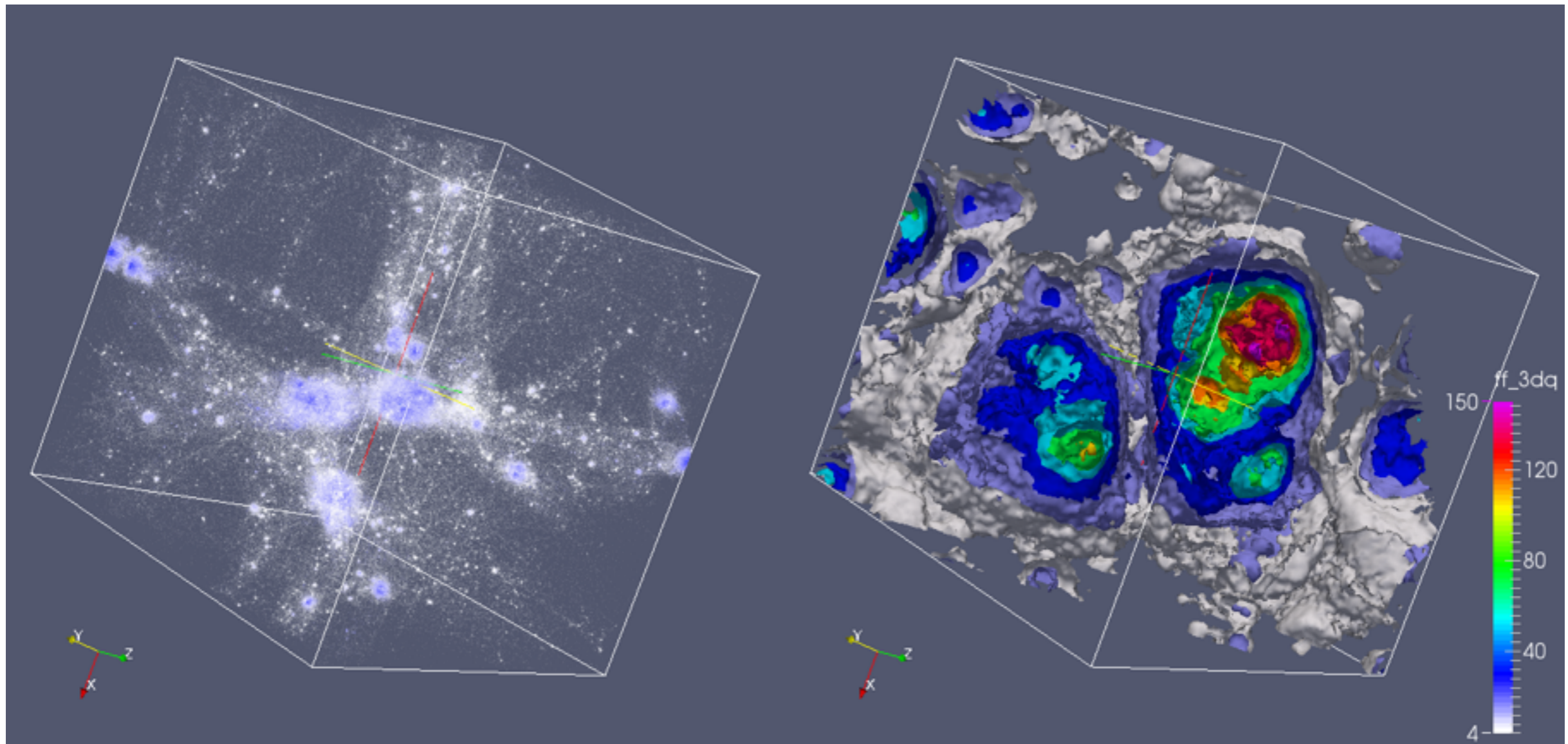
Flip-Flop Field in Lagrangian space in  $512^2$  simulation  
(smooth initial conditions)

