

MODELLING COSMIC VOID STATISTICS

NICO HAMAUS

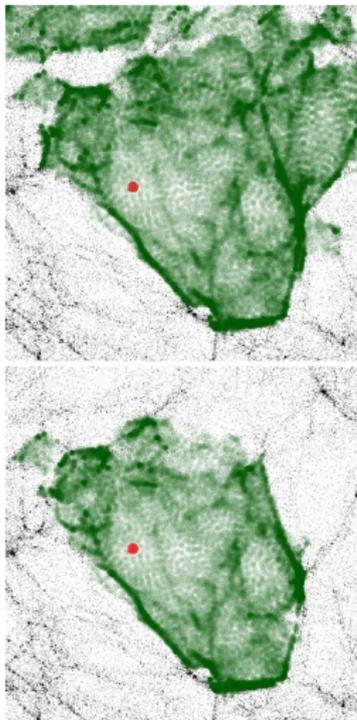
in collaboration with

Benjamin Wandelt, Paul Sutter, Guilhem Lavaux
Alice Pisani, Florent Leclercq, Jens Jasche, Michael Warren

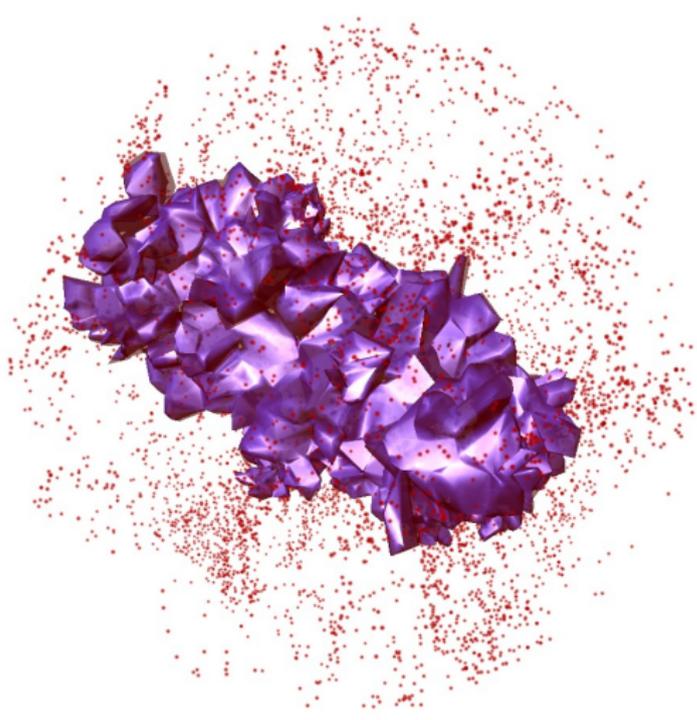


- 1 Introduction
- 2 Density profile [arXiv:1309.5087, arXiv:1403.5499]
- 3 2-point statistics in redshift space [in prep.]
- 4 Outlook

Voids in the cosmic web

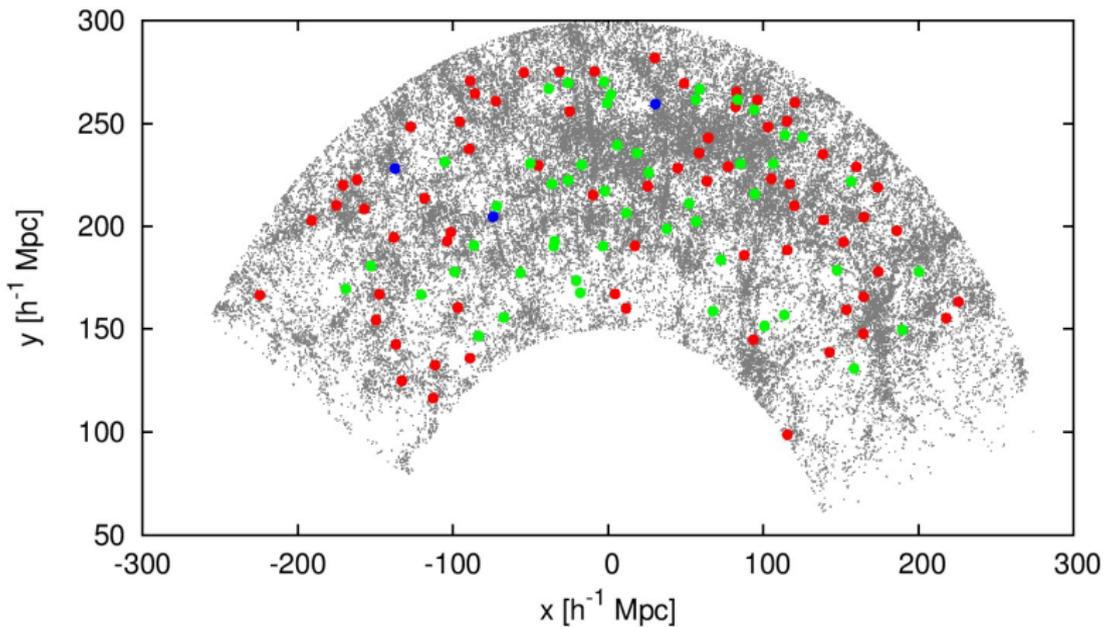


Zobov: Neyrinck (2008)



Sutter, Lavaux, Wandelt, Weinberg (2012)

SDSS voids



$R = 5-15 h^{-1} \text{ Mpc}$
 $R = 15-25 h^{-1} \text{ Mpc}$

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•

$R = 25-45 h^{-1} \text{ Mpc}$

•

Sutter et al. (2012)

Void profile

Estimate density profile by “stacking” tracer particles around void centers

$$\rho_v(r) = \frac{3}{4\pi} \sum_i \frac{m_i(\mathbf{r}_i)}{(r + \delta r)^3 - (r - \delta r)^3}$$

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Estimator for velocity profile

$$v_v(r) = \frac{1}{N(r)} \sum_i \mathbf{v}_i(\mathbf{r}_i) \cdot \frac{\mathbf{r}_i}{r_i} V_c(\mathbf{r}_i)$$

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With linear theory

$$v_v(r) = -\frac{1}{3} \Omega_m^\gamma H r \Delta(r) \qquad \Delta(r) = \frac{3}{r^3} \int_0^r \left(\frac{\rho_v(q)}{\bar{\rho}} - 1 \right) q^2 dq$$

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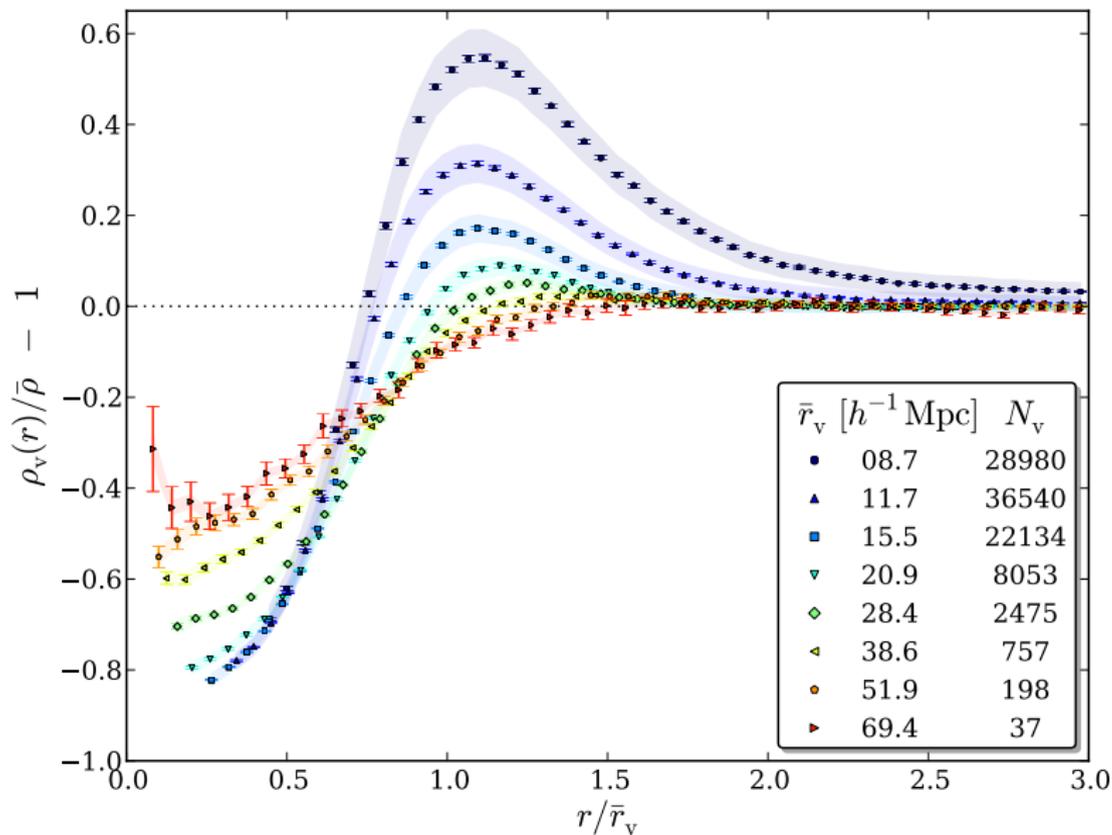
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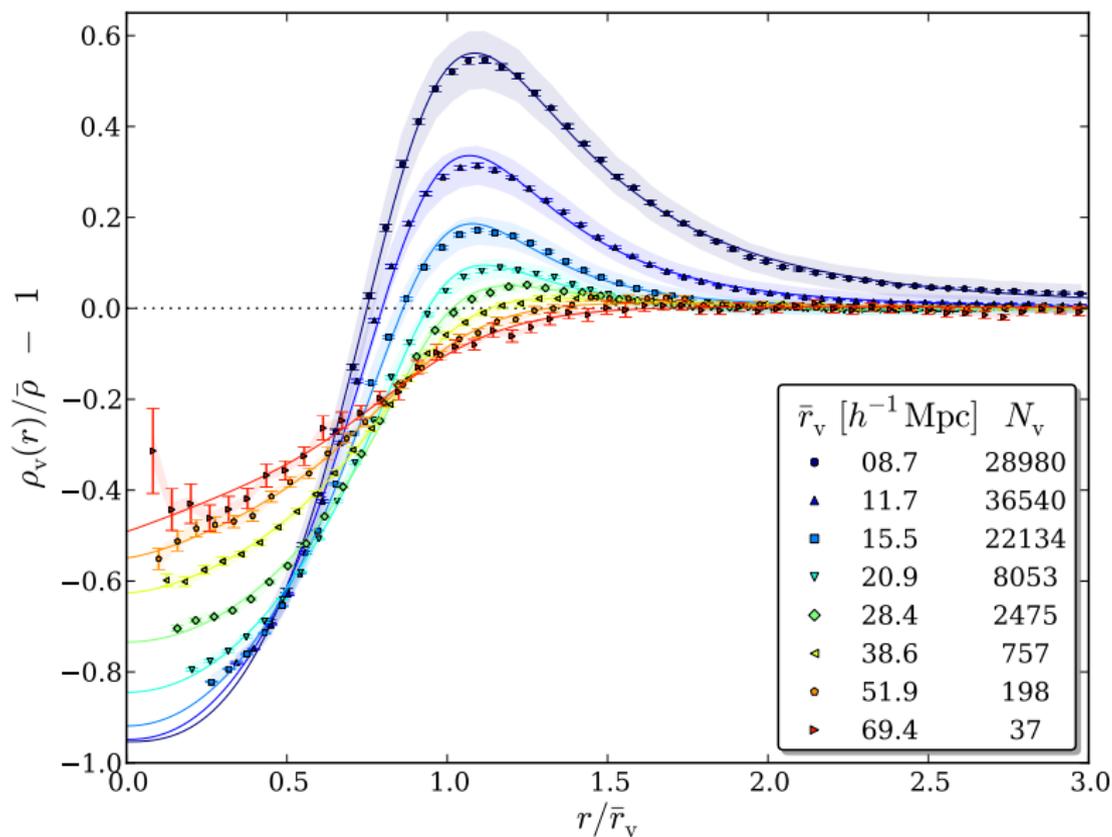
Empirical best-fit model (4 parameters)

