

Large-scale structure and the intrinsic alignment of galaxies

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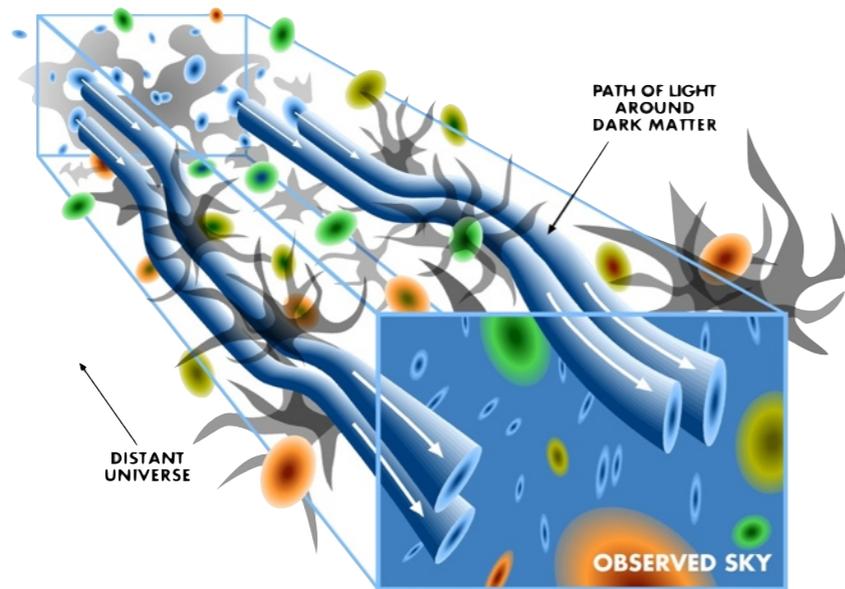
IAU 308 - Tallinn
June 28, 2014



Overview

- Weak lensing and intrinsic alignments (IA)
- Measuring IA in observations
- Modeling IA
 - Tidal (linear) alignment
 - IA as a probe of structure and galaxy/halo physics

IA and weak lensing



$$\gamma^{obs} = \gamma^I + \gamma^G$$

IA auto-correlation \rightarrow

$$\langle \gamma_i^{obs} \gamma_j^{obs} \rangle = \langle \gamma_i^G \gamma_j^G \rangle + \langle \gamma_i^I \gamma_j^G \rangle + \langle \gamma_i^G \gamma_j^I \rangle + \langle \gamma_i^I \gamma_j^I \rangle$$

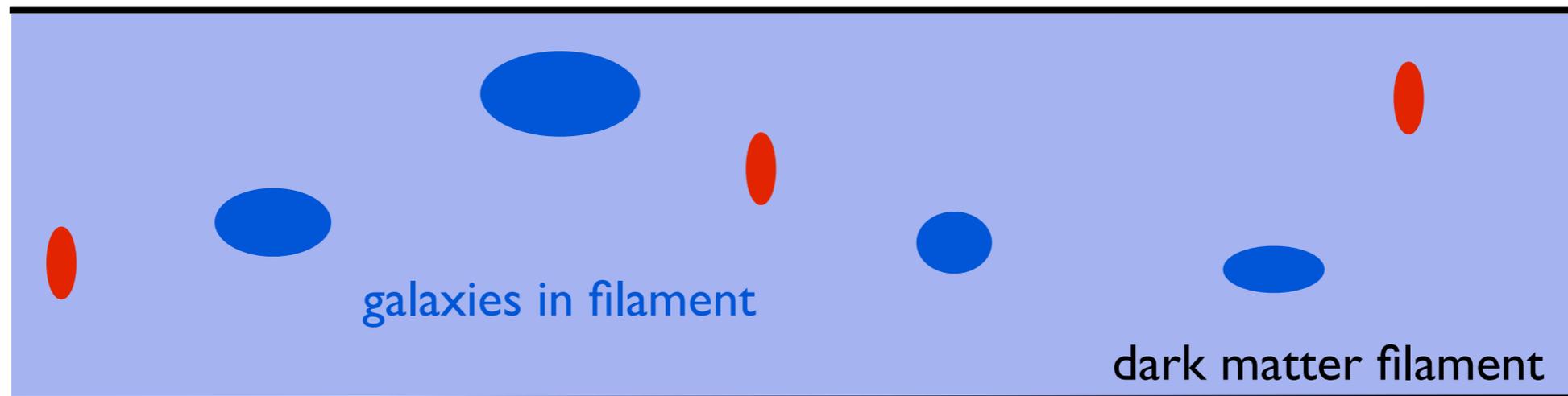
lensing signal \nearrow

IA-lensing cross-correlation \nearrow

*harder to remove

(Hirata & Seljak 2004)

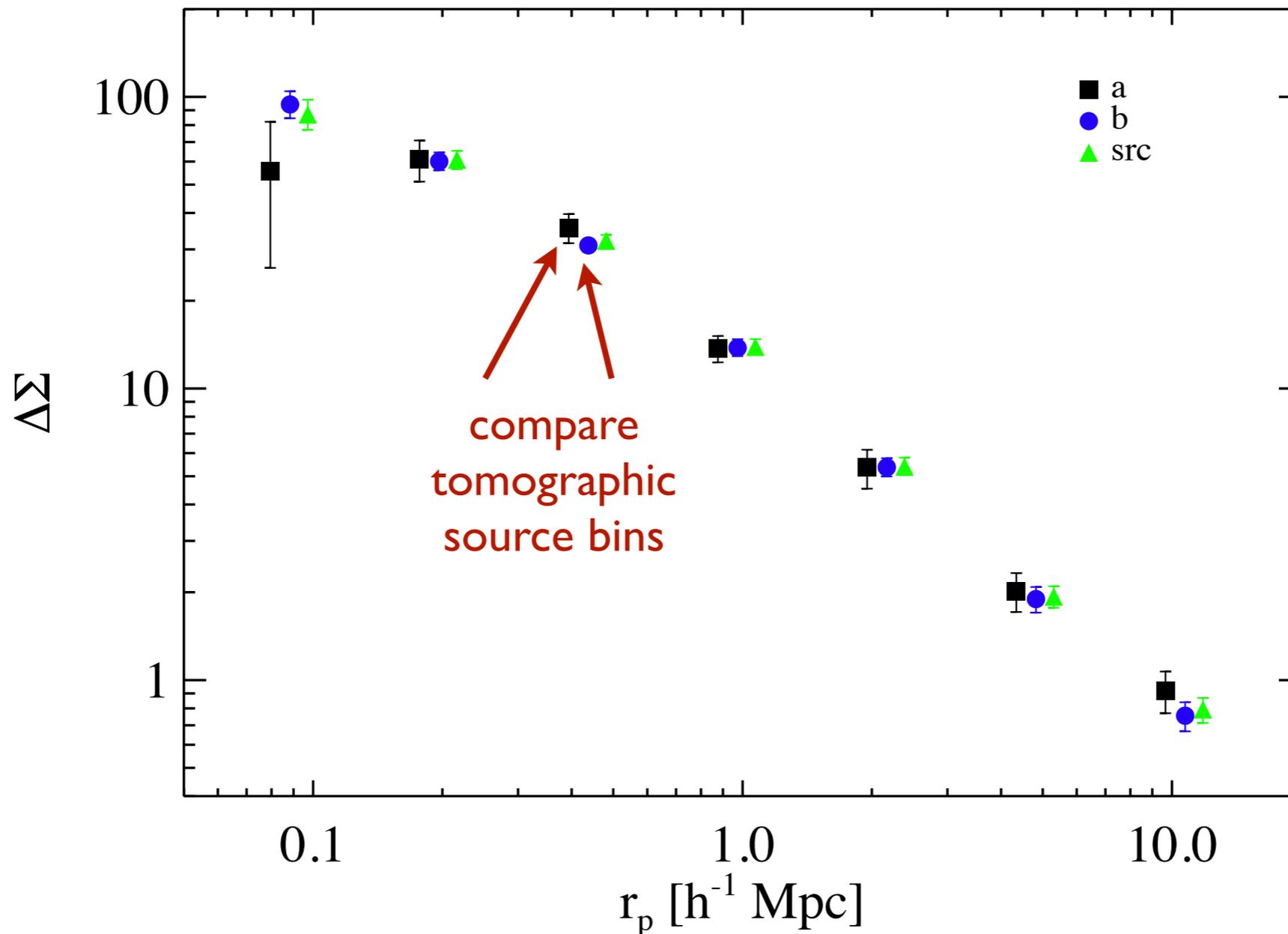
distorted background galaxies, seen behind and around filament