# flip-flops

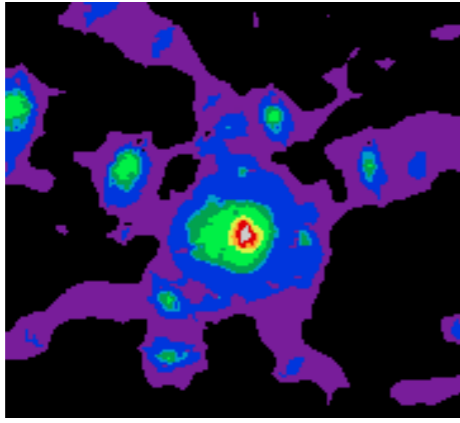


Shandarin, Medvedev, Hidding 2014, in preparation

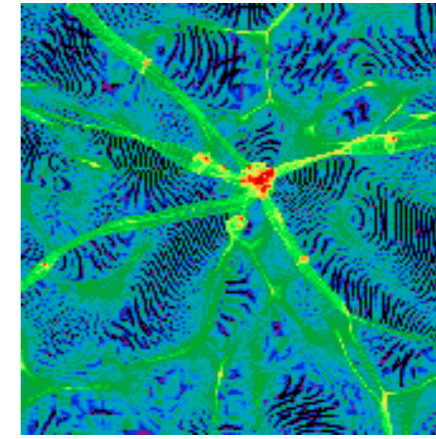
3D simulation in LCDM model: box size 1 Mpc/h,  $N_p = 256^3$



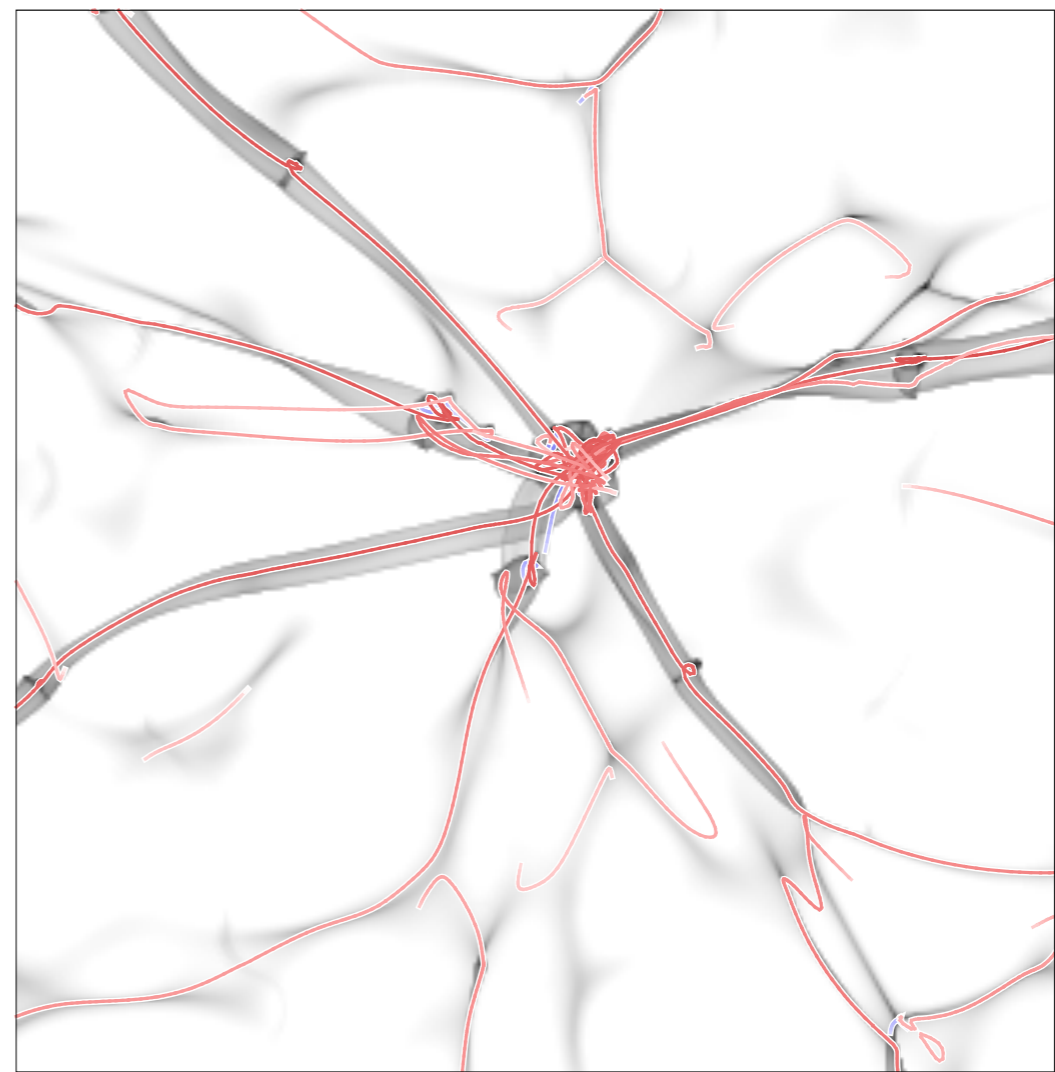
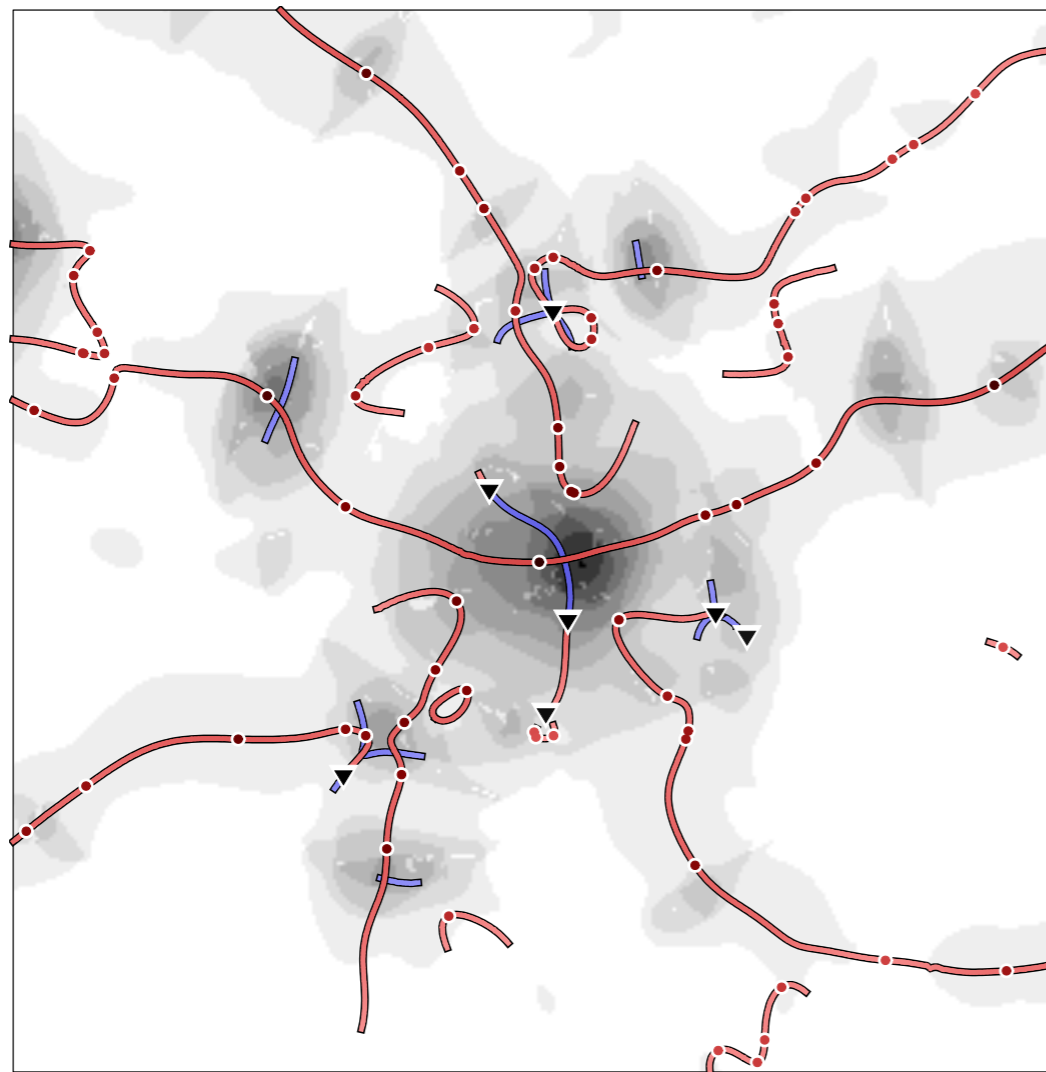
# Lagrangian Skeleton and Flip-Flop Field



Lagrangian



Eulerian



# Summary

The **Zel'dovich approximation** provides the necessary concepts and language for describing complex geometry and dynamics of DM structures.

The sequence of formation of generic elements in DM structure is

- 1) pancakes,
- 2) filaments on the crossings of pancakes,
- 3) halos on the crossing of filaments

Halos are embedded in filaments and filaments are embedded in pancakes.

The skeleton of the Cosmic Web derived from Zel'dovich approximation allows to trace the dynamical evolution of the Cosmic Web and provides quantitative characteristics of the web.

# Summary

New fields: **number of streams** and **flip-flop** fields reveal new properties of the cosmic web. They are easy to compute from standard cosmological simulations.

**Number of streams field** as a function of Eulerian coordinates allows to set **physical limit** on the total volume and mass of the voids:

for LCDM **volume fraction** is  $\sim 93\%$  and **mass fraction** 24%

The **number of flip-flops** as a function of Lagrangian coordinate stores the information about the substructure of DM halos.