$$\frac{\rho_v(r)}{\bar{\rho}} - 1 = \delta_c \frac{1 - (r/r_s)^\alpha}{1 + (r/r_v)^\beta}$$

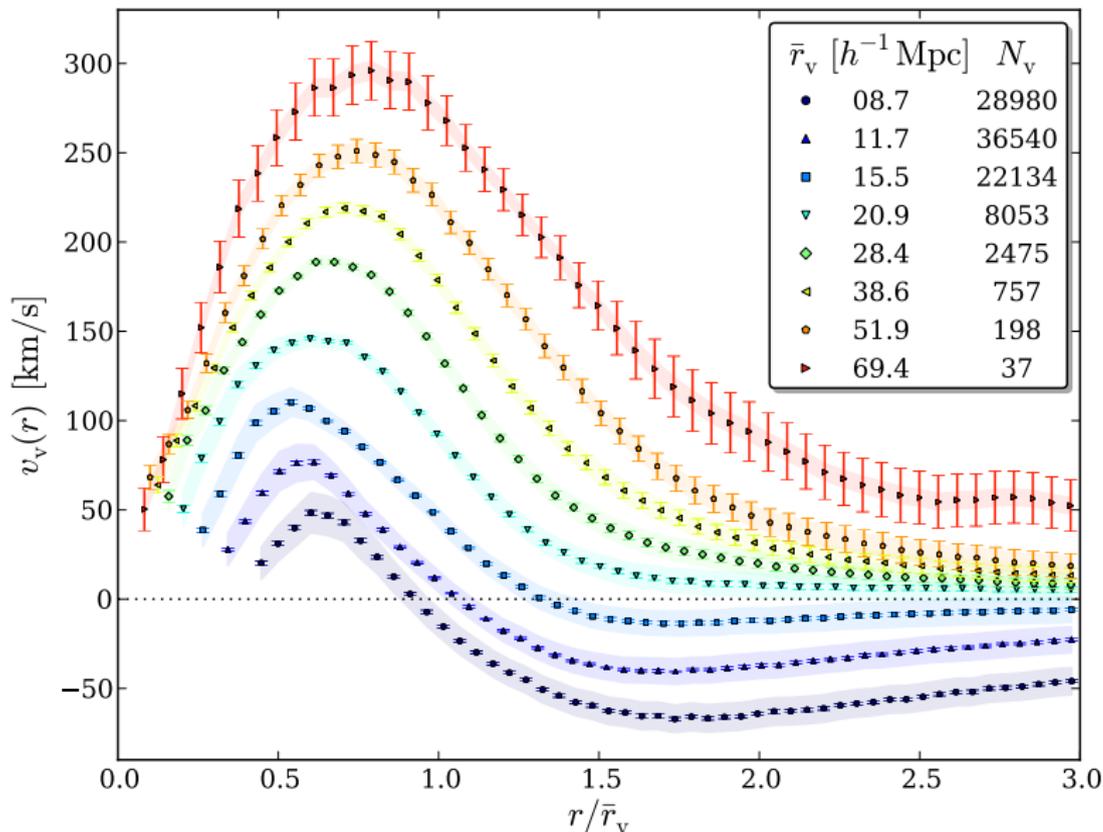
Void profile: density



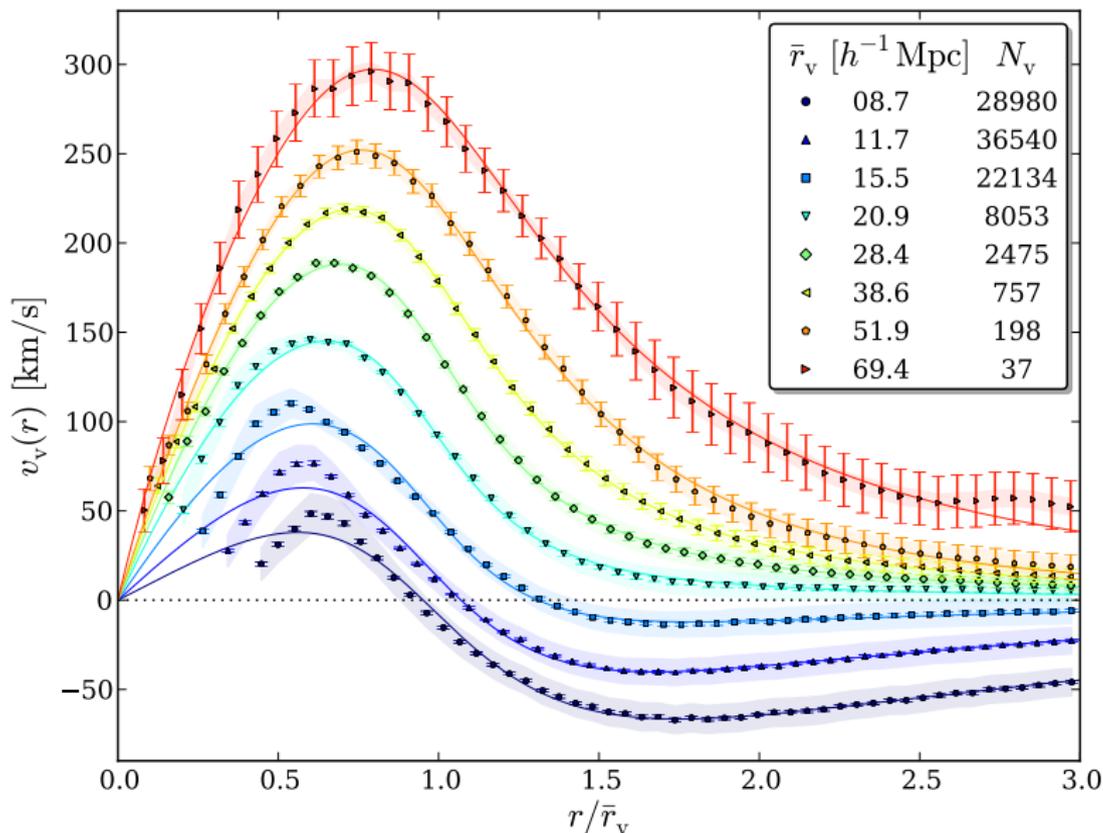
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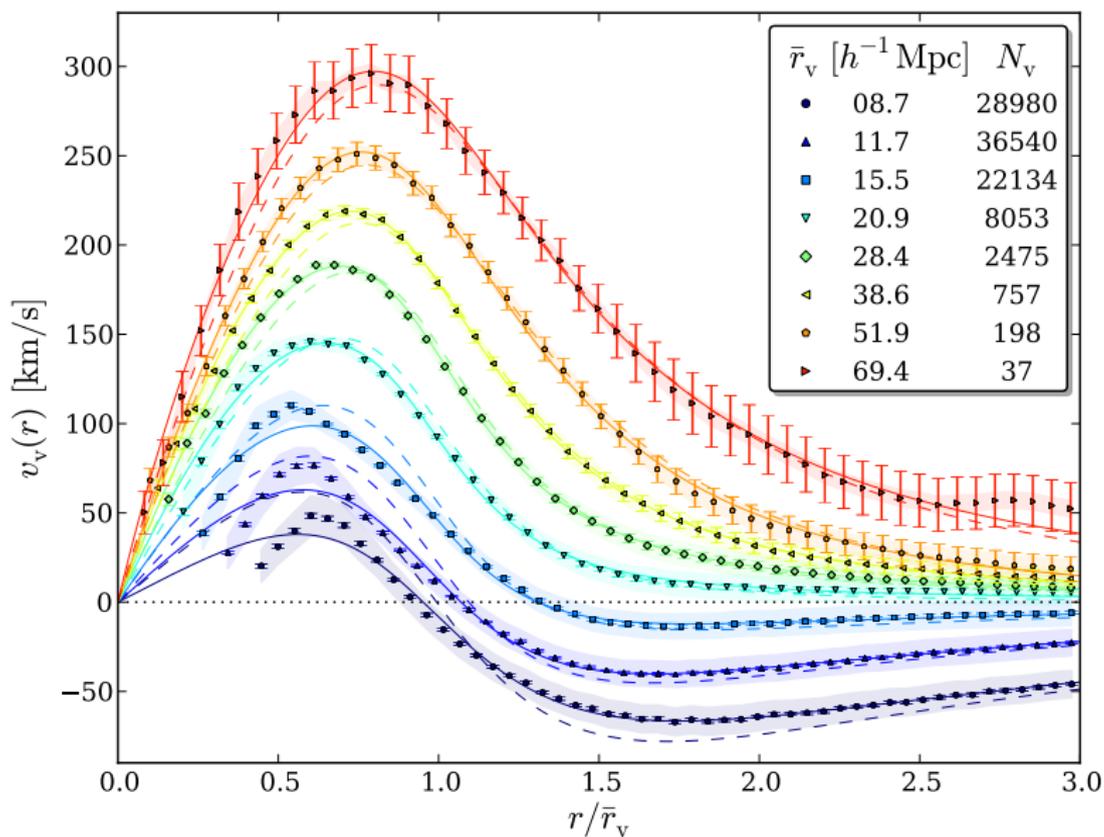
Void profile: velocity