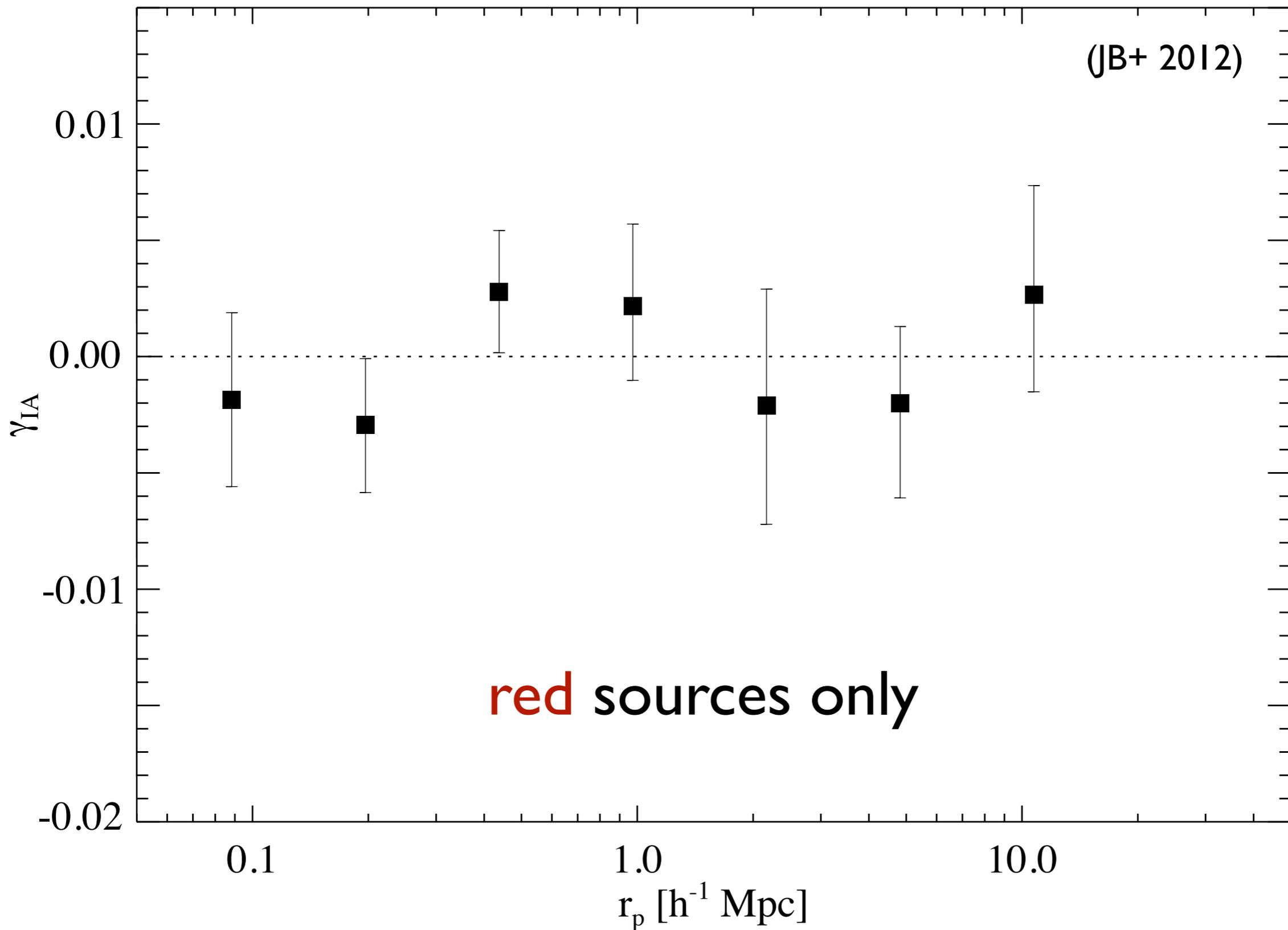
Separating IA and lensing signals

$$\widetilde{\Delta\Sigma} \sim \left\langle \widetilde{\Sigma}_c \widetilde{\gamma}_t \right\rangle = \left\langle \widetilde{\Sigma}_c \left(\Delta\Sigma \Sigma_c^{-1} + \gamma_{t,IA} \right) \right\rangle$$