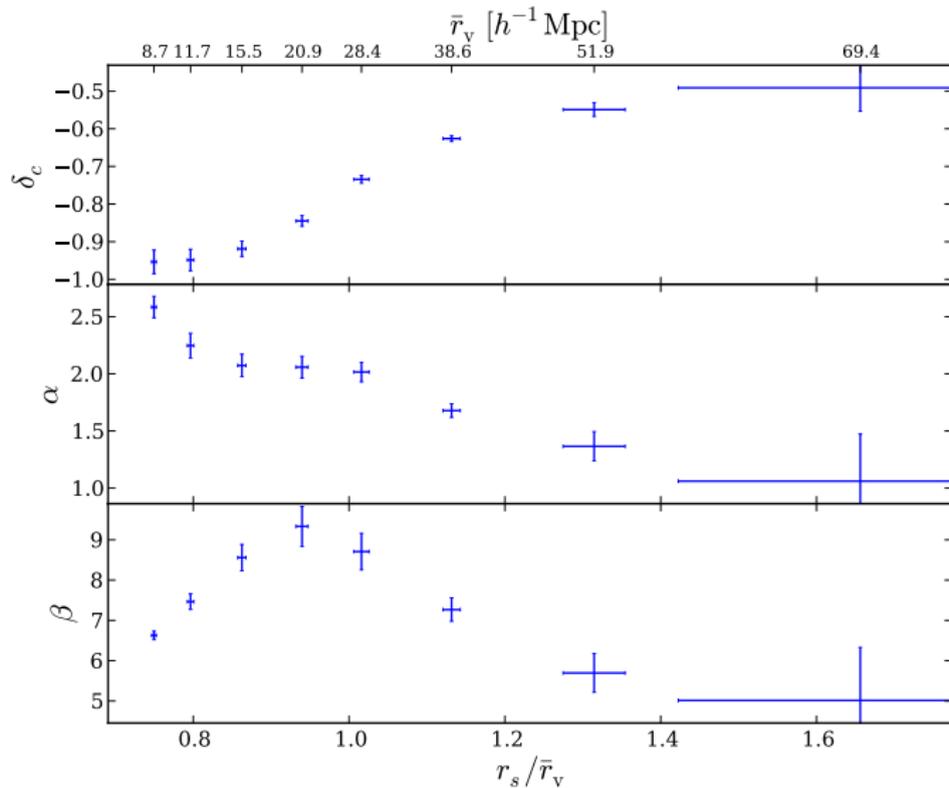
Void profile: velocity



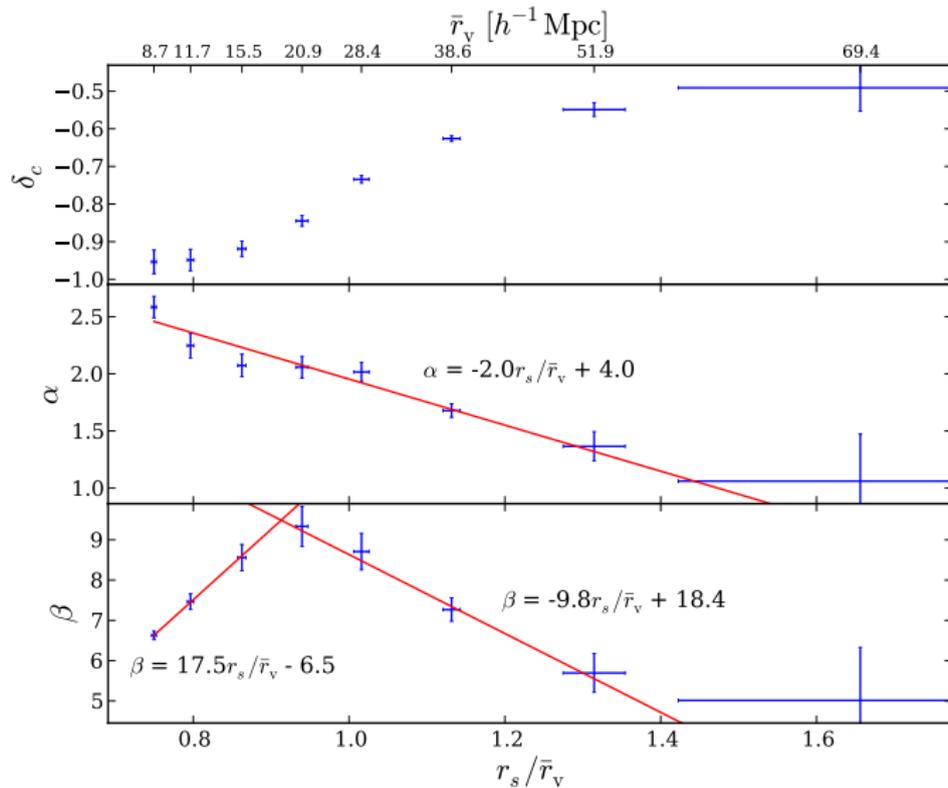
Void profile: velocity



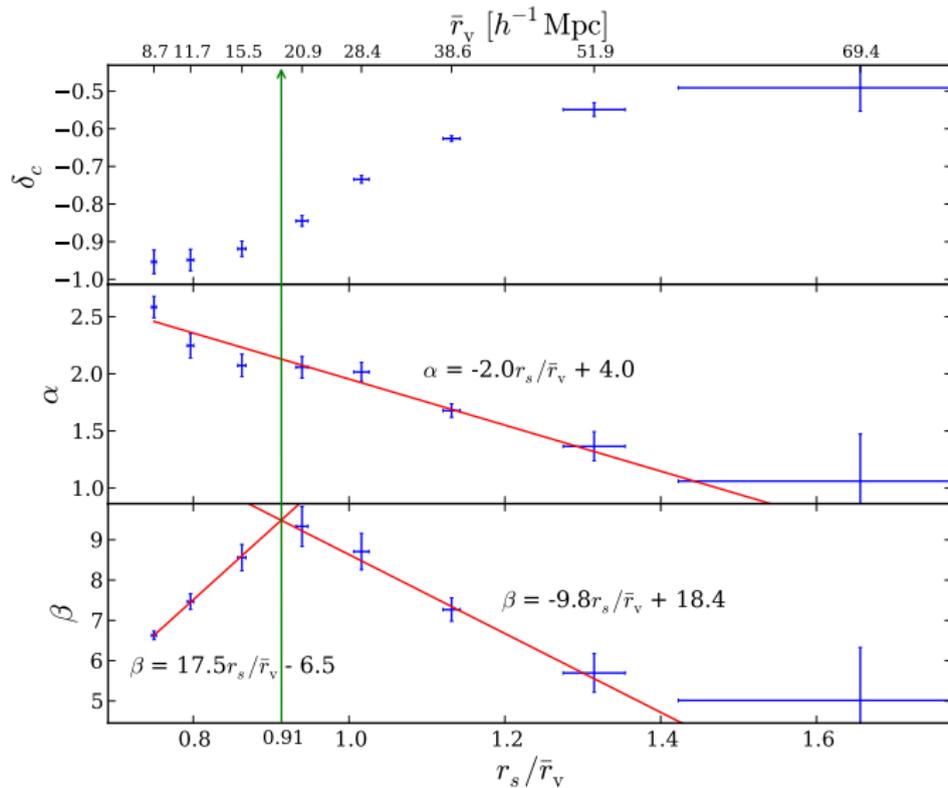
Void profile: parameters



Void profile: parameters



Void profile: parameters



Void profile: universality

Voids in redshift space

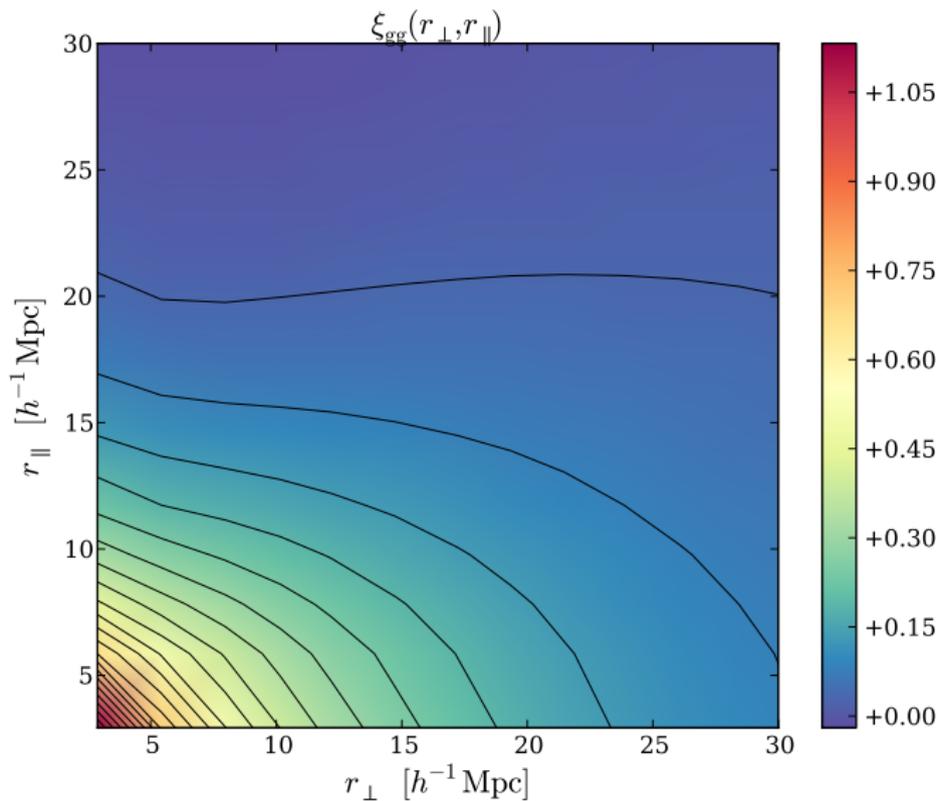
Peculiar motions of galaxies cause **redshift-space distortions**:

$$\mathbf{r}_s = \mathbf{r} + \mathbf{v}_{\parallel} H^{-1}(z)$$

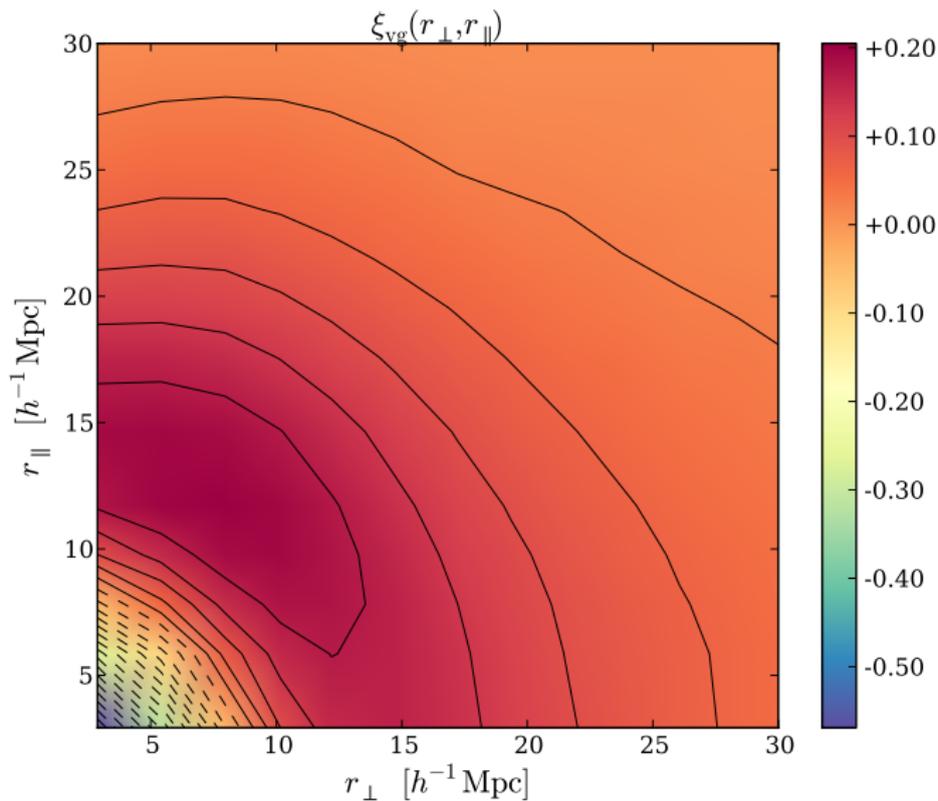
- ➔ \perp to line of sight:
Pancakes of God from linear growth
- ➔ \parallel to line of sight:
Fingers of God from nonlinear collapse
- ➔ Galaxy power spectrum no longer isotropic, what about void power spectrum?

Melott et al. (1998)

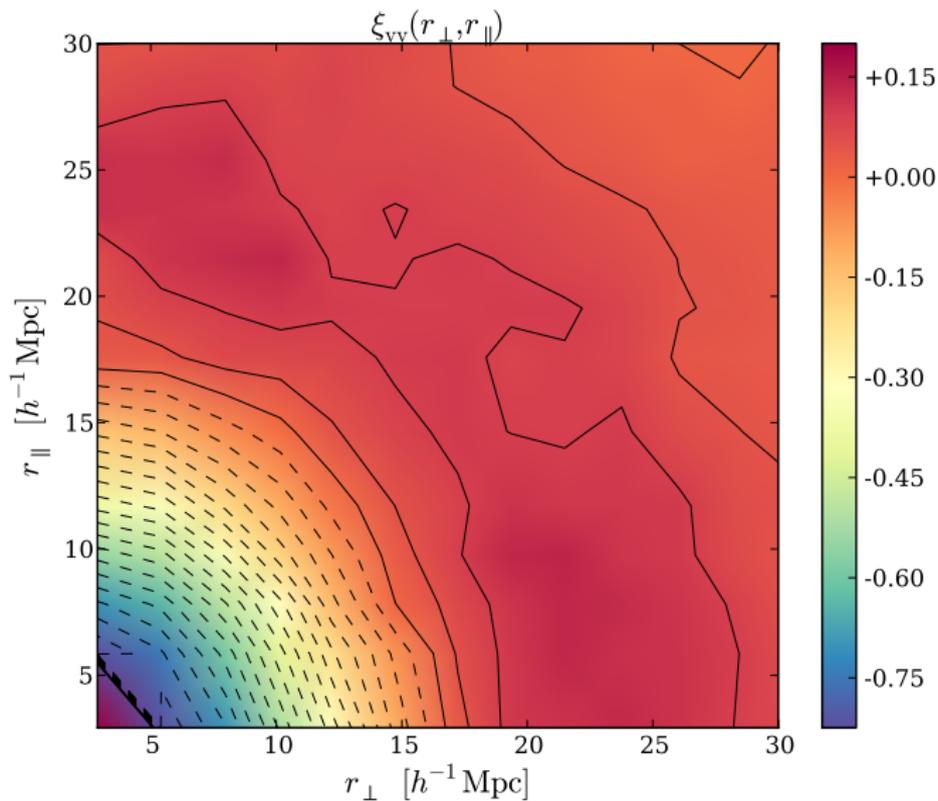
Voids in redshift space: 2D correlation function