observed signal
lensing signal
IA contamination

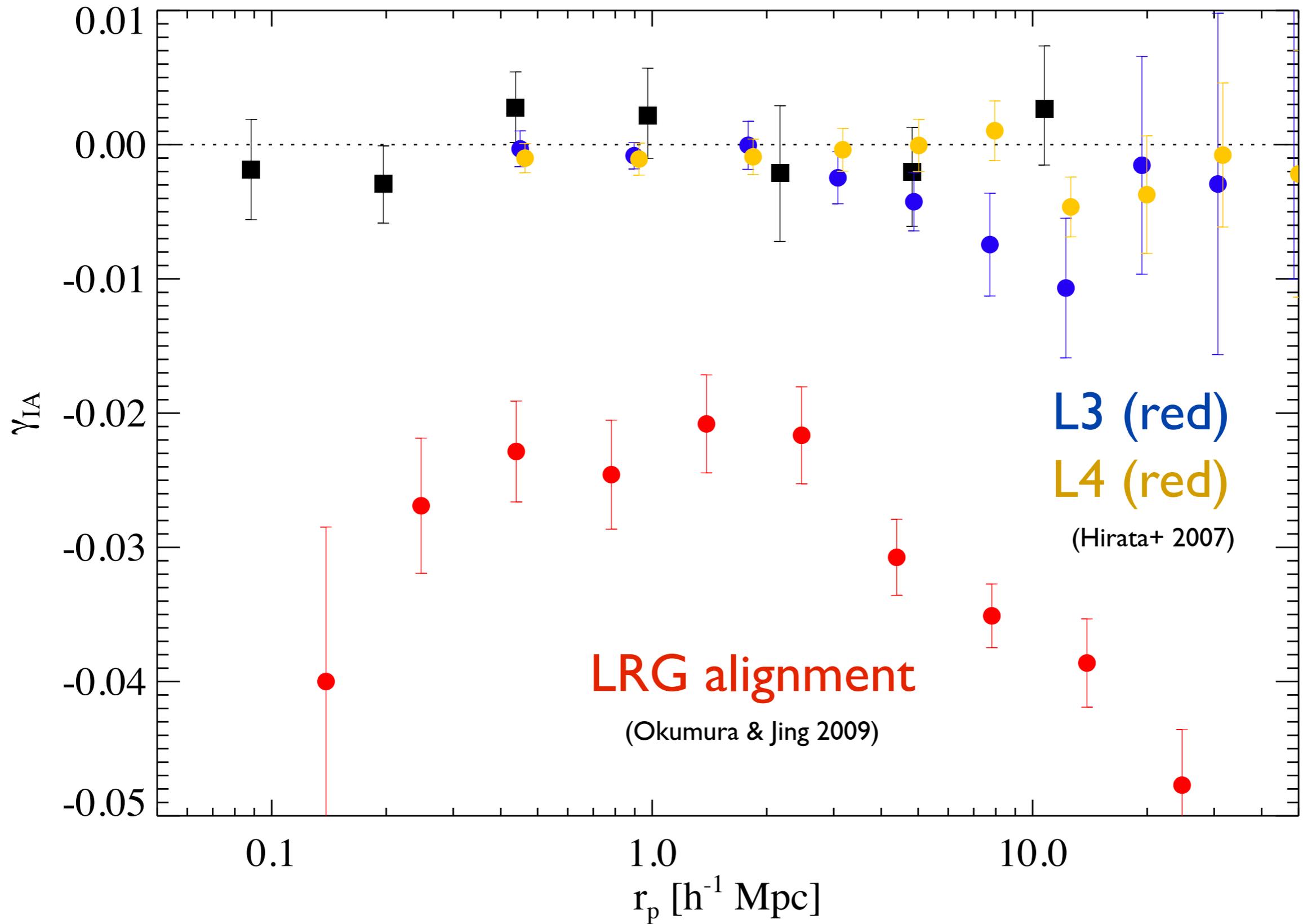


IA per excess lens-source pair



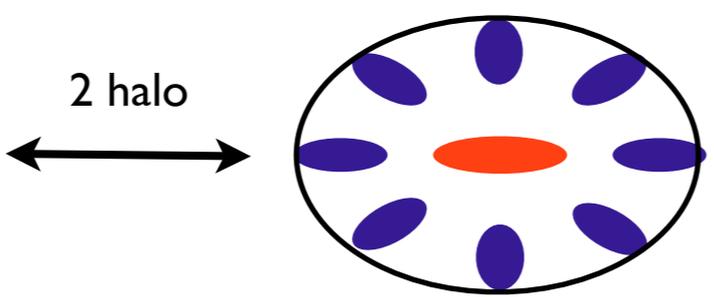
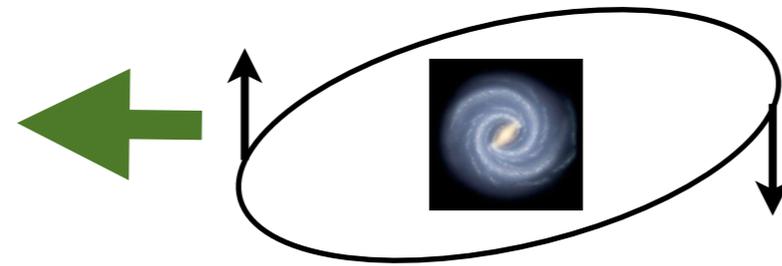
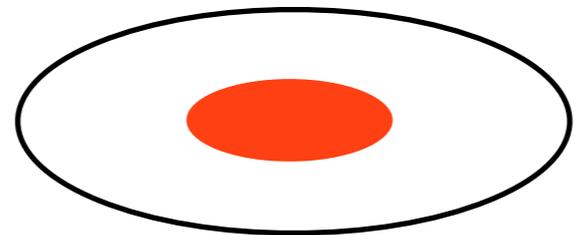
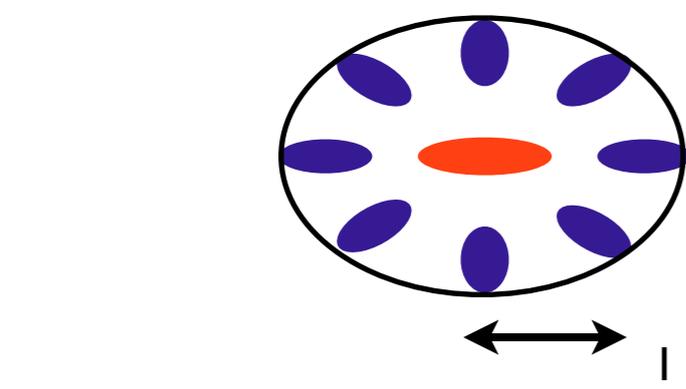
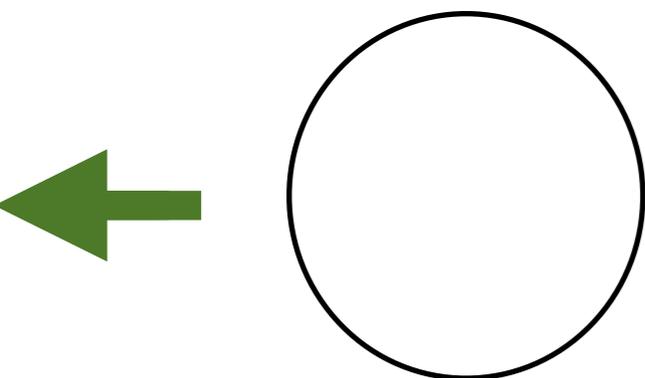
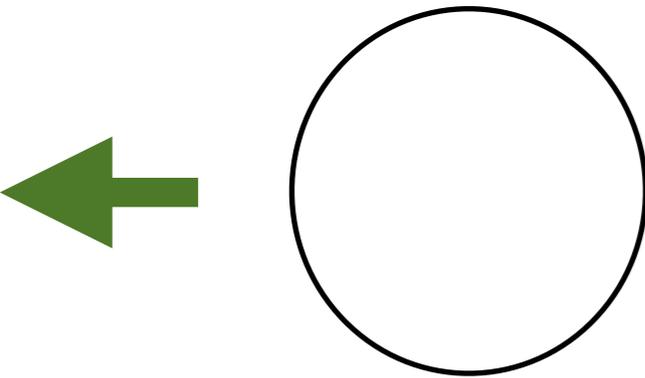
SDSS DR7 LRG lenses ($\sim 62\text{k}$) at $z=0.16-0.36$