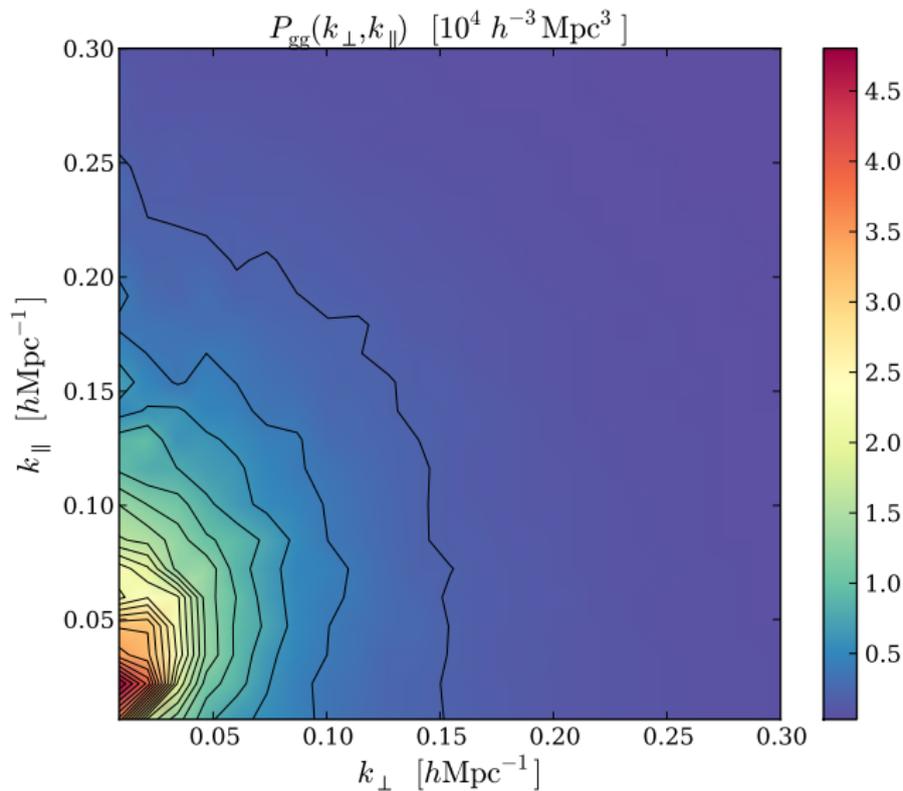
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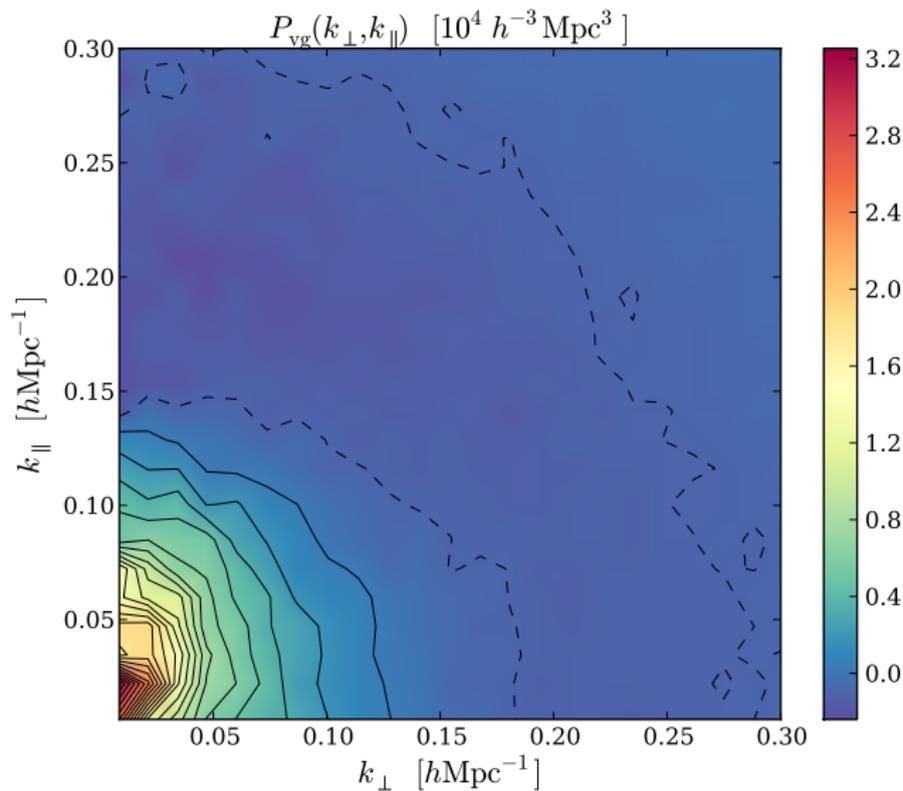
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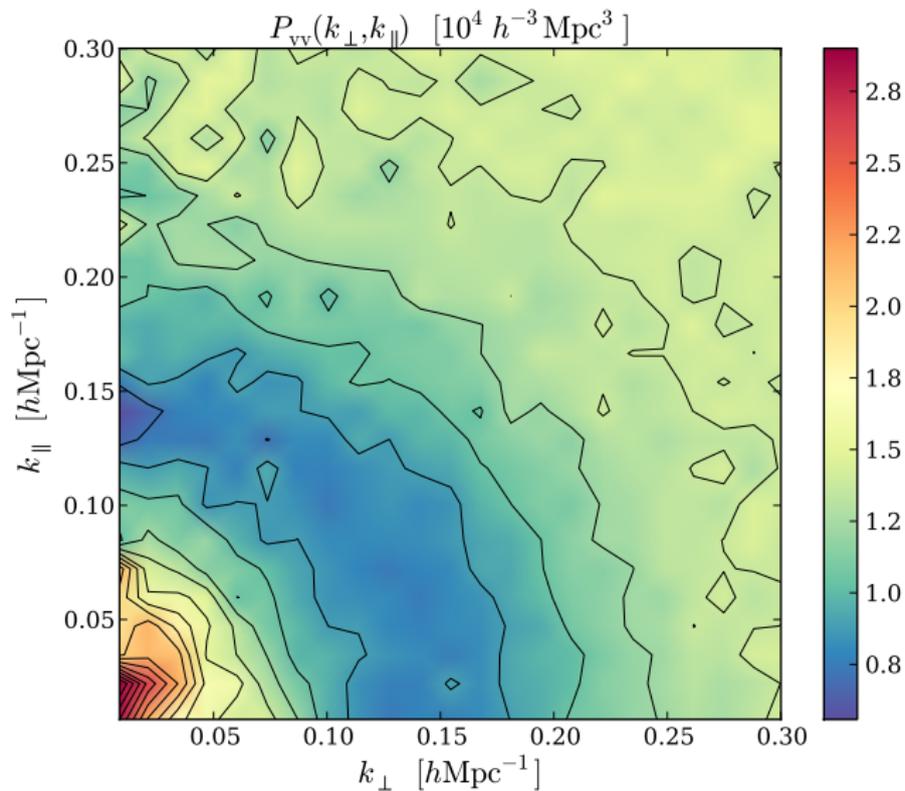
Voids in redshift space: 2D power spectrum



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Voids in redshift space: Alcock-Paczynski test

Perform *Alcock-Paczynski test* to constrain cosmological parameters:

- Angular separation $\delta r_{\perp} = D_A(z) \delta\Theta$
- Radial separation $\delta r_{\parallel} = cH^{-1}(z) \delta z$

Angular diameter distance & Hubble rate

$$D_A(z) = c \int_0^z H^{-1}(z') dz' \quad , \quad H(z) = H_0 \sqrt{\Omega_m(1+z)^3 + \Omega_\Lambda}$$

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Any deviation from the fiducial cosmology causes geometric distortions. \Rightarrow Measure **ellipticity** of power spectrum:

$$\epsilon(k_i) \equiv \int_{\Delta k_i} k_{\parallel}^2 P(\mathbf{k}) d^3 k \quad / \quad \int_{\Delta k_i} k_{\perp}^2 P(\mathbf{k}) d^3 k$$

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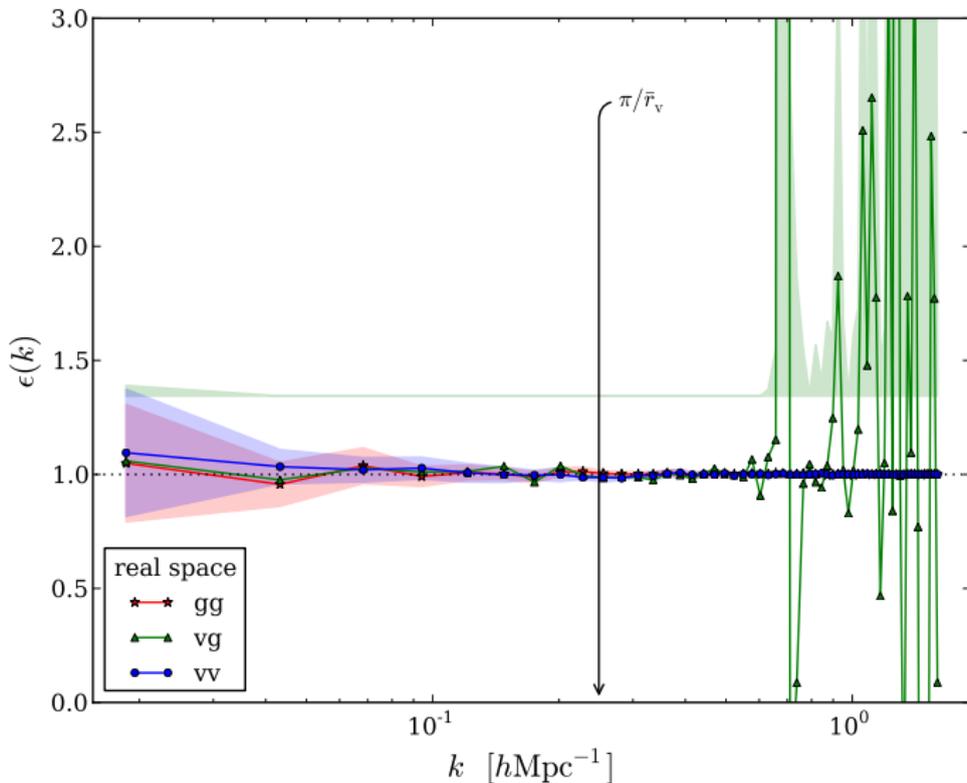
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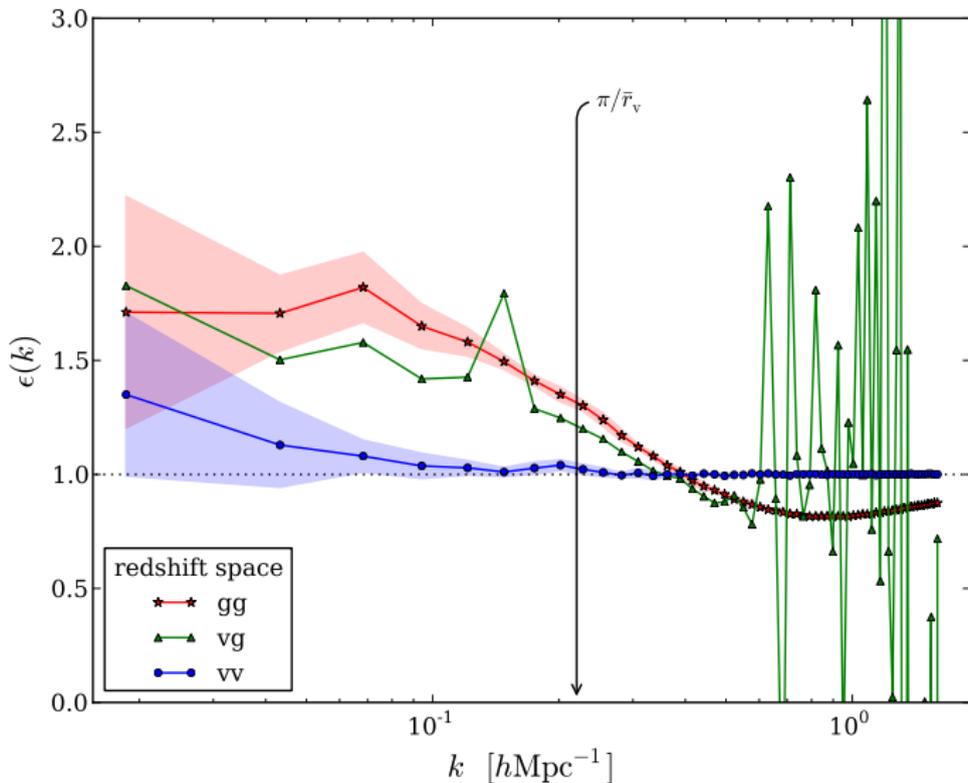
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Problem: degeneracy with redshift-space distortions