IA per excess lens-source pair



IA modeling

tidal field: T



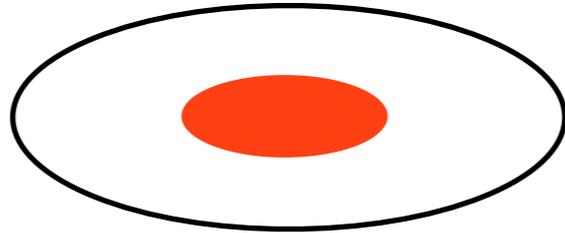
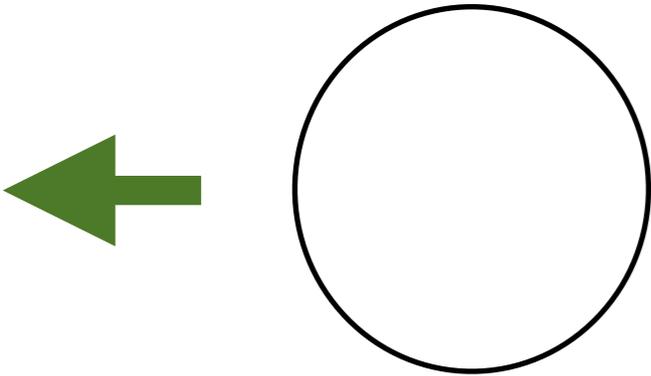
tidal alignment:
linear in T
(Catelan+ 2001)

tidal torquing:
quadratic in T
(Talks by Lee, Codis, Pichon)

hybrid halo model
(e.g. Schneider&Bridle 2009)

IA modeling

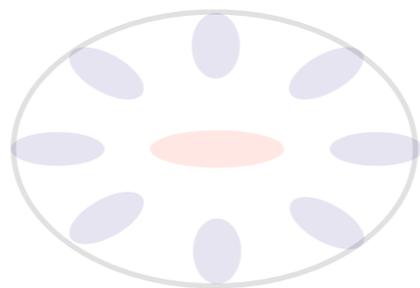
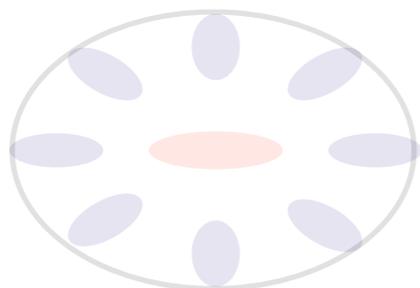
tidal field: T



tidal alignment:
linear in T
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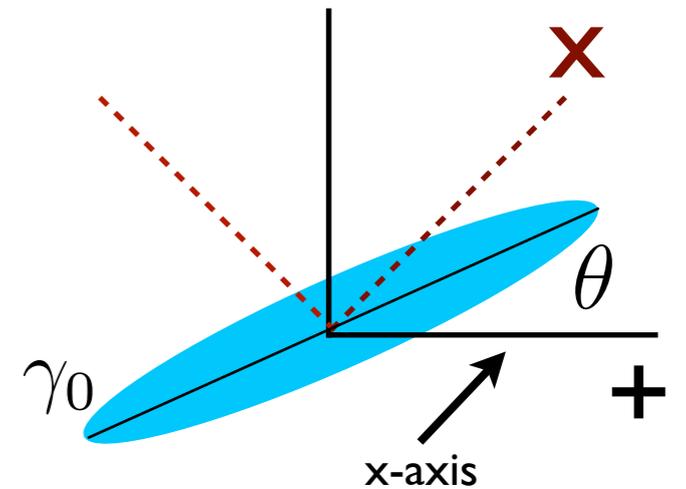
hybrid halo model
(e.g. Schneider&Bridle 2009)

Tidal alignment model

(Catelan+ 2001)

$$\gamma_{(+,\times)}^I = -\frac{C_1}{4\pi G} (\nabla_x^2 - \nabla_y^2, 2\nabla_x \nabla_y) S[\Psi_P]$$

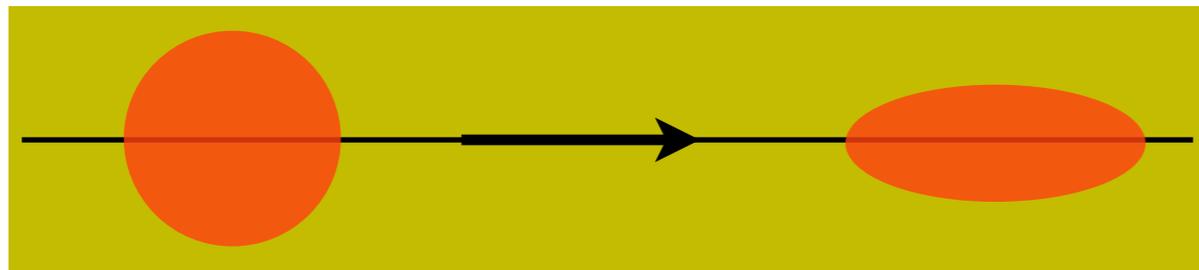
- Ellipticity aligns with tidal field
- Should dominate on large scales: $\sim P(k)$



$$\gamma_+ = \gamma_0 \cos 2\theta$$

$$\gamma_\times = \gamma_0 \sin 2\theta$$

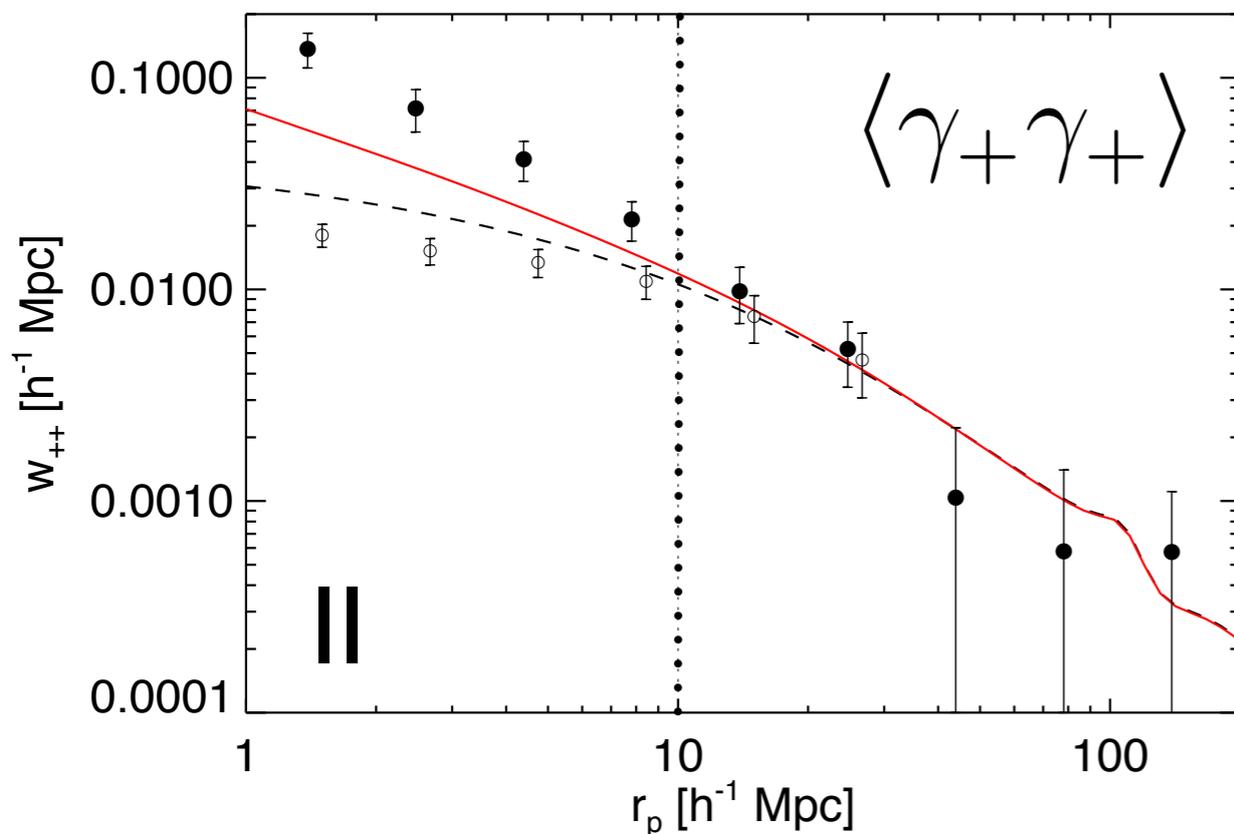
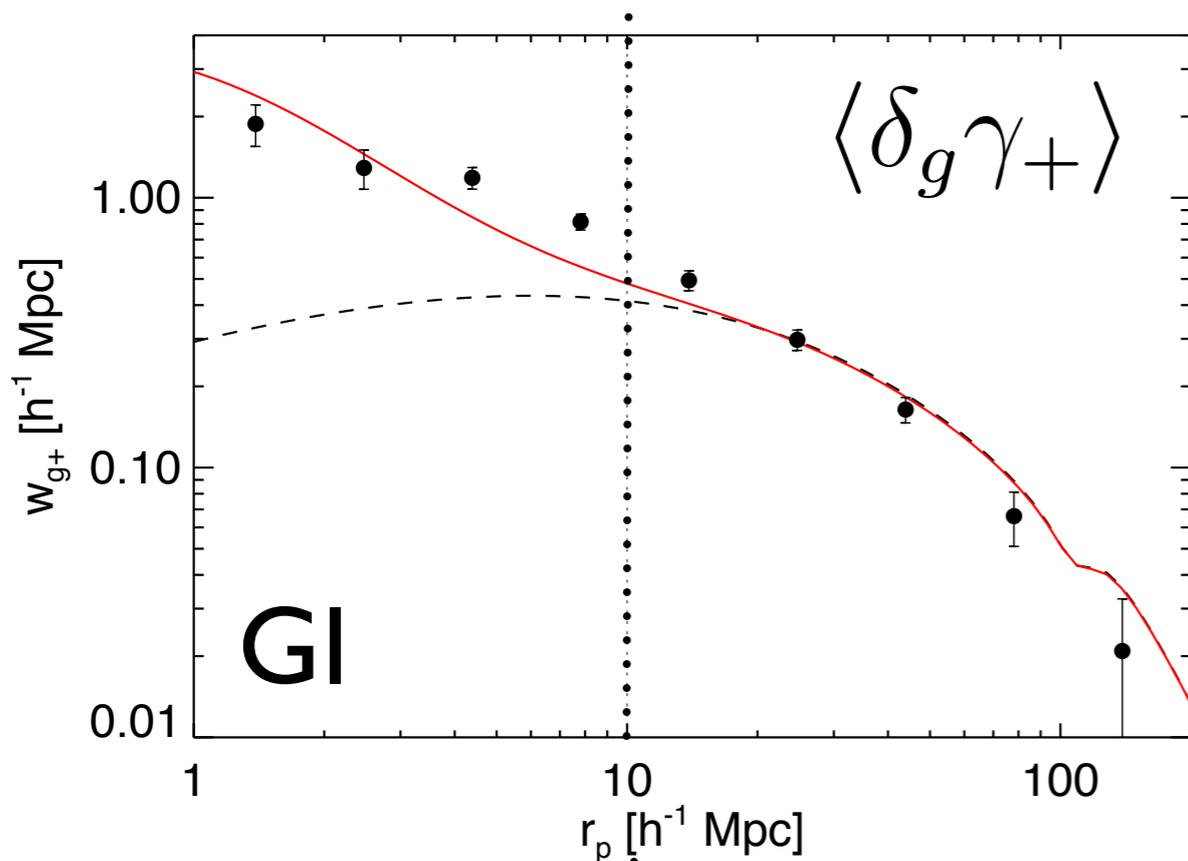
tidal field



The model works!