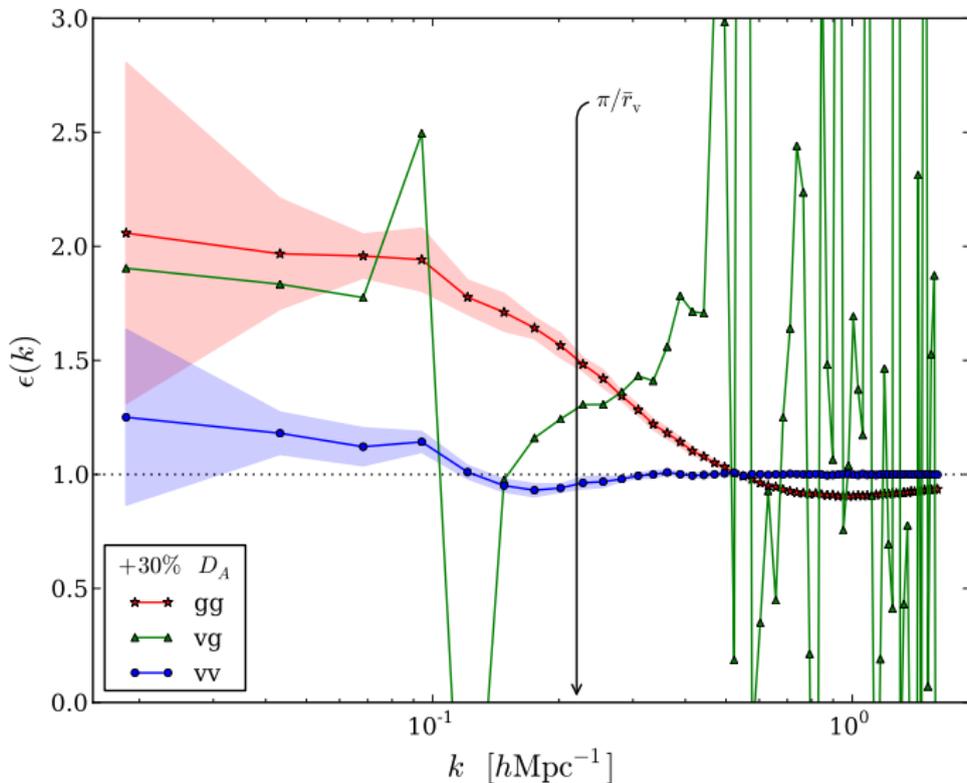
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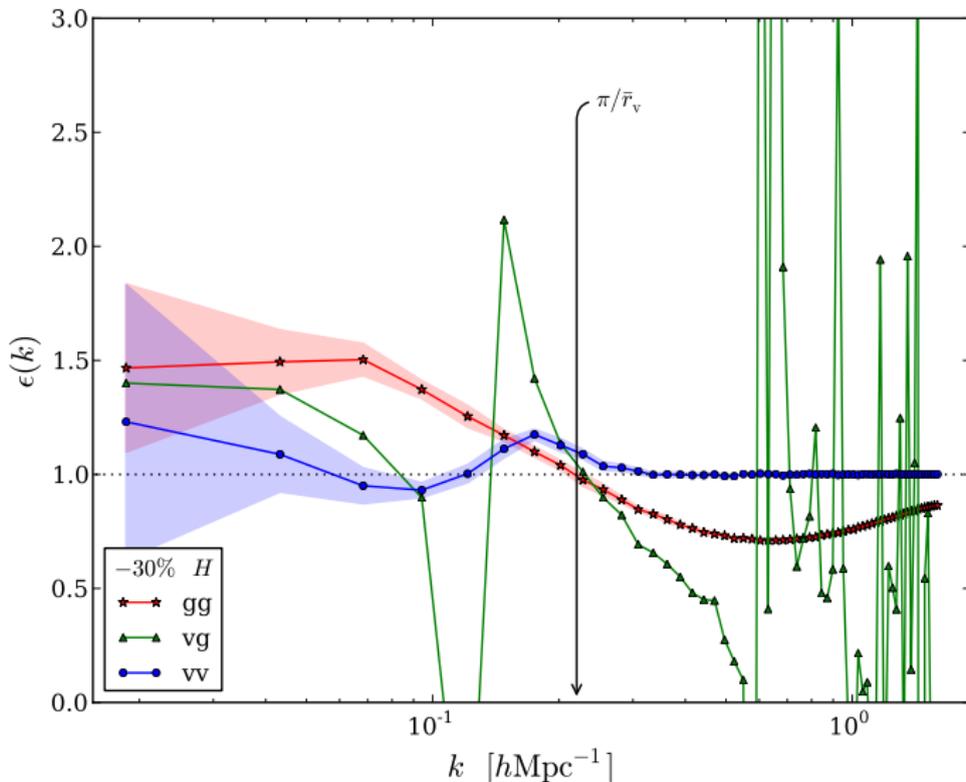
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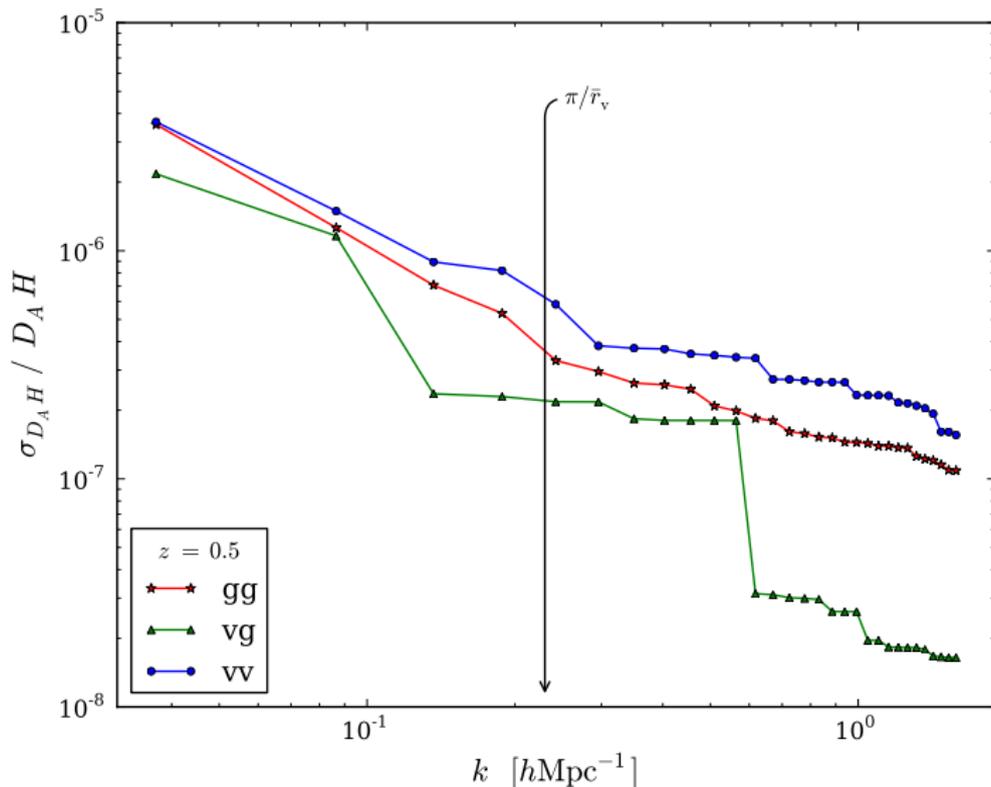
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Outlook

- Universal void density- and velocity profiles emerge from self-similarity / fractal nature of large-scale structure?
- Void auto-correlations provide pristine Alcock-Paczynski test?

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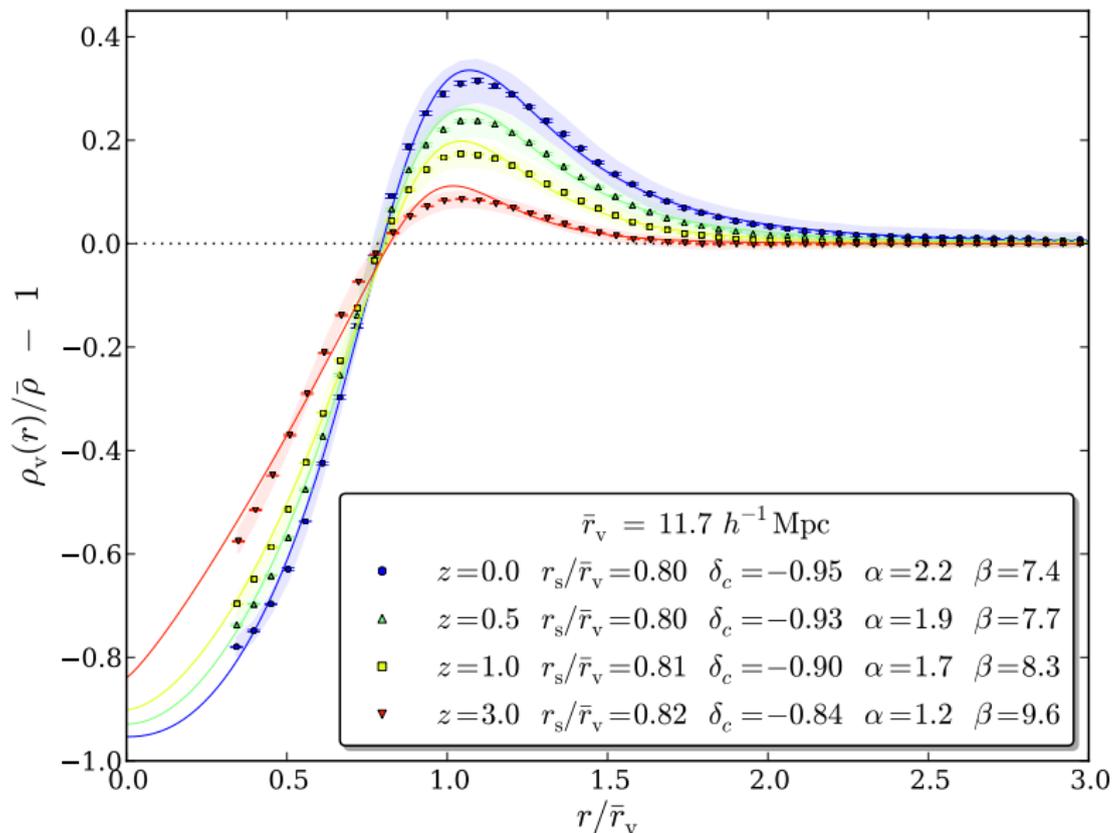
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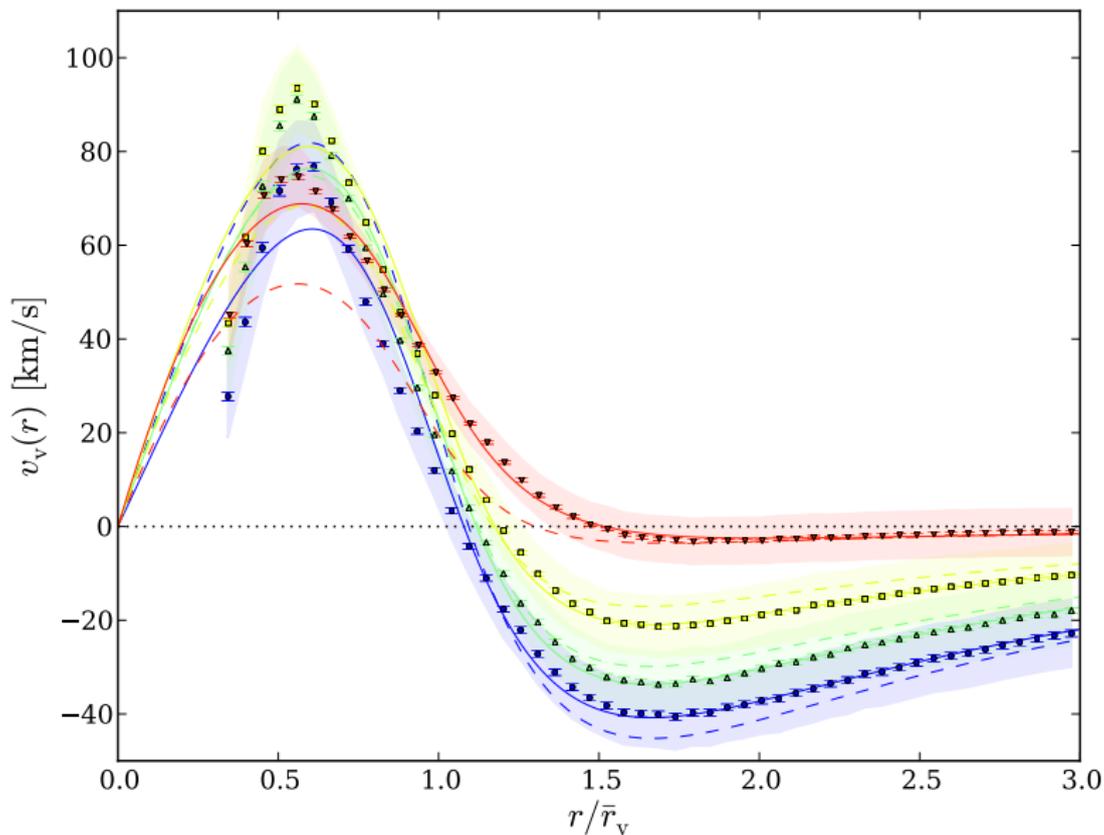
Questions ?