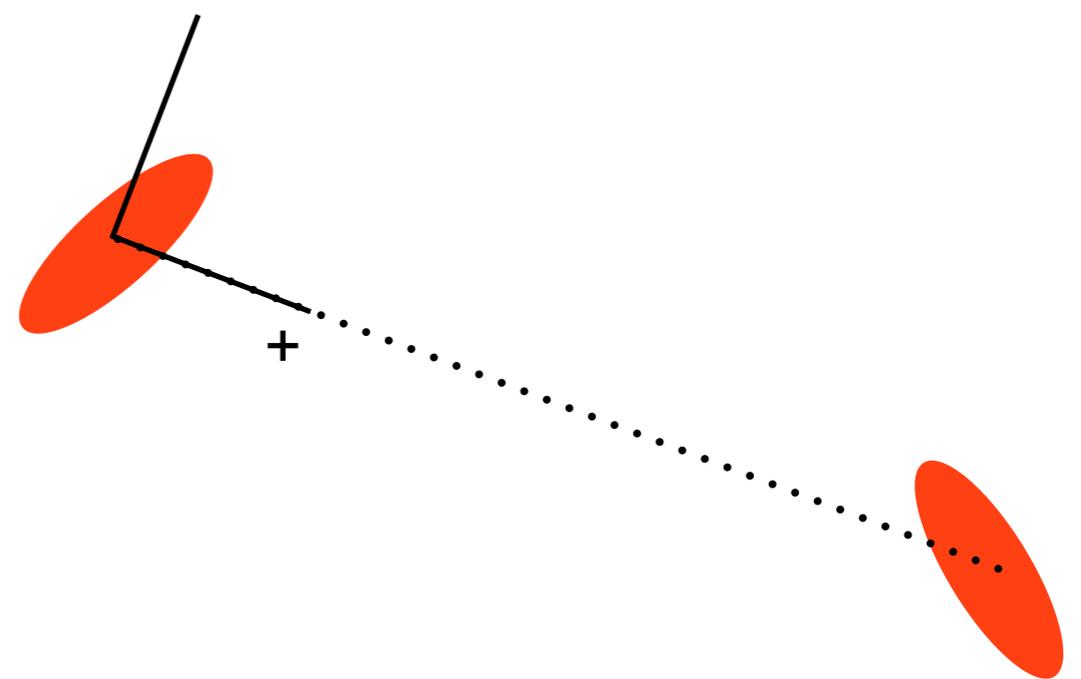
Luminous red galaxies (LRGs)
from SDSS DR7

$0.16 < z < 0.47$ (Okumura+ 2009; Okumura & Jing 2009)

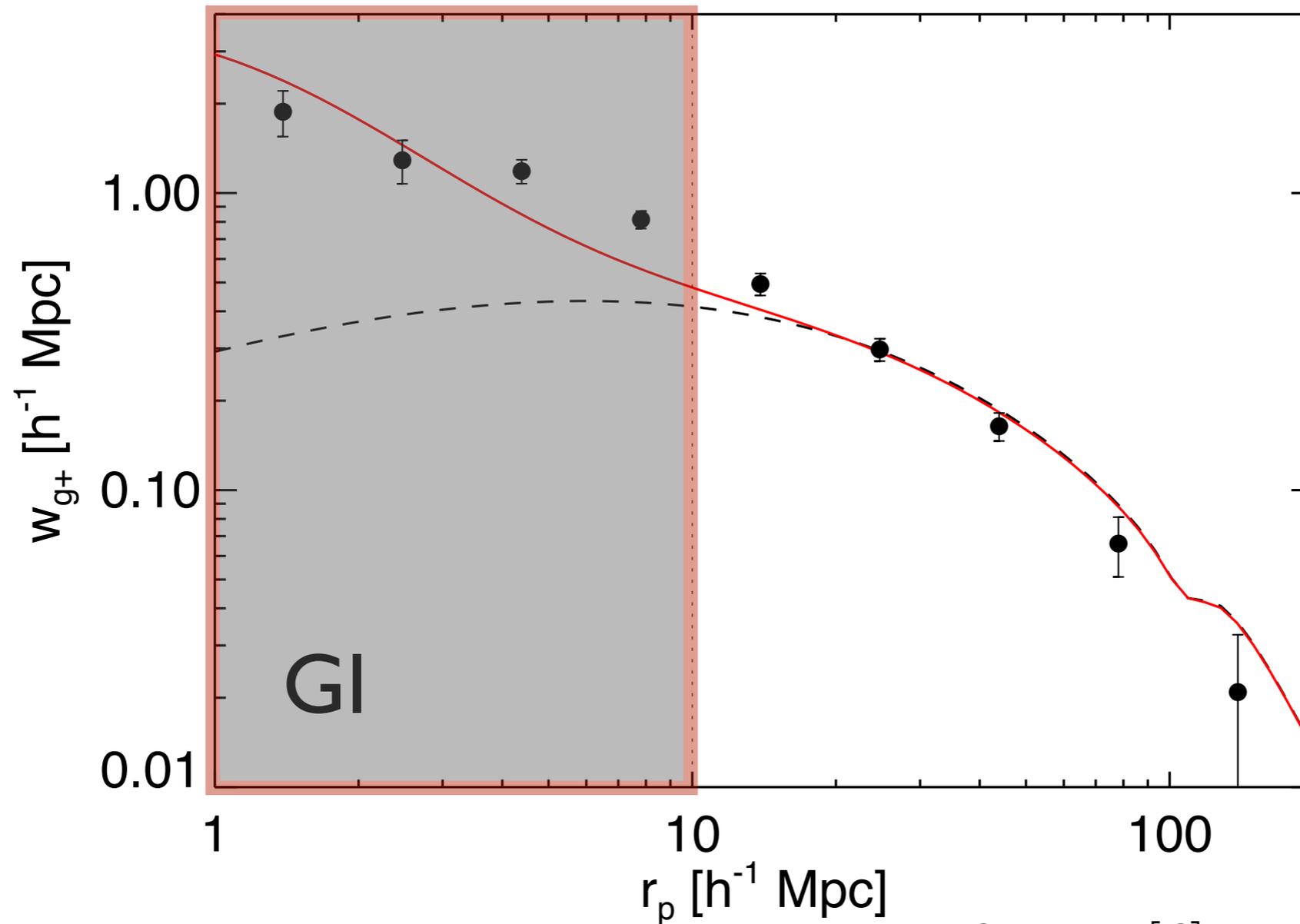


Test	$C_1 \rho_{\text{crit}}$ (JB+ 2011)
w_{g+}	0.125 ± 0.007
w_{++}	0.123 ± 0.014

(c.f. Joachimi, Mandelbaum+ 2011)