Thank you !

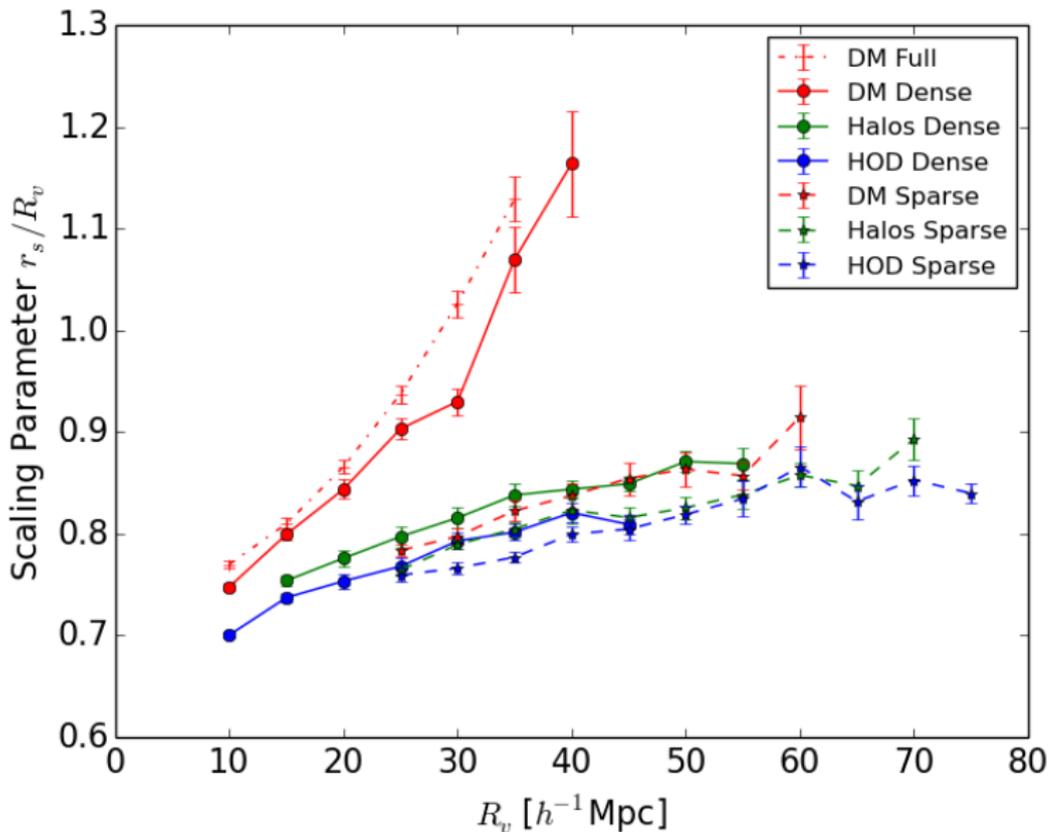
Void profile: universality



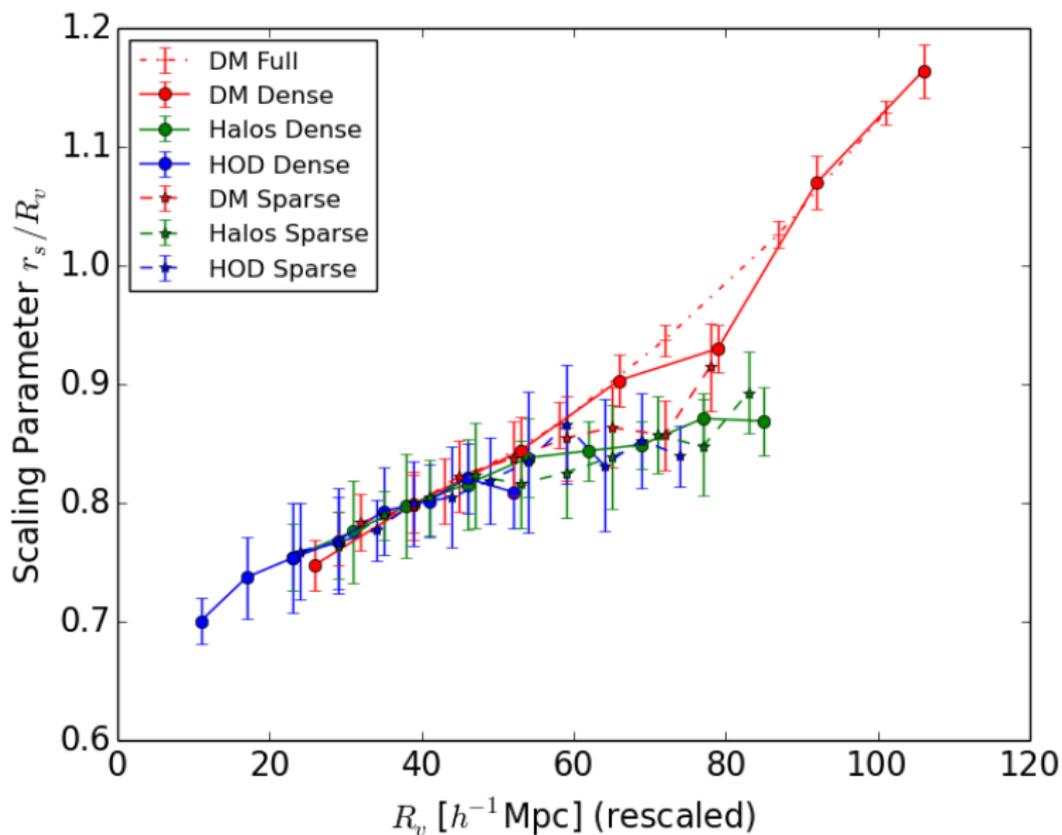
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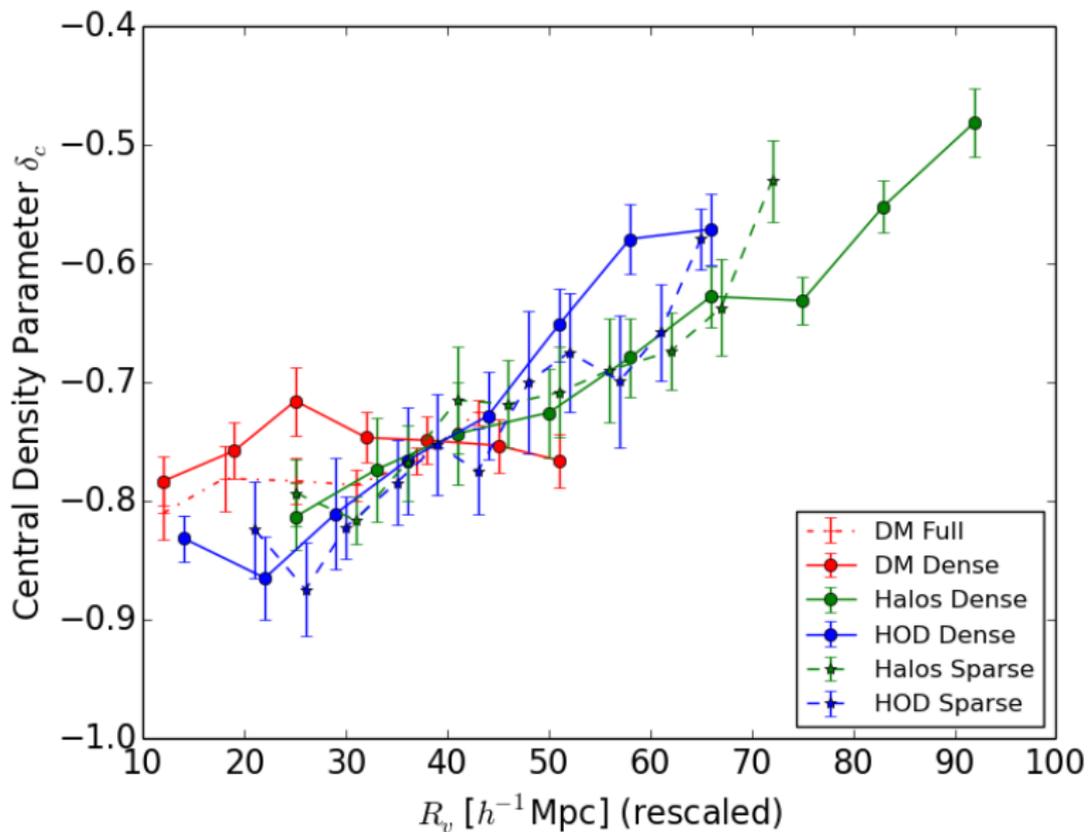
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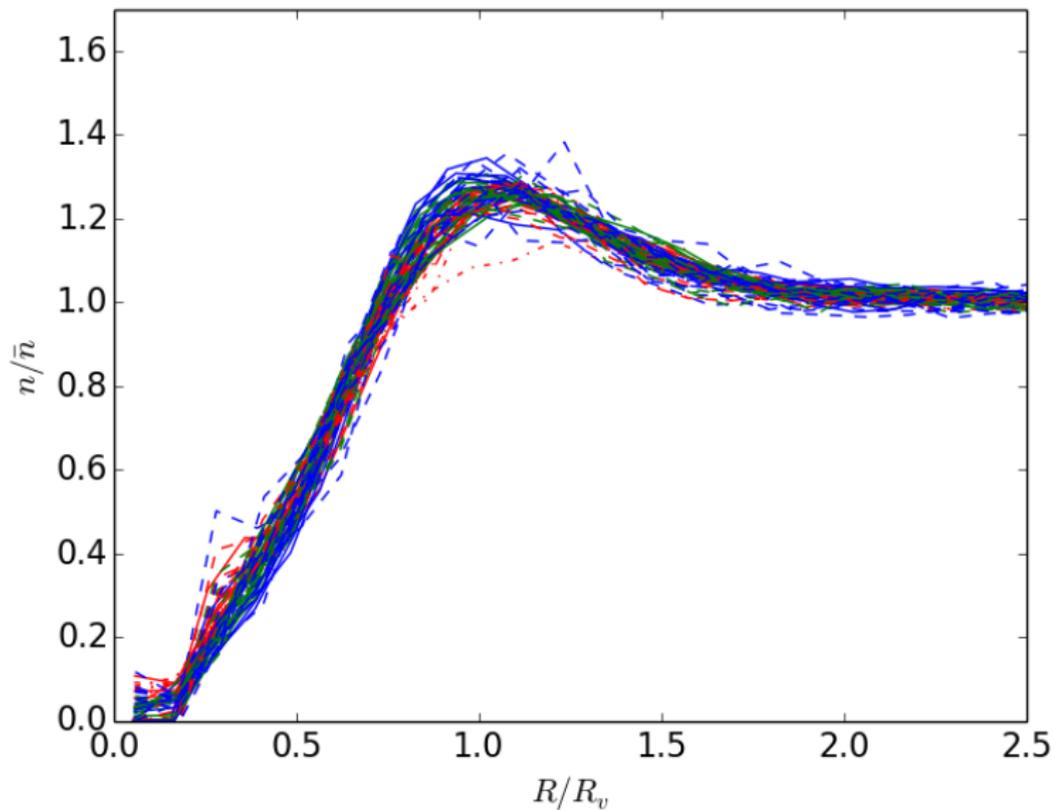