Modeling smaller scales



$$\delta_g = \mathcal{F}[\delta] \approx b_1 \delta + b_2 \delta^2 + \dots$$

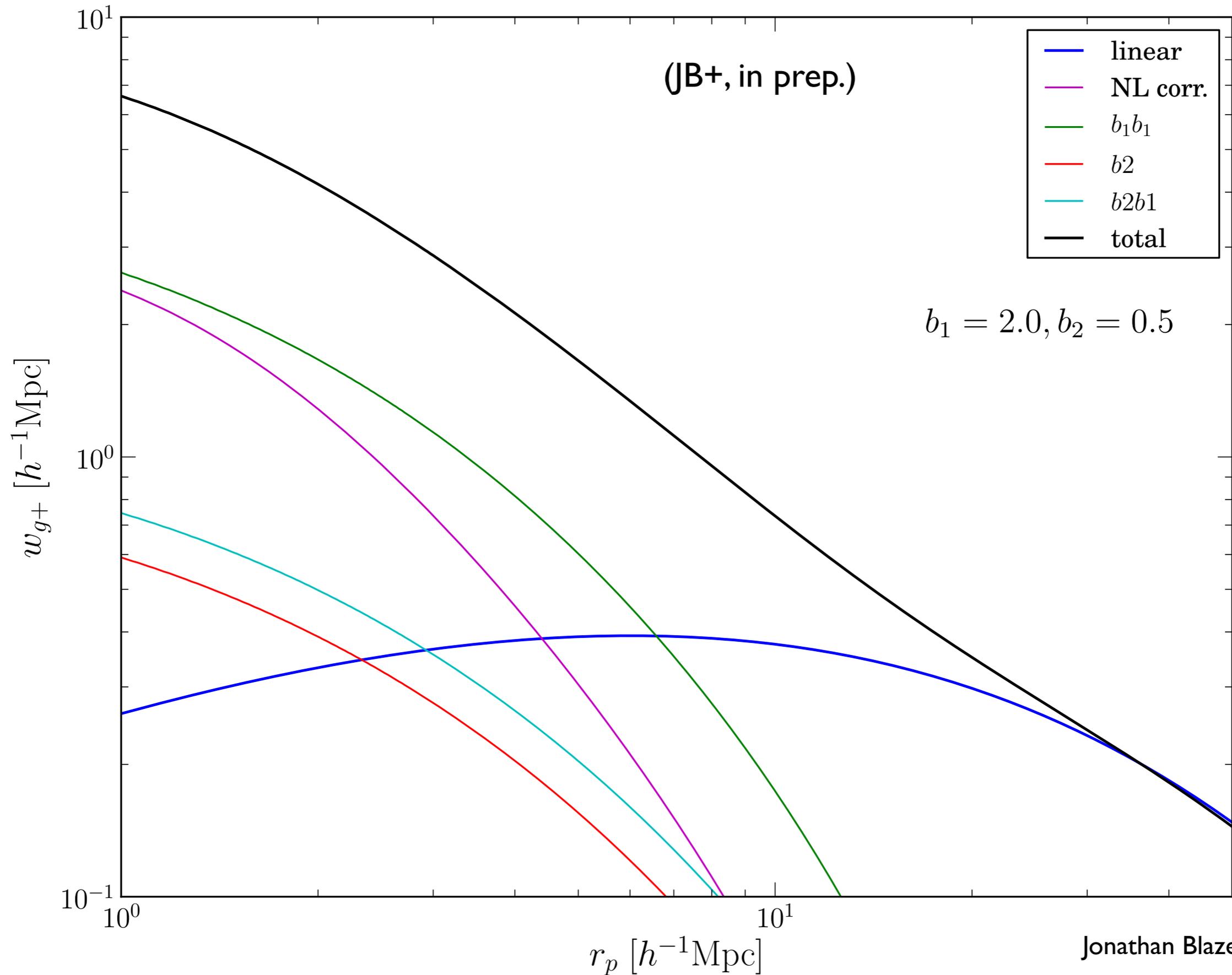
$$\langle \delta_g | (1 + \delta_g) \gamma_+ \rangle = \underbrace{b_1 \langle \delta | \gamma_+ \rangle}_{\text{“NLA model”}} + \underbrace{b_2 \langle \delta \delta | \gamma_+ \rangle}_{\text{nonlinear bias}} + \underbrace{b_1^2 \langle \delta | \delta \gamma_+ \rangle + b_2 b_1 \langle \delta \delta | \delta \gamma \rangle}_{\text{density weighting}} + \mathcal{O}(P_{\text{lin}}^3)$$

“NLA model”

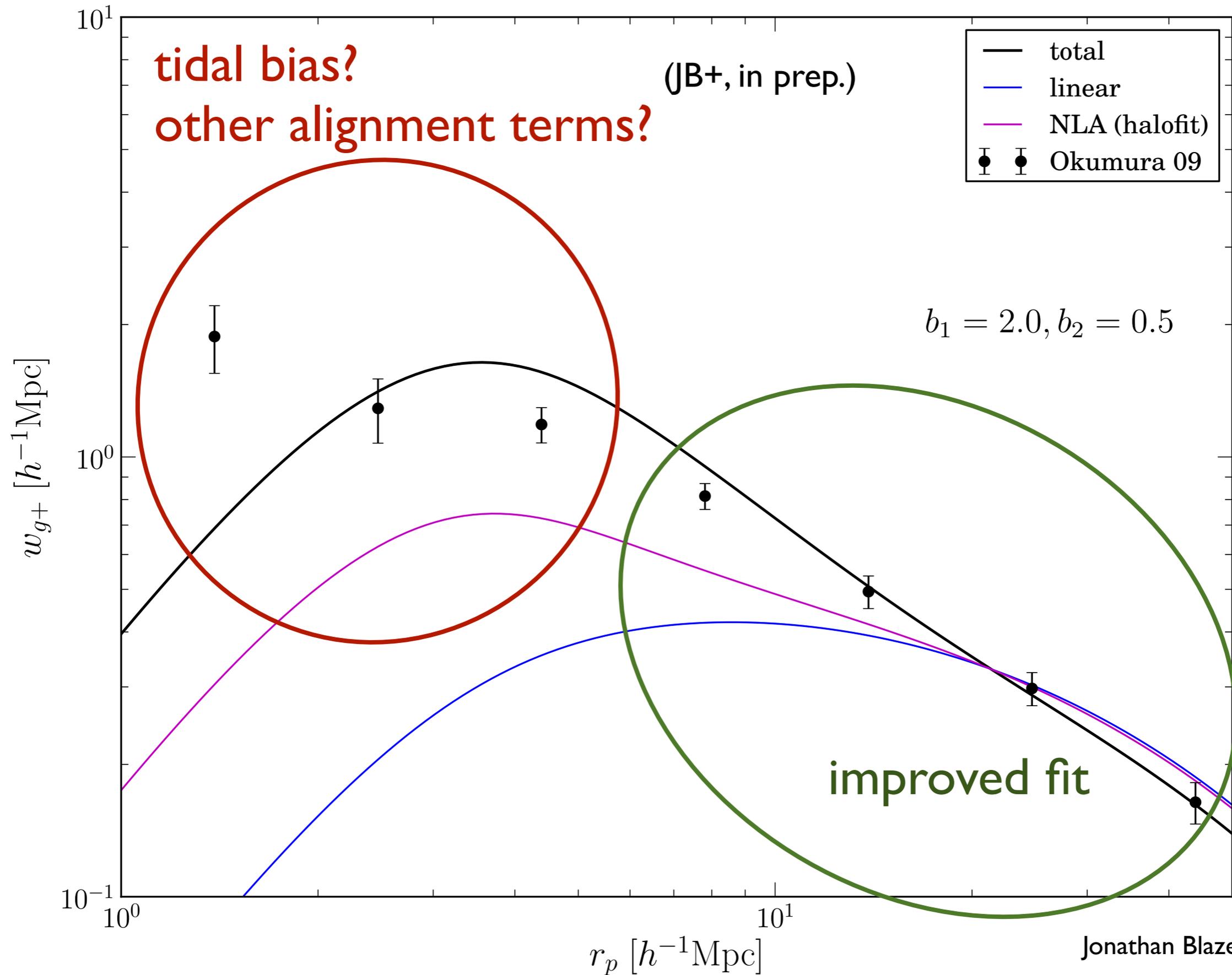
nonlinear bias

density weighting

Additional contributions

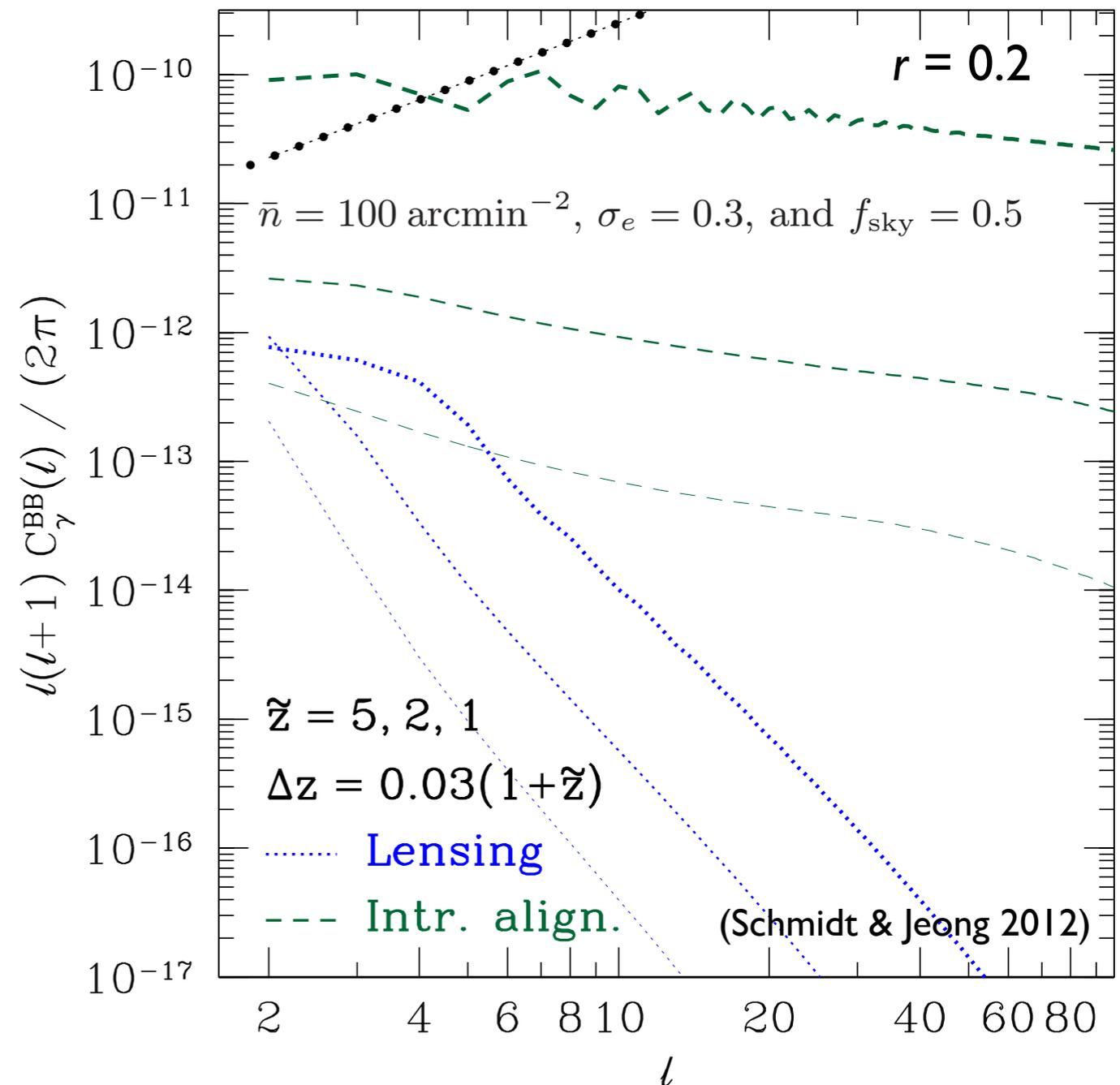


Fit to data



Intrinsic alignments as a signal

- halo and galaxy formation and evolution (e.g. JB+ 2011)
- additional tracer of LSS (e.g. Chisari & Dvorkin 2013)
- density field reconstruction (c.f. Lee & Pen 2000)
- nonlinear and tidal galaxy bias (JB+, in prep.)
- tidal field B-modes from primordial gravity waves (Schmidt & Jeong 2012; Chisari+ 2014)



Summary

- We have made a lot of progress in understanding IA - more is needed!
- Indirect measurement constraints IA contamination in SDSS to $\sim 2\%$ - apply method to new data sets
- Tidal alignment provides a good description for LRGs on large scales - nonlinear corrections improve the model on smaller scales
- Future/ongoing work:
 - Nonlinear alignment effects
 - Compare with N-body and hydro sims
 - IA measurements on new data sets, including DES
- See papers: JB+ 2011 (JCAP 5, 10); JB+ 2012 (JCAP 4, 41); JB+ in prep.