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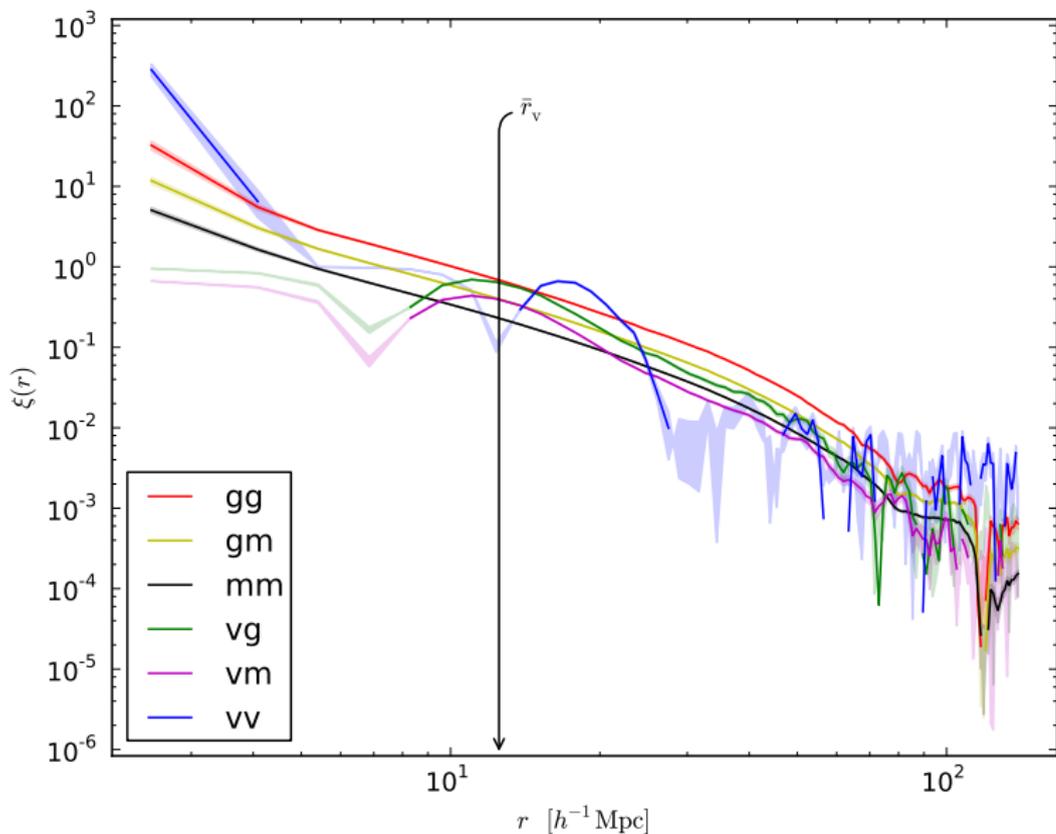
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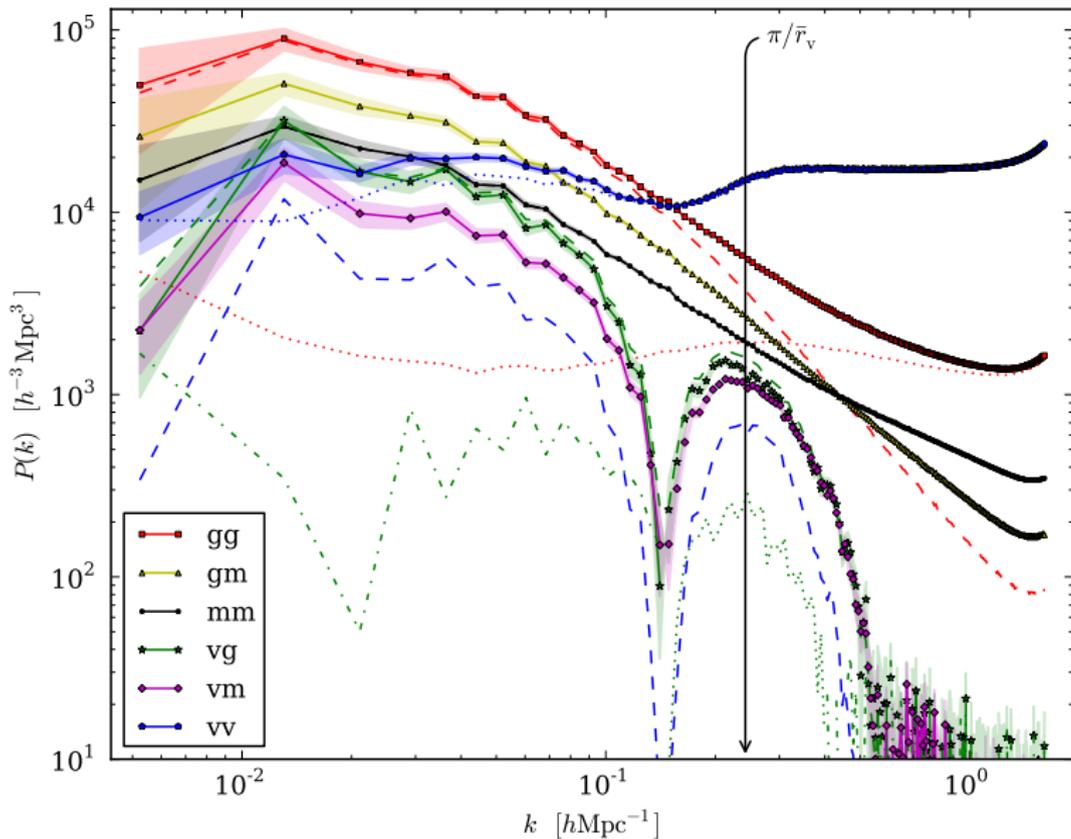
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Voids in redshift space: 1D correlation function



Voids in redshift space: 1D power spectrum



Void number